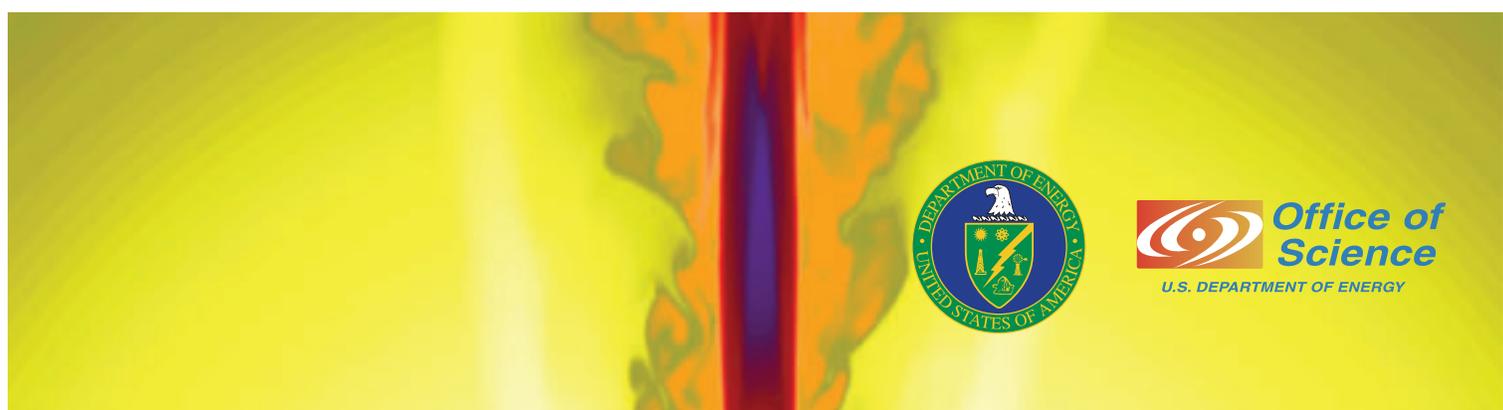




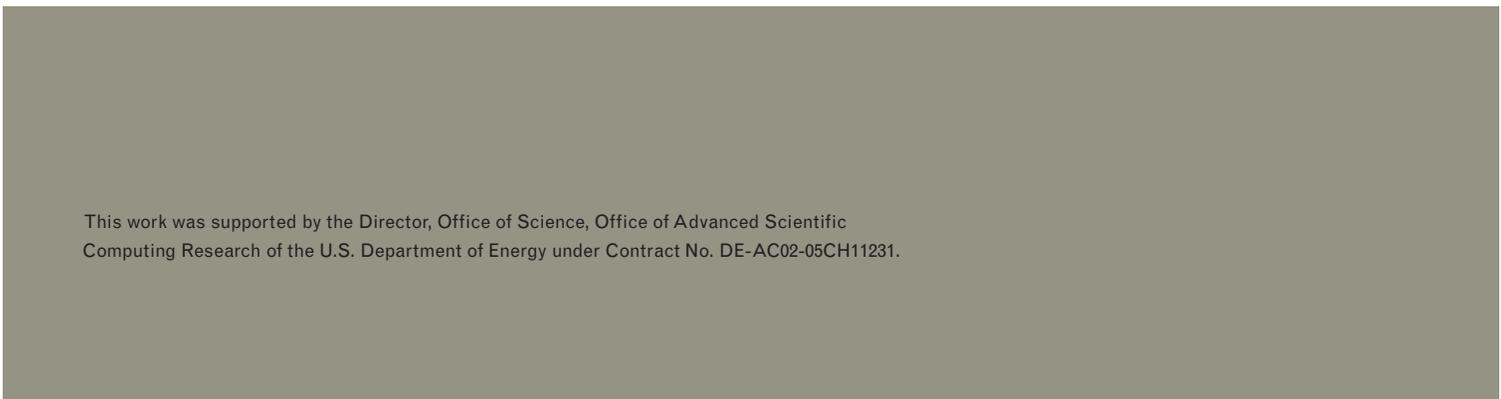
SCIENTIFIC DISCOVERY

A progress report on the U.S. Department of Energy's Scientific Discovery through Advanced Computing (SciDAC) Program





Cover: When the cores of massive stars collapse to form black holes and accretion disks, relativistic jets known as gamma ray bursts pass through and then explode the star. SciDAC's Supernova Science Center used simulations like the one of the cover to test and confirm theories about the origin of gamma-ray bursts.



This work was supported by the Director, Office of Science, Office of Advanced Scientific Computing Research of the U.S. Department of Energy under Contract No. DE-AC02-05CH11231.

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SciDAC

Scientific Discovery through
Advanced Computing

Executive Summary

Greetings:

Five years ago, the U.S. Department of Energy Office of Science launched an innovative software development program with a straightforward name – Scientific Discovery through Advanced Computing, or SciDAC. The goal was to develop scientific applications to effectively take advantage of the terascale supercomputers (capable of performing trillions of calculations per second) then becoming widely available.

From the most massive explosions in our universe to our planet's changing climate, from developing future energy sources to understanding the behavior of the tiniest particle, the SciDAC program has lived up to its name by helping scientists make important discoveries in many scientific areas.

Here are some examples of the achievements resulting from SciDAC:

- For the first time, astrophysicists created fully resolved simulations of the turbulent nuclear combustion in Type 1a supernovae, exploding stars which are critical to better understanding the nature of our universe.
- Using climate modeling tools developed and improved under SciDAC, climate change scientists in the United States are making the largest contribution of global climate modeling data to the world's leading body on climate studies, the Intergovernmental Panel on Climate Change.
- To better understand combustion, which provides 80 percent of the energy used in the United States,

scientists created the first laboratory-scale flame simulation in three dimensions, an achievement which will likely help improve efficiency and reduce pollution.

- Looking toward future energy resources, magnetic fusion researchers, applied mathematicians and computer scientists have worked together to successfully carry out advanced simulations on the most powerful modern supercomputing platforms. They discovered the favorable result that for the larger reactor-scale plasmas of the future such as ITER (the planned international fusion experiment), heat losses caused by plasma turbulence do not continue to follow the empirical trend of increasing with the size of the system.
- To make the most of existing particle accelerators and reduce the cost of building future accelerators, teams developed new methods for simulating improvements. In addition to helping us understand the most basic building blocks of matter, accelerators make possible nuclear medicine.
- Physicists studying the Standard Model of particle interactions have, after 30 years of trying, been able to model the full spectrum of particles known as hadrons at the highest level of accuracy ever. The results could help lead to a deeper understanding of the fundamental laws of physics.

These and other impressive scientific breakthroughs are the results of hundreds of researchers working in multidisciplinary teams at national

laboratories and universities across the country.

The SciDAC Program consisted of three research components:

- Creating a new generation of scientific simulation codes to take full advantage of the powerful computing capabilities of terascale supercomputers
- Creating the mathematical and computing systems software to enable these scientific simulation codes to effectively and efficiently use terascale computers
- Creating a distributed science software infrastructure to enable scientists to effectively collaborate in managing, disseminating and analyzing large datasets from large-scale computer simulations and scientific experiments and observations.

Not only did the researchers work in teams to complete their objectives, but the teams also collaborated with one another to help each other succeed. In addition to advancing the overall field of scientific computing by developing applications which can be used by other researchers, these teams of experts have set the stage for even more accomplishments.

This report will describe many of the successes of the first five years of SciDAC, but the full story is probably too extensive to capture – many of the projects have already spurred researchers in related fields to use the applications and methods developed under SciDAC to accelerate their own research, triggering even more breakthroughs.

While SciDAC may have started out as a specific program, Scientific Discovery through Advanced Computing has become a powerful concept for addressing some of the biggest challenges facing our nation and our world.

Michael Strayer

SciDAC Program Director
U.S. Department of Energy
Office of Science



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Scientific Discovery through Advanced Computing

Over the past 50 years, computers have transformed nearly every aspect of our lives, from entertainment to education, medicine to manufacturing, research to recreation. Many of these innovations are made possible by relentless efforts by scientists and engineers to improve the capabilities of high performance computers, then develop computational tools to take full advantage of those capabilities.

As these computer systems become larger and more powerful, they allow researchers to create more detailed models and simulations, enabling the design of more complex products, ranging from specialized medicines to fuel-efficient airplanes. Critical to the success of these efforts is an every-increasing understanding of the complex scientific processes and principles in areas such as physics, chemistry and biology.

Over the past five years, teams of scientists and engineers at national laboratories operated by the U.S. Department of Energy, along with researchers at universities around the country, have been part of a concerted program to accelerate both the performance of high performance computers and the computing allocations used to advance our scientific knowledge in areas of critical importance to the nation and the world.

Launched in 2001, this broad-based program is called Scientific Discovery through Advanced Computing, or SciDAC for short. The \$57 million-per-year program was designed to accelerate the devel-

opment of a new generation of tools and technologies for scientific computing. SciDAC projects were selected to capitalize on the proven success of multidisciplinary scientific teams. In all, 51 projects were announced involving collaborations among 13 DOE national laboratories and more than 50 colleges and universities.

SciDAC is an integrated program that has created a new generation of scientific simulation codes. The codes are being created to take full advantage of the extraordinary computing capabilities of today's terascale computers (computers capable of doing trillions of calculations per second) to address ever larger, more complex problems. The program also includes research on improved mathematical and computing systems software that will allow these codes to use modern parallel computers effectively and efficiently. Additionally, the program is developing "collaboratory" software to enable geographically separated scientists to effectively work together as a team, to control scientific instruments remotely and to share data more readily.

Today, almost five years after it was announced, the SciDAC program has lived up to its name and goal of advancing scientific discovery through advanced computing.

Teaming Up for Better Science

In the 1930s, physicist Ernest Orlando Lawrence pioneered a new approach to scientific research. Rather than working alone in their labs, physicists, engineers, chemists and other scientists were brought together to form multidisciplinary teams which could bring a range of knowledge and expertise to bear on solving scientific challenges. This "big science" approach formed the basis for the national laboratories operated by the Department of Energy.

Since the 1950s, scientists at the DOE national labs have increasingly employed high performance computers in their research. As a result, computational science has joined experimental science and theoretical science as one of the key methods of scientific discovery.

With SciDAC, this approach has come full circle, with multidisciplinary teams from various research institutions working together to develop new methods for scientific discovery, while also developing technologies to advance collaborative science.

And as with other research accomplishments, the results are being shared through articles published in scientific and technical journals. Additionally, the tools and techniques developed under the SciDAC program are freely available to other scientists working on similar challenges.

A Full Spectrum of Scientific Discovery

The SciDAC program was created to advance scientific research in all mission areas of the Department of Energy's Office of Science. From looking back into the origins of our

universe to predicting global climate change, from researching how to burn fuels more efficiently and cleanly to developing environmentally and economically sustainable energy sources, from investigating the behavior of the smallest particles to learning how to combine atoms to create new nanotechnologies, computational science advanced under SciDAC affects our lives on many levels.

Today, almost five years after it was announced, the SciDAC program has lived up to its name and goal of advancing scientific discovery through advanced computing. This progress report highlights a number of the scientific achievements made possible by SciDAC and also looks at the computing methods and infrastructure created under the program which are now driving computational science advances at research institutions around the world.

The achievements range from understanding massive exploding stars known as supernovae to studying the tiniest particles which combine to form all the matter in the universe. Other major accomplishments include better methods for studying global climate, developing new sources of energy, improving combustion to reduce pollution and designing future facilities for scientific research.

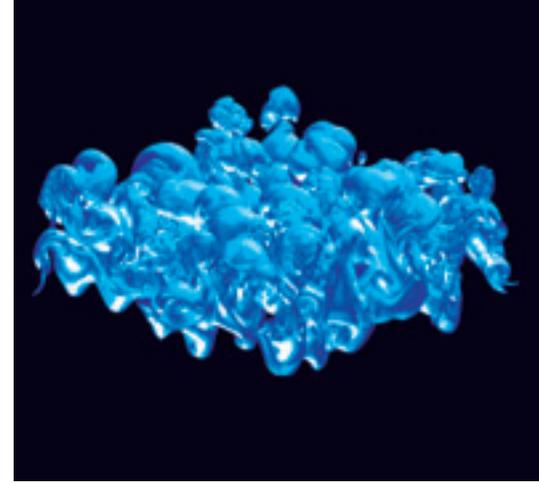
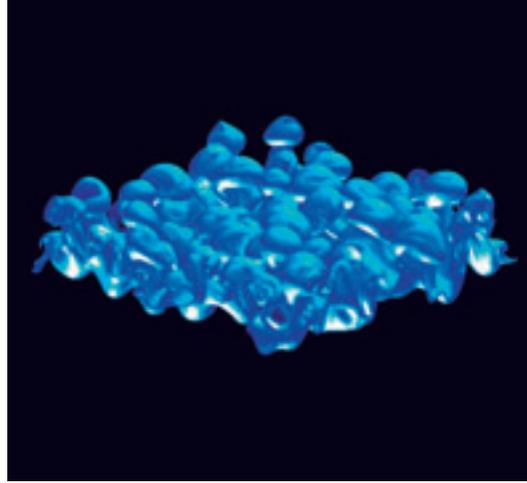
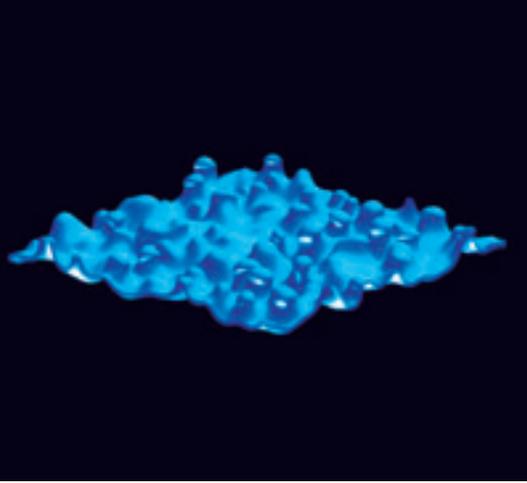
Computing then and now

While the achievements of the SciDAC teams are impressive on their own, they also demonstrate just how far scientific computing has come in the past 20 years.

Back in 1986, a state-of-the-art supercomputer was a Cray-2, which had a peak speed of 2 gigaflops (2 billion calculations) per second and a then-phenomenal 2 gigabytes of memory. Such machines were rare, located only at a limited number of research institutions, and access was strictly controlled. Scientific programs were typically written by individuals, who coded everything from scratch. Tuning the code to improve performance was done by hand – there were no tools freely available, no means of creating visualizations. The lucky scientists were able to submit jobs remotely using a 9600 baud modem.

Today, inexpensive highly parallel commodity clusters provide teraflops-level performance (trillions of calculations per second) and run on open source software. Extensive libraries of tools for high performance scientific computing are widely available – and expanding, thanks to SciDAC. Computing grids offer remote access at 10 gigabits per second, allowing hundreds or thousands of researchers to use a single supercomputer. Scientists can do code development on desktop computers which have processors faster than the Cray-2 and offer more memory. Once an application runs, powerful visualization tools provide detailed images which can be manipulated, studied and compared to experimental results.





Supernovae Science:

Earth-Based Combustion Simulation Tools Yield Insights into Supernovae

In the far reaches of space, a small star about the size of Earth simmers away, building up the heat and pressure necessary to trigger a massive thermonuclear explosion.

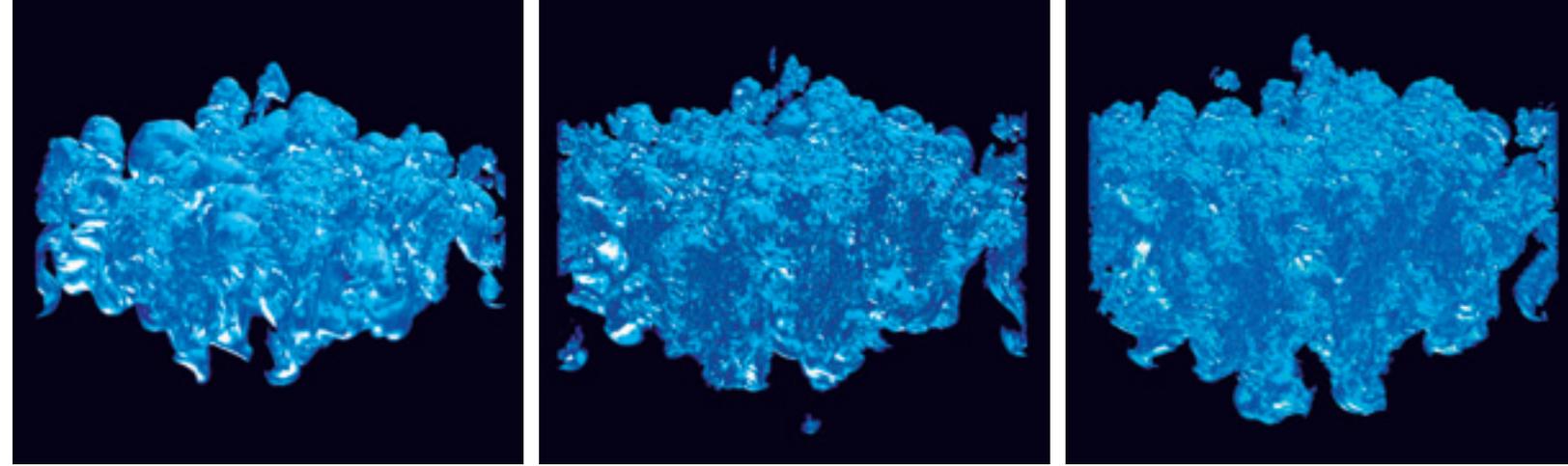


FIGURE 1. This series shows the development of a Rayleigh-Taylor unstable flame. The interface between the fuel and ash is visualized. The buoyant ash rises upward as the fuel falls downward in the strong gravitational field, wrinkling the flame front. This increases the surface area of the flame and accelerates the burning. At late times, the flame has become fully turbulent.

The star, known as a white dwarf, consists of oxygen and carbon. At its center, the star is about two billion times denser than water. Already 40 percent more massive than our Sun, the star continues to gain mass as a companion star dumps material onto the surface of the white dwarf.

As this weight is added, the interior of the star compresses, further heating the star. This causes the carbon nuclei to fuse together, creating heavier nuclei. The heating process, much like a pot of water simmering on the stove, continues for more than 100 years, with plumes of hot material moving through the star by convection and distributing the energy.

Eventually, the nuclear burning occurs at such a rate that convection cannot carry away all the energy that is generated and the heat builds dramatically. At this point, a burning thermonuclear flame front moves quickly out from the core of the star, burning all the fuel in a matter of seconds. The result is a Type Ia supernova, an exploding star which is one of the brightest objects in the universe. Because of their brightness, supernovae are of great interest to astrophysicists studying the origin, age and size of our universe.

Although astronomers have been observing and recording supernovae for more than 1,000 years, they still don't understand the exact mechanisms which cause a white dwarf to explode

as a supernova. Under SciDAC, teams of researchers worked to develop computational simulations to try to understand the processes. It turns out that an algorithm developed to simulate a flame from a Bunsen burner in a laboratory is well suited – with some adaptation – to studying the flame front moving through a white dwarf.

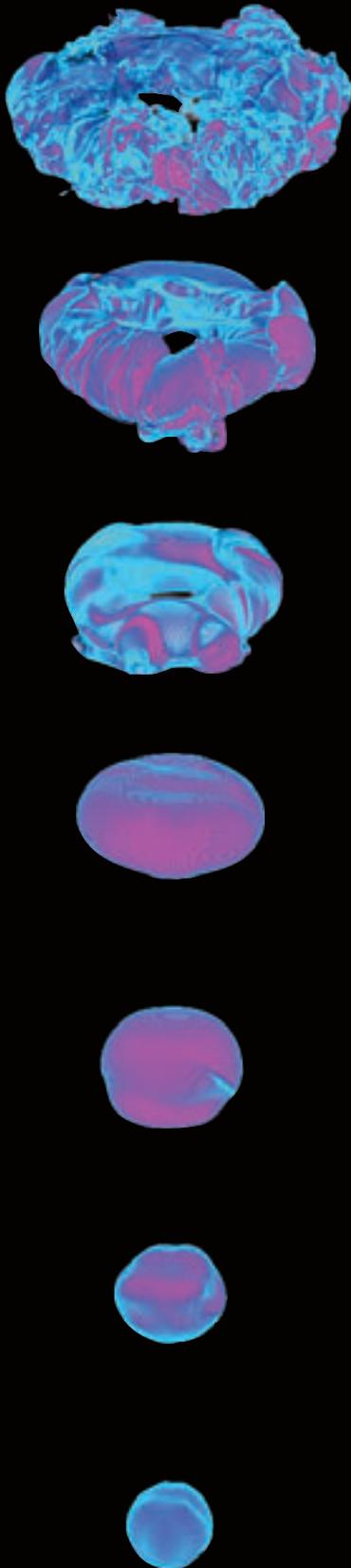
The SciDAC program brought together members of the Supernova Science Center project and LBNL mathematicians and computational scientists associated with the Applied Partial Differential Equations Integrated Software Infrastructure Center (ISIC). “We had two groups in very different fields,” said Michael Zingale, a professor at the State University of New York–Stony Brook and member of the Supernovae Science Center. “We met with these mathematicians who had developed codes to study combustion and through SciDAC, we started using their tools to study astrophysics.”

The adaptive mesh refinement (AMR) combustion codes developed at the Center for Computational Sciences and Engineering (CCSE) at Lawrence Berkeley National Laboratory are very effective for modeling slow-moving burning fronts. The behavior of the flame front leading up to a supernova is very similar, and the scientists worked together to adapt the codes to astrophysical environments.

Computational scientists tackle such problems by dividing them into small cells, each with different values for density, pressure, velocity and other conditions. The conditions are then evolved one time step at a time, using equations which say how things like mass, momentum and energy change over time, until a final time is reached. The astrophysicists had been using a code that included modeling the effects of sound waves on the flame front, which limited them to very small time steps. So, modeling a problem over a long period of time exceeded the availability of computing resources. The CCSE code allowed the astrophysicists to filter out the sound waves, which aren't important in this case, and take much bigger time steps, making more efficient use of their supercomputing allocations.

The net result, according to Zingale, is that the project team was able to carry out studies which had never been possible before. Their findings have been reported in a series of papers published in scientific journals such as *Astrophysical Journal*.

While observational astronomers are familiar with the massive explosion which characterizes a Type Ia supernova, one of the unknowns is where the flame starts, or if it starts in more than one location. When the flame front starts out, it is moving at about one one-thousandth of the



speed of sound. As it burns more fuel, the flame front speeds up. By the time the star explodes, the front must be moving at half the speed of sound. The trick is, how does it accelerate to that speed?

One answer explored by the SciDAC team is the effect of a wrinkled flame. As the flame becomes wrinkled, it has more surface area, which means it burns more fuel, which in turn accelerates the flame front. The AMR codes were used to model small areas of the flame front in very high resolution.

As the flame burns, it leaves in its wake “ash,” which through intense heat and pressure is fused into nickel. This hot ash – measuring several billion degrees Kelvin – is less dense than the fuel. The fuel ahead of the front is relatively cooler, at about 100 million degrees Kelvin, and denser than the ash. These conditions make the flame front unstable. This instability, known as the Rayleigh-Taylor instability, causes the flame to wrinkle.

As the flame starts out from the center and burns through most of the star, the front between the fuel and the ash is sharp and the front is called a “flamelet.”

Astrophysicists had predicted that as the flame front burns further out from the center, the lower density of the star would cause it to burn less vigorously and become more unstable due to increased turbulence and mixing of fuel and ash. This represents a different mode of combustion known as a “distributed burning regime.” Scientists believe that such combustion occurs in the late stage of a supernova explosion.

Using the AMR combustion codes and running simulations for 300,000 processor hours at DOE’s National Energy Research Scientific Computing Center, the Supernova Science Center team was able to create the first-ever three-dimensional simulations of such an event. The results are feeding into the astrophysics community’s knowledge base.

“While other astrophysicists are modeling entire stars at a different scale, they need to know what is going on at scales their models can’t resolve – and we’re providing that model for them,” Zingale said. “Although it takes a large amount of computing time, it is possible now thanks to these codes. Before our SciDAC partnership, it was impossible.”

FIGURE 2. This series of images show the growth of a buoyant reactive bubble. At the density simulated, the rise velocity is greater than the laminar flame velocity, and the bubble distorts significantly. The vortical motions transform it into a torus. Calculations of reactive rising bubbles can be used to gain more insight into the early stages of flame propagation in Type Ia supernovae.

From Soundwaves to Supernova:

SciDAC Provides Time, Resources to Help Astrophysicists Simulate a New Model

Once every 30 to 50 years – and then just for a few milliseconds – an exploding star known as a core-collapse supernova is the brightest object in the universe, brighter than the optical light of all other stars combined.

Supernovae have been documented for 1,000 years and astrophysicists know a lot about how they form, what happens during the explosion and what's left afterward. But for the past 40 years, one problem has dogged astrophysicists – what is the mechanism that actually triggers the massive explosion? Under SciDAC, a number of computational projects were established to simulate plausible explanations. One group created simulations showing a new mechanism for core-collapse supernovae. This model indicates that the core of the star generates sound waves which in turn become shock waves powerful enough to trigger the explosion.

Understanding these explosions and their signals is important, not

only because of their central role in astronomy and nucleosynthesis, but because a full understanding of supernovae may lead to a better understanding of basic physics.

The massive stars which become supernovae feature a number of conditions which are also found on Earth – winds, fluid flows, heating and cooling. But whereas a hurricane may blast along at 150 kilometers an hour on earth, the winds on these stars rage at tens of kilometers per second. The temperature of the star is 100 million times that of Earth, and the density of the object can be 14 orders of magnitude more dense than lead.

In short, studying supernovae is a very complicated problem. Key to finding an answer is understanding

the physics of all the interactions in the star.

One thing that is known is that supernovae produce neutrinos, particles with very little mass which travel through space and everything in their path. Neutrinos carry energy from the deep interior of the star, which is being shaken around like a jar of supersonic salad dressing, and deposit the energy on the outer region. One theory holds that if the neutrinos deposit enough energy throughout the star, this may trigger the explosion.

To study this, a group led by Adam Burrows, professor of astronomy at the University of Arizona and a member of the SciDAC Supernova Science Center, developed codes for simulating the behavior of a supernovae core in two dimensions. While a 3D version of the code would be optimum, it would take at least five more years to develop and would require up to 300 times as much computing time. As it was, the group ran 1.5 million hours of calculations at DOE's National Energy Research Scientific Computing Center.

But the two-dimensional model is suitable for Burrows' work, and the instabilities his group is interested in studying can be seen in 2D. What they found was that there is a big overturning motion in the core, which leads to wobbling, which in turn creates sound waves. These waves then carry energy away from the core, depositing it farther out near the mantle.

According to Burrows, these oscillations could provide the "power

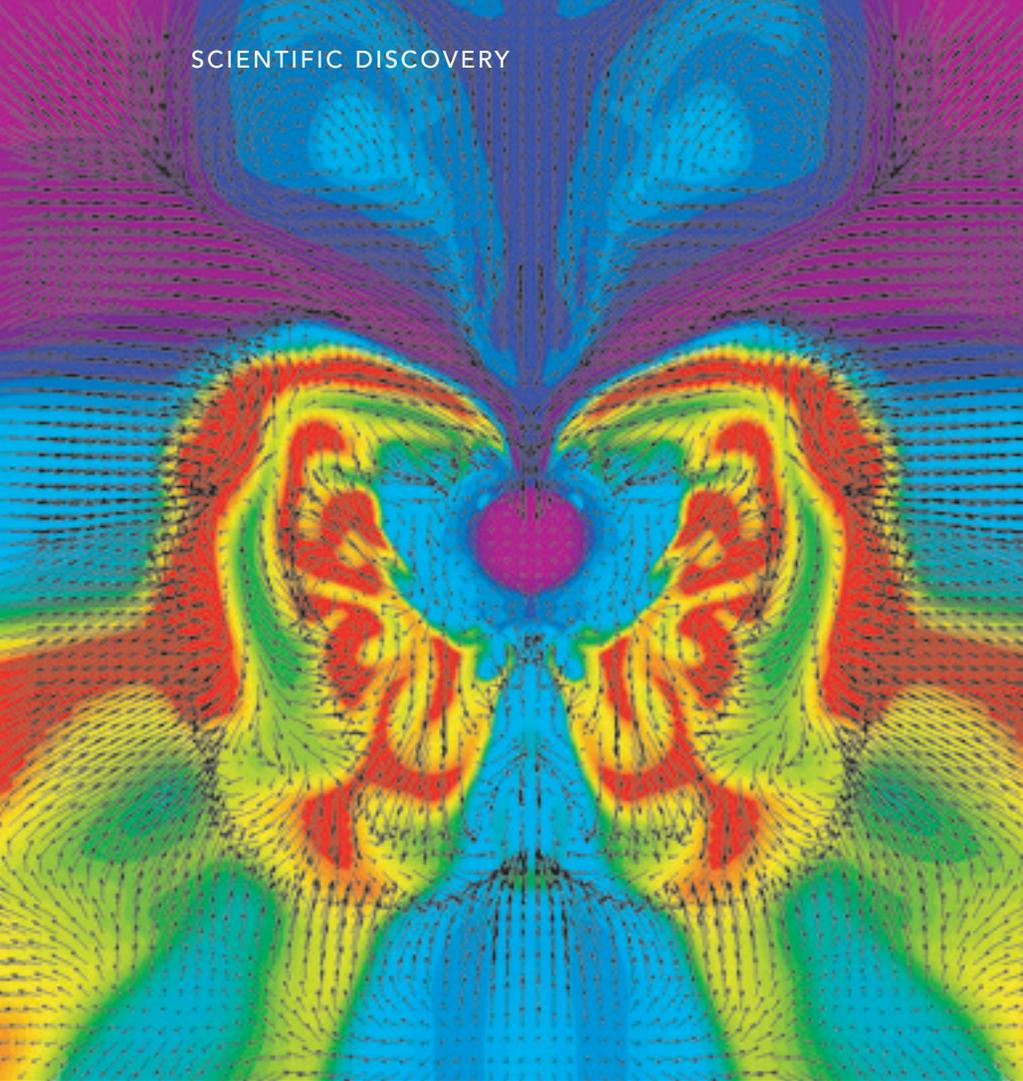
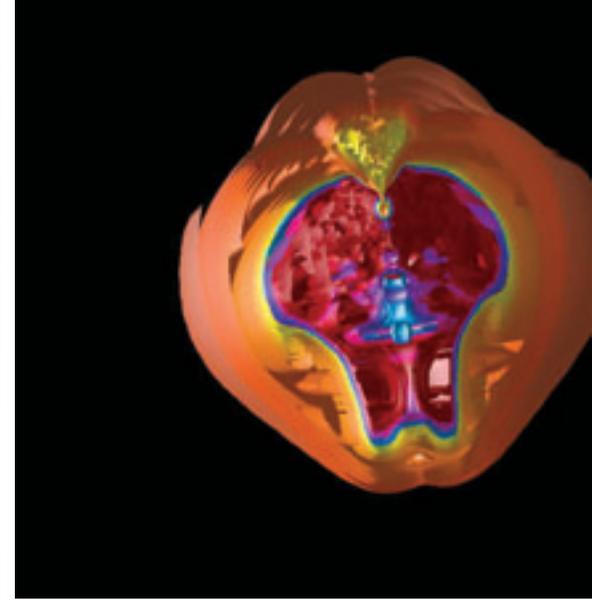


FIGURE 1. A 2D rendition of the entropy field of the early blast in the inner 500 km of an exploding supernova. Velocity vectors depict the direction and magnitude of the local flow. The bunching of the arrows indicates the crests of the sound waves that are escalating into shock waves. These waves are propagating outward, carrying energy from the core to the mantle and helping it to explode. The purple dot is the protoneutron star, and the purple streams crashing in on it are the accretion funnels. (All images courtesy of Adam Burrows, University of Arizona)



source” which puts the star over the edge and causes it to explode. To imagine what such a scenario would look like, think of a pond into which rocks are thrown, causing waves to ripple out. Now think of the pond as a sphere, with the waves moving throughout the sphere. As the waves move from the denser core to the less dense mantle, they speed up. According to the model, they begin to crack like a bullwhip, which creates shockwaves. It is these shockwaves, Burrows believes, which could trigger the explosion.

So, what led the team to this model? They came up with the idea by following the pulsar – the neutron star which is the remains of a supernova. They wanted to explore the origin of the high speed which pulsars seem to be born with and this led them to create a code that allowed the core to move. However, when they implemented this code, the core not only recoiled, but oscillated, and generated sound waves.

The possible explosive effect of the oscillations had not been considered before because previous simulations of the conditions inside the core used smaller time steps, which consumed more computing resources. With this limitation, the simulation ran their course before the onset of oscillations. With SciDAC support, however, Burrows’ team was able to develop new codes with larger time steps, allowing them to model the oscillations for the first time.

Calling the simulation a “real

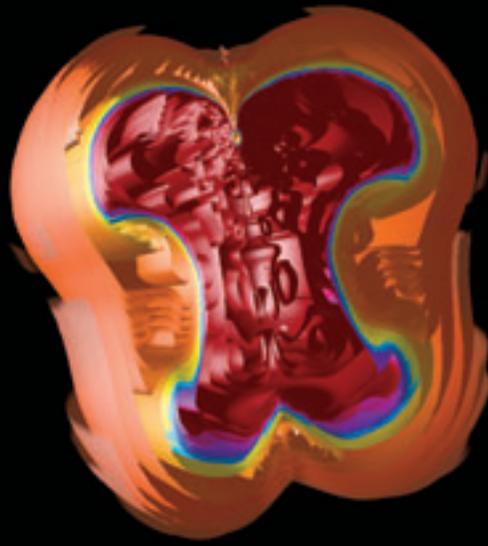


FIGURE 2. These shells are isodensity contours, colored according to entropy values. The red (blast area) indicates regions of high entropy, and the orange outer regions have low entropy. In the image on the left, matter is accreting from the top onto the protoneutron star in the center (the orange dot that looks like the uvula in the throat). The image is oriented so that the anisotropic explosion is emerging down and towards the viewer. The scale is roughly 5000 km. The image on the right shows the same explosion later in time and at a different orientation with respect to the viewer.

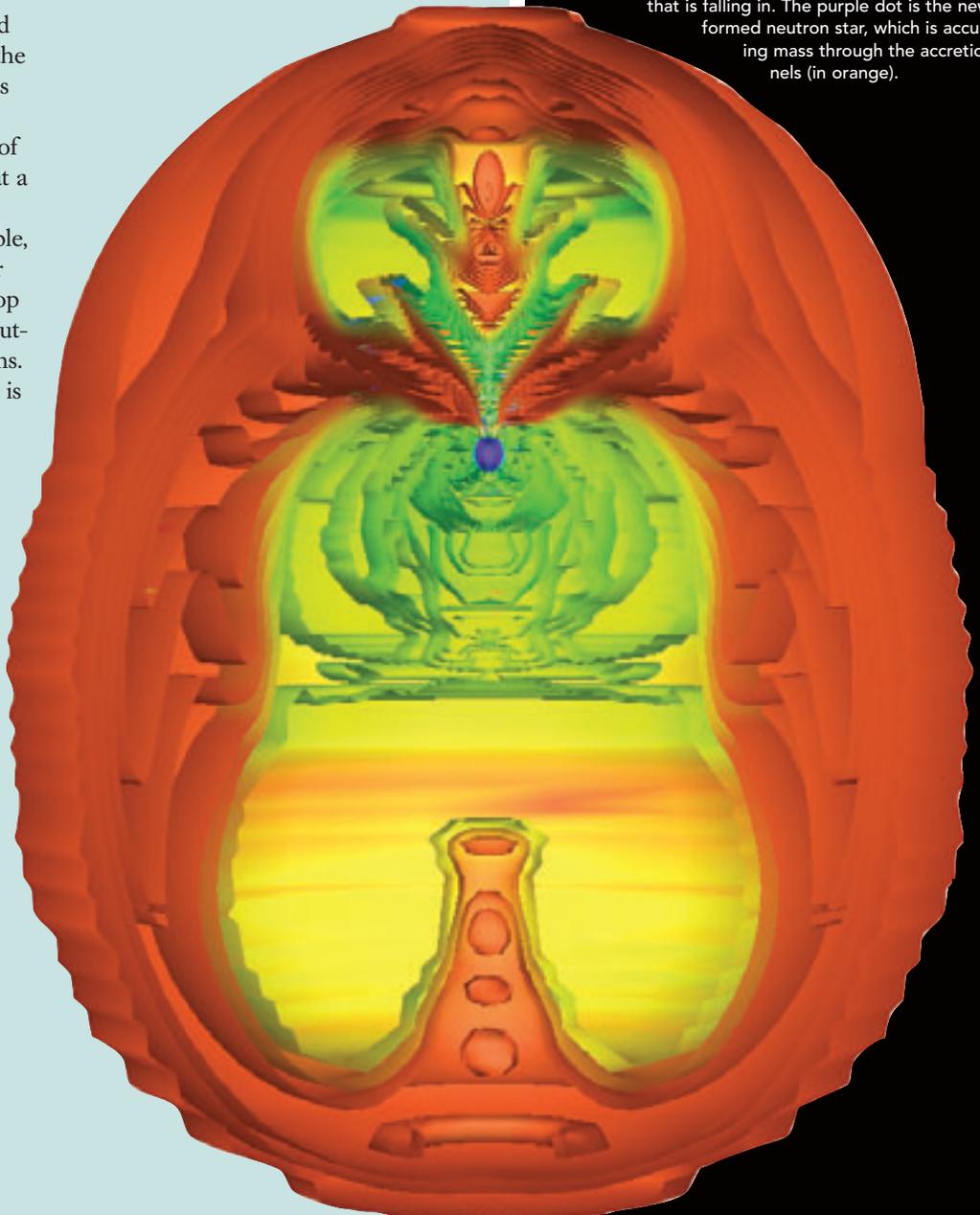
numerical challenge,” Burrows said the resulting approach “liberated the inner core to allow it to execute its natural multidimensional motion.” This motion led to the excitation of the core, causing the oscillations at a distinct frequency.

SciDAC made the work possible, he said, by providing support over five years – enough time to develop and test the code – and the computing resources to run the simulations. The results look promising, but as is often the case, more research is needed before a definitive mechanism for triggering a supernova is determined.

The group published a paper on their research in the *Astrophysical Journal*. Neutrino transfer is included as a central theme in a 2D multi-group, multi-neutrino, flux-limited transport scheme. It is approximate but has the important components, and is the only truly 2D neutrino code with results published in the archival literature.

“The problem isn’t solved,” Burrows said. “In fact, it’s just beginning.”

FIGURE 3. Another isodensity shell colored with entropy, showing simultaneous accretion on the top and explosion on the bottom. The inner green region is the blast, and the outer orange region is the unshocked material that is falling in. The purple dot is the newly formed neutron star, which is accumulating mass through the accretion funnels (in orange).



Insight into Supernovae:

SciDAC's Terascale Supernova Initiative
Discovers New Models for Stellar Explosions,
Spinning Pulsars

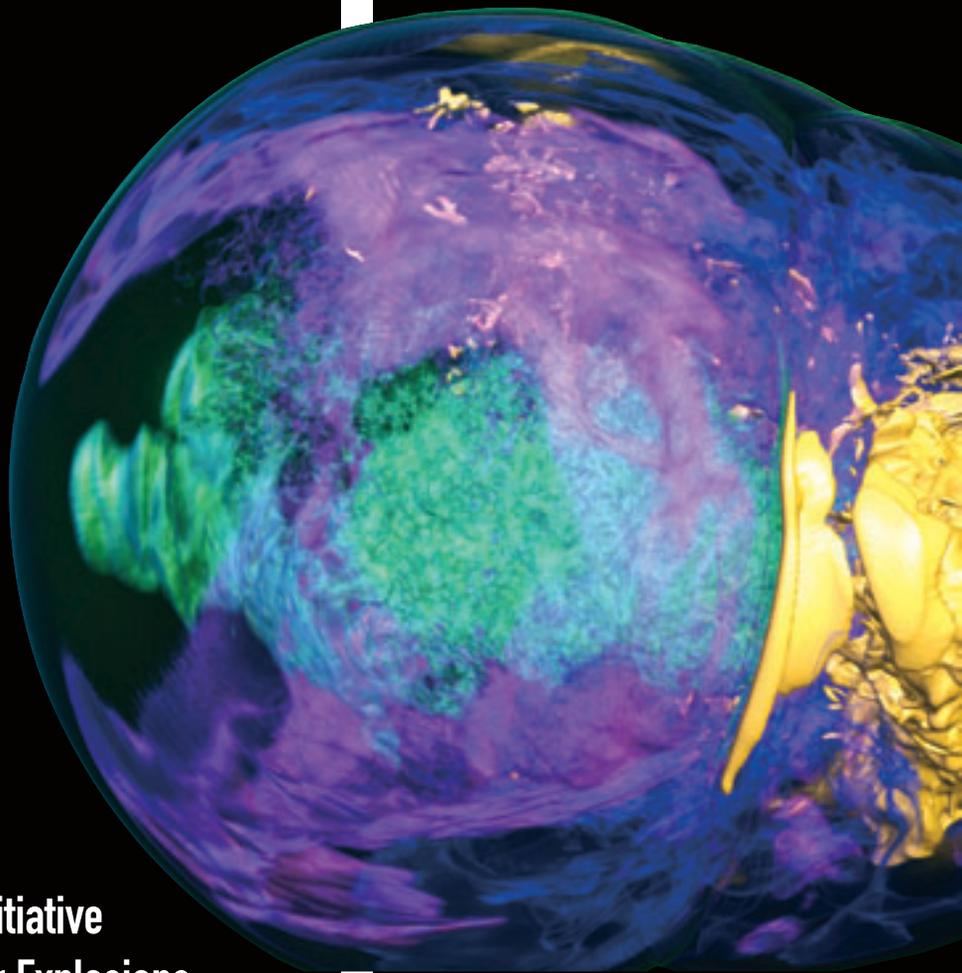
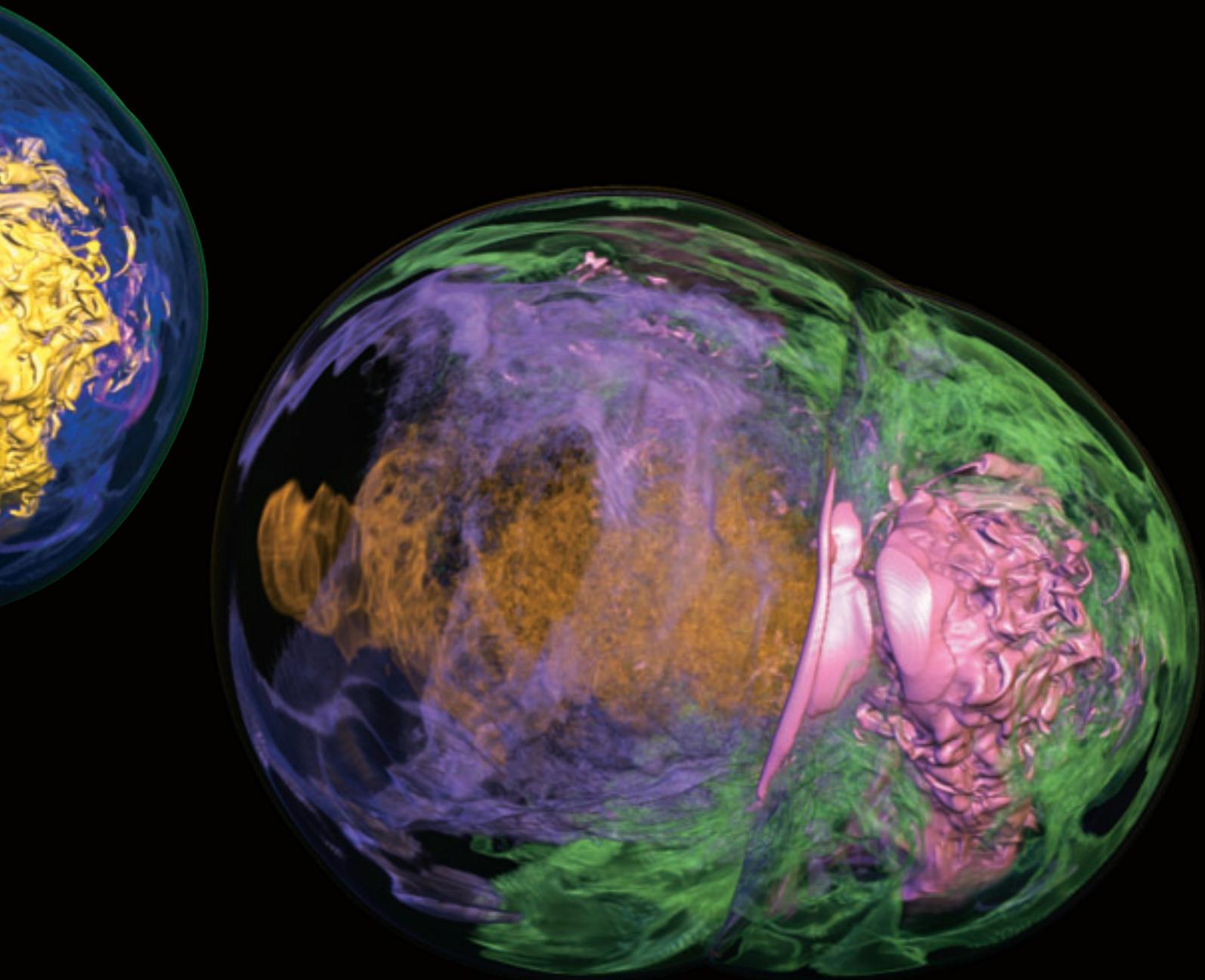


FIGURE 1: This image provides a snapshot of the angular momentum of the layered fluid flow in the stellar core below the supernova shock wave (the outer surface) during a core collapse supernova explosion. Pink depicts a significant flow rotating in one direction directly below the shock wave. Gold depicts a deeper flow directly above the proto-neutron star surface, moving in the opposite direction. The supernova shock wave instability (SASI) leads to such counter rotating flows and is important in powering the supernova, and may be responsible for the spin of pulsars (rotating neutron stars). The inner flow (in green) spins up the proto-neutron star. These three-dimensional simulations were performed by John Blondin (NCSU) under the auspices of the Terascale Supernova Initiative, led by Tony Mezzacappa (ORNL). The visualization was performed by Kwan-Liu Ma (University of California, Davis).

The massive stellar explosions known as core-collapse supernovae are not only some of the brightest and most powerful phenomena in the universe, but are also the source of many of the elements which make up our universe – from the iron in our blood cells to the planet we live on to the solar systems visible in the night sky.

Supernova science has advanced dramatically with the advent of powerful telescopes and other instruments for observing and measuring the deaths of these stars, which are 10 times more massive than our sun, or more. At the same time, increasingly accurate applications for simulating supernovae have been developed to



take advantage of the growing power of bigger and faster supercomputers.

Scientists now know that these massive stars are layered like onions. Around the iron core are layer after layer of lighter elements. As the stellar core becomes more massive, gravity causes the core to collapse to a certain point at which it rebounds like a

compressed ball. This results in a shock wave which propagates from the core to the outermost layers, causing the star to explode.

Despite these gains in understanding, however, a key question remains unanswered: What triggers the forces which lead a star to explode? Under SciDAC, the Terascale Supernova

Initiative (TSI) was launched, consisting of astrophysicists, nuclear physicists, computer scientists, mathematicians and networking engineers at national laboratories and universities around the country. Their singular focus was to understand how these massive stars die.

“When these stars die in stellar

explosions known as core-collapse supernovae, they produce many of the elements in the universe. In fact, they are arguably the single most important source of elements,” said Tony Mezzacappa, an astrophysicist at Oak Ridge National Laboratory and principal investigator for the TSI. “Learning what triggers supernovae is tantamount to understanding how we came to be in the universe.”

Ascertaining the explosion mechanism is one of the most important questions in physics. It is a complex, multi-physics problem involving fluid flow, instability, and turbulence, stellar rotation, nuclear physics, particle physics, radiation transport and magnetic fields.

“Much of modern physics comes to play a role in a core-collapse supernova,” Mezzacappa says. “The cosmos is very much a physics laboratory.”

Once the explosion mechanism is accurately predicted, scientists will be able to predict all the other associated byproducts and phenomena, such as the synthesis of elements. The key to solving this question depends on developing more detailed computer models, such as those advanced under SciDAC, and detailed observations of core-collapse supernovae, which only occur about twice every century in our galaxy (although many such supernovae are observed each year from outside of our galaxy). As models are developed and improved, their accuracy can be tested by comparing the results with the observed data. By adjusting any parameters of the model (of course, the goal in developing sophisticated models is to minimize the number of free parameters), scientists can generate results which are closer and closer to the actual data.

Because the simulations are computationally intensive, Mezzacappa and John Blondin, a professor at North Carolina State University, started by developing codes to look at the fluid flow at the core of a supernova. To focus on this one area, they removed the other physics components from their code, knowing they would have

to add them again later. They were interested in learning how the core flow behaves during the explosion as one goes from two spatial dimensions to three. They first ran the code in two dimensions, then in 3D. When they ran the simulations, they started to see instability in the shock wave. But, Mezzacappa said, they initially believed that the physics in their code should not have led to that behavior. Their first reaction was that the code contained an error, but further study led them into a new area of supernova research.

Despite decades of supernova theory focused on the formation and propagation of the shock wave, apparently no one considered whether the shock wave was stable, according to Mezzacappa. The TSI team’s research led them to conclude that the shock wave is unstable; and if it is perturbed, it grows in an unbounded, nonlinear fashion. As it spreads, it becomes more and more distorted. The TSI group found that the instability of the shock wave could help contribute to the explosion mechanism. For example, a distorted shock wave could explain why supernovae explode asymmetrically.

Under further study, the TSI researchers found that as they added more physics back into their code, the instability did not go away, giving them more confidence in their findings. Calling this new discovery the Stationary Accretion Shock Instability, or SASI, the team published their findings, which were then corroborated by other supernova researchers. As a result, Mezzacappa said, the work has fundamentally changed scientists’ thinking about the explosion phenomenon. And, he adds, it probably would not have happened without the multidisciplinary approach fostered by SciDAC.

For starters, SciDAC provided the team with access to some of the world’s fastest supercomputers – a resource otherwise unavailable to researchers like Blondin. This access to the Cray X1E and Cray XT3 sys-

tems at Oak Ridge gave Blondin the computing horsepower he needed to run his simulations in 3D.

But scaling up from two to three dimensions is not just a matter of running on a larger computer. It is also a matter of developing more detailed codes.

In this case, neutrinos are central to the dynamics of core-collapse supernovae. The explosion releases 10^{53} ergs of radiation, almost all of it in the form of neutrinos. The explosion energy (the kinetic energy of the ejected material) is only 10^{51} ergs. This intense emission of radiation plays a key role in powering the supernova, heating the interior material below the shock wave and adding energy to drive the shock wave outward. But this radiation transport – the production, movement and interaction of the neutrinos – is one of the most difficult things to simulate computationally. It requires that the supercomputer solve a huge number of algebraic equations. Applied mathematicians on the TSI team developed methods for solving the equations on terascale supercomputers.

“This is another example of how SciDAC makes a world of difference – we could not have done these simulations without the help of the applied mathematicians,” Mezzacappa said. “SciDAC really enabled us to think about this problem in all of its complexity in earnest for the first time.”

But this also led to a new challenge. Suddenly, the team was faced with managing massive amounts of data with no infrastructure for analyzing or visualizing these terabytes of data. “We didn’t know what we had,” Mezzacappa said.

The TSI team partnered with another SciDAC project led by Micah Beck of the University of Tennessee, which developed a new data transfer solution known as logistical networking. Logistical networking software tools allow users to create local storage “depots” or utilize shared storage depots deployed worldwide to easily accomplish long-haul data transfers,

temporary storage of large datasets (on the order of terabytes) and pre-positioning of data for fast on-demand delivery.

The group provided the hardware and software needed to create a private network linking Oak Ridge and North Carolina State, allowing the data to be moved to the university. There, Blondin used a visualization cluster to analyze and visualize the data, allowing the team to pursue the 3D science. The TSI network now spans the U.S., linking the two original sites with facilities like the National Energy Research Scientific Computing Center in California to the State University of New York at Stony Brook. As the TSI project demonstrated the benefits of logistical networking, researchers in the fusion and combustion communities also built their own logistical networks.

Once the data from the 3D simulations could be analyzed, the team made another discovery that was not

possible in the two-dimensional simulations. Once a core-collapse supernova explodes, what remains is known as a neutron star. In some cases, this neutron star spins and emits radiation as light and is called a pulsar.

What is not known, however, is what causes a neutron star to get spun up.

A previous theory held that neutron star spin is generated when the stellar core spins up as it collapses during the supernova, much like an ice skater who increases his rotational speed by pulling his arms close to his body. But this simple model cannot explain all the observed characteristics of pulsars and at the same time explain the predicted characteristics of stars at the onset of collapse.

In TSI's 3D models, the SASI produces two counter-rotating flows between the proto-neutron star and the propagating shock wave. As the innermost flow settles onto the proto-neutron star, it imparts angular momentum and causes it to spin, just

as you can spin a bicycle tire by swiping your hand across the tread. The team computed their spin predictions and found that the results were within the observed range of pulsars spins.

"This was another breakthrough made possible by SciDAC," Mezzacappa said. "It gives us new possibilities for explaining the spins of pulsars."

The key to their results in both cases, Mezzacappa noted, was scaling their application to simulate complex phenomena in 3D, rather than two dimensions. "In other scientific articles, supernova researchers have written that they don't see much difference between simulations in 2D and 3D, but now we know they are very different, and have shown how this difference is related to the supernova mechanism and byproducts," Mezzacappa said. "Mother Nature clearly works in a 3D universe."

CCSM:

Advanced Simulation Models for Studying Global Climate Change

One of the most widely discussed scientific questions of the past 20 years is the issue of global climate change. But it's not just a matter of science. The question of global climate change also involves economic, environmental, social and political implications. In short, it's a complex question with no easy answers. But detailed climate modeling tools are helping researchers gain a better understanding of how human activities are influencing our climate.

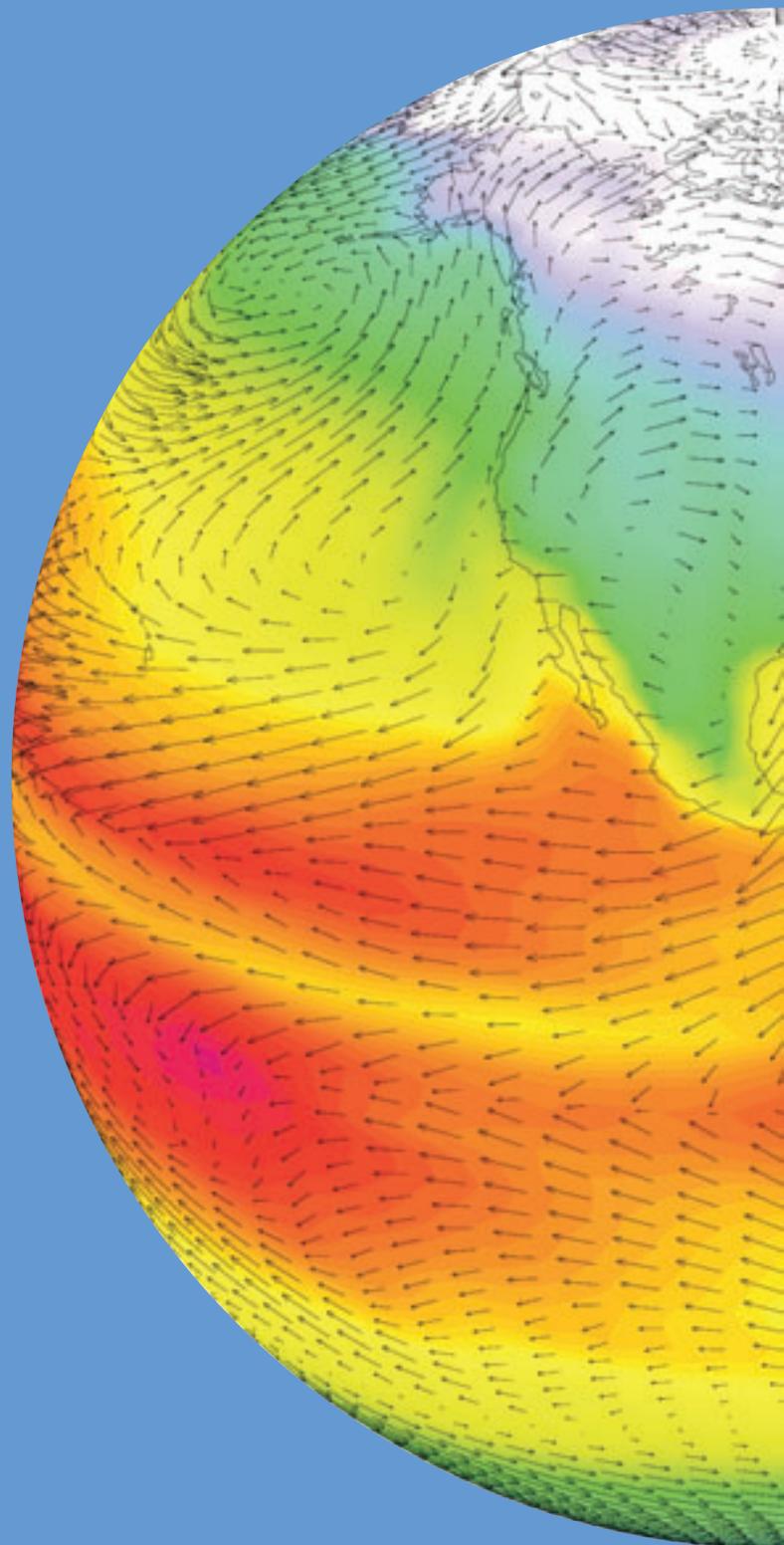
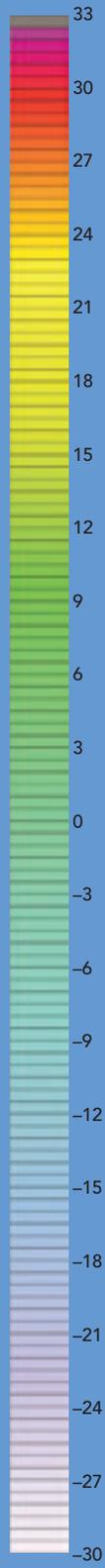
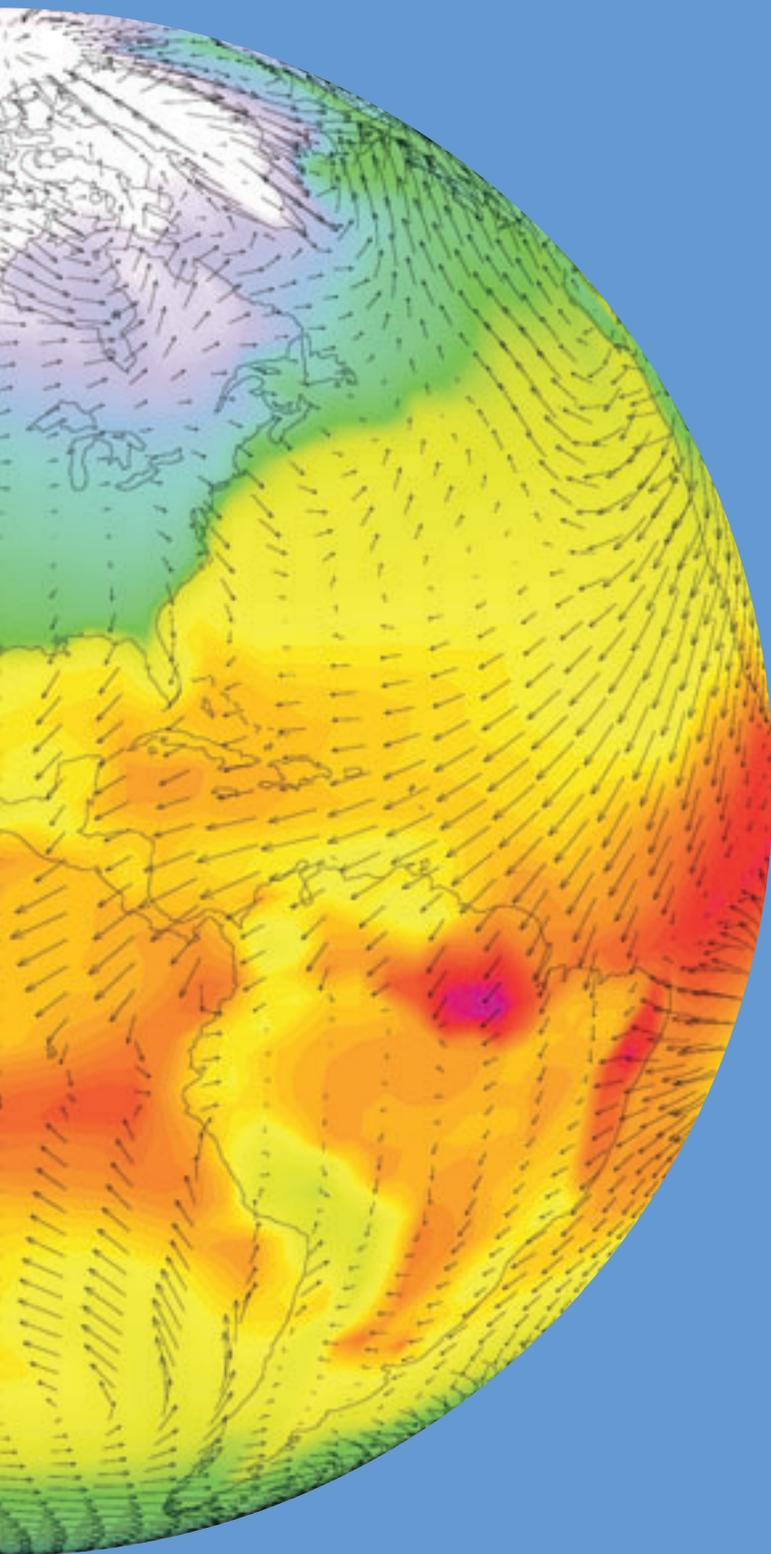


FIGURE 1. Surface temperature (in °C) averaged over December of year 9 of the fully coupled chemistry simulation. Vectors represent the surface wind.



To try to ensure that decision-makers around the world have access to the most accurate information available, the Intergovernmental Panel on Climate Change (IPCC) was established in 1988 under the auspices of the World Meteorological Organization and the United Nations Environment Program. The role of the IPCC is to assess on a comprehensive, objective, open and transparent basis the scientific, technical and socio-economic information relevant to understanding the scientific basis of risk of human-induced climate change, its potential impacts and options for adaptation and mitigation.

In accordance with its mandate, the major activity of the IPCC is to prepare at regular intervals comprehensive and up-to-date assessments of relevant for the understanding of human induced climate change, potential impacts of climate change and options for mitigation and adaptation. The First Assessment Report was completed in 1990, the Second Assessment Report in 1995 and the Third Assessment Report in 2001. The Fourth Assessment Report is scheduled to be completed in 2007. SciDAC-sponsored research has enabled the United States climate modeling community to make significant contributions to this report. In fact, the United States is the largest contributor of climate modeling data to the report, as compared to the 2001 report in which no U.S.-generated data was included.

Two large multi-laboratory SciDAC projects are directly relevant to the activities of the IPCC. The first, entitled "Collaborative Design and Development of the Community Climate System Model for Terascale Computers," has made important software contributions to the recently released third version of the Community Climate System Model (CCSM3.0), developed by the National Center for Atmospheric Research, DOE laboratories and the academic community. The second project, entitled "Earth System Grid

II: Turning Climate Datasets into Community Resources,” aims to facilitate the distribution of the copious amounts of data produced by coupled climate model integrations to the general scientific community.

A key chapter in the Fourth Assessment Report will look at Global

ran extensive simulations using the CCSM and another climate modeling application, PCM, to project climate change through the end of the 21st century.

The results, reported in the March 18, 2005 issue of *Science* magazine, indicate that “even if the concentrations of greenhouse gases in the atmosphere had been stabilized in the year 2000, we are already committed to further global warming of about another half degree and an additional 320 percent sea level rise caused by thermal expansion by the end of the 21st century. ... At any given point in time, even if concentrations are stabilized, there is a commitment to future climate changes that will be greater than those we have already observed.”

As one of many research organizations from around the world providing input for the IPCC report, the DOE Office of Science is committed to providing data that is as scientifically accurate as possible. This commitment to scientific validity is a key factor behind the SciDAC program to refine and improve climate models.

In order to bring the best science to bear on future climate prediction, modelers must first try to accurately simulate the past. The record of historical climate change is well documented by weather observations and measurements starting around 1890. So, after what is known about the physics and mathematics of atmospheric and ocean flows is incorporated in a simulation model running on one of DOE’s supercomputers, the model is tested by “predicting” the climate from 1890 to the present.

The input to such a model is not the answer, but rather a set of data describing known climate conditions – what the atmospheric concentrations of greenhouse gases were, what solar fluctuations occurred, and what volcanic eruptions took place. While past climate models were not able to replicate the historical record with its natural and induced variability, the present generation of coupled ocean, atmosphere, sea ice, and land compo-

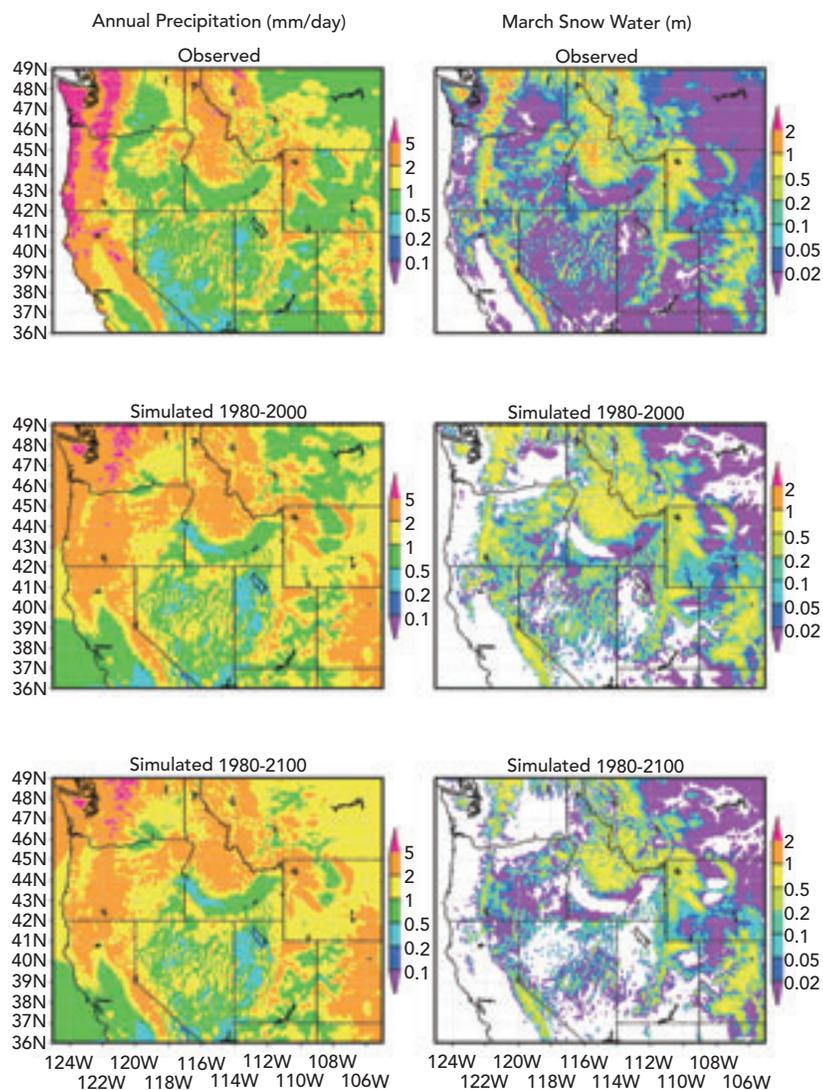


FIGURE 2. Annual precipitation and March snow water from an IPCC simulation using the subgrid orography scheme in Ghan and Shippert (2005).

Climate Projections, which will examine climate change to the year 2100 and beyond. Gerald Meehl, one of two lead authors of the chapter and a scientist at the National Center for Atmospheric Research in Colorado,

nents do a remarkably good job of predicting the past century's climate. This accuracy gives confidence that when the same model is used to generate predictions into the future based on various emission scenarios, the results will have a strong scientific basis and are more than scientific opinions of what might happen.

The Community Climate System Model, CCSM3, is centered at NCAR, arguably a world leader in the field of coupled climate models. This model continues to be developed jointly by the National Science Foundation and DOE to capture the best scientific understanding of climate. In addition to being used for climate change studies, the CCSM is being used to study climate-related questions ranging from ancient climates to the physics of clouds. However, the DOE national laboratories and the National Center for Atmospheric Research periodically use the model for assessment purposes such as the IPCC project.

The SciDAC CCSM Consortium developing the model has also made sure that it runs effectively on the most powerful supercomputers at DOE's leading computing centers – the Leadership Class Facility (LCF) in Tennessee and the National Energy Research Scientific Computing Center (NERSC) in California. DOE contributed significant allocations of supercomputer time to the international efforts (along with NCAR and the Japanese Earth Simulator Center) toward the completion of the IPCC runs with CCSM3.

The fruit of these labors is substantial. Enabled by these sizable allocations, the CCSM3 is the largest single contributor to the IPCC Assessment Report 4 database. Not only is the model represented in more transient scenarios than any other model, more statistically independent realizations of each of these scenarios have been integrated than with any other model. Also, the resolution of the atmosphere component is exceeded by only one other model

(which has not been ensemble-integrated).

This increase in production is scientifically important in many ways. The need for large numbers of individual simulations is especially important when attempting to characterize the uncertainty of climate change. For instance, in looking at recent climate change (over the last century), running more statistically independent simulations of the 20th century allows researchers to produce a much better defined pattern of climate changes. To model future climate change, scientists will run multiple simulations with a given level of greenhouse gas emissions to quantify the range of possible outcomes. And when researchers want to study possible outcomes for a range of emissions, they need to run even more simulations.

Running simulations for extended times are also important to help researchers detect and take into account conditions which result from flaws in the climate model. Such flaws cause "drift," or an unsubstantiated change in temperature. While drift-free control runs covering several hundred years are useful for studying climate change over a few decades, examining climate change of the entire 20th century requires a control run of at least 1,000 years to avoid misinterpretation of the resulting climate model. In summary, large ensembles and long control runs are necessary to better understand the clear patterns of climate change.

Through SciDAC, DOE also sponsors the technology being used to make available the results to scientists worldwide. The Earth System Grid (ESG) project collects all the data from all the runs of models in the U.S., as well as from other countries, and makes the data accessible to scientists throughout the world. This allows scientists in the international climate research community to analyze more extensive datasets, compare results and document differences in model predictions. In fact,

over 200 papers are in press referencing this data. (You can see a list at http://www-pcmdi.llnl.gov/ipcc/diagnostic_subprojects.php.)

When researchers encounter disagreement among models, scientific debates and discussions ensue about the physical mechanisms governing the dynamics of climate. This debate is typically vigorous and points in directions that models can be improved. To take advantage of such discussions and to improve the validity of climate models, DOE sponsors the Program for Model Intercomparison and Diagnosis (PCMDI), which places the scientific basis of climate modeling front and center in its work. Such rigorous peer review and discussion is critical to ensuring that government officials have access to the most scientifically valid information when making decisions regarding climate change.

This SciDAC project is a collaboration between six DOE National Laboratories (ANL, LANL, LBNL, LLNL, ORNL, PNNL) and NCAR. The lead investigators representing these institutions are R. Jacob, P. Jones, C. Ding, P. Cameron-Smith, J. Drake, S. Ghan, W. Collins, W. Washington and P. Gent.

Plasma Physics

Terascale Simulations Predict Favorable Confinement Trend for Reactor-Scale Plasmas

Most people do not think about turbulence very often, except when they are flying and the captain turns on the “Fasten Seat Belts” sign. The kind of turbulence that may cause problems for airplane passengers involves swirls and eddies that are a great deal larger than the aircraft. But in fusion plasmas, much smaller-scale turbulence, called microturbulence, can cause serious problems—specifically, instabilities and heat loss that could stop the fusion reaction.



In fusion research, all of the conditions necessary to keep a plasma dense and hot long enough to undergo fusion are referred to as confinement. The retention of heat, called energy confinement, can be threatened by microturbulence, which can make particles drift across, rather than along with, the plasma flow. At the core of a fusion reactor such as a tokamak, the temperatures and densities are higher than at the outside edges. As with weather, when there

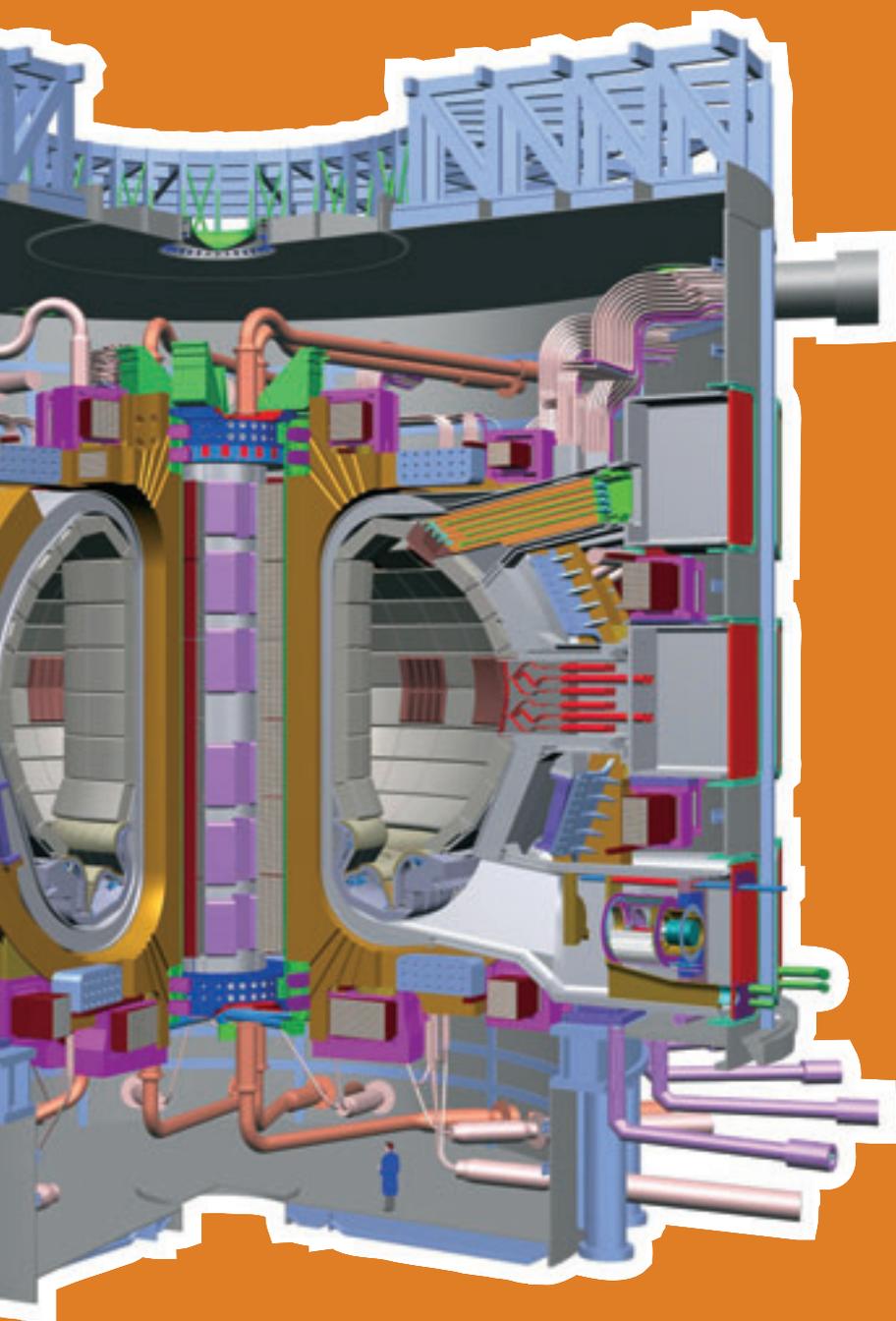


FIGURE 1. Cutaway illustration of the ITER tokamak.

be built at Cadarache in southern France, is one of the highest strategic priorities of the DOE Office of Science.

Underlining America's commitment to ITER, U.S. Energy Secretary Samuel Bodman stated on June 28, 2005, "Plentiful, reliable energy is critical to worldwide economic development. Fusion technologies have the potential to transform how energy is produced and provide significant amounts of safe, environmentally friendly power in the future. The ITER project will make this vision a reality."

ITER is expected to produce 500 million thermal watts of fusion power—10 times more power than is needed to heat the plasma—when it reaches full operation around the year 2016. As the world's first production-scale fusion reactor (Figure 1), ITER will help answer questions about the most efficient ways to configure and operate future commercial reactors.

The growth of the microinstabilities that lead to turbulent transport has been extensively studied over the years, not only because understanding this process is an important practical problem, but also because it is a true scientific grand challenge which is particularly well suited to be addressed by modern terascale computational resources. And the latest news, confirmed by multiple simulations using different codes, is good: in reactors the size of ITER, heat losses caused by plasma turbulence no longer follow the empirical trend of increasing with the size of the plasma. Instead, the rate of heat loss levels off and stabilizes.

Progress in understanding plasma ion dynamics has been impressive. For example, studies show that electrostatic ion temperature gradient (ITG) driven turbulence can be sup-

are two regions with different temperatures and densities, the area between is subject to turbulence. In a tokamak, turbulence can allow charged particles in the plasma to move toward the outer edges of the reactor rather than fusing with other particles in the core. If enough particles drift away, the plasma loses temperature and the fusion reaction cannot be sustained.

One troublesome result of tokamak experiments to date is that as

the size of the plasma increases, the relative level of heat loss from turbulent transport also increases. Because the size (and therefore cost) of a fusion experiment is determined largely by the balance between fusion self-heating and turbulent transport losses, understanding this process is of utmost importance for the design and operation of fusion devices such as the multi-billion-dollar ITER project. ITER (Latin for "the way"), a multinational tokamak experiment to

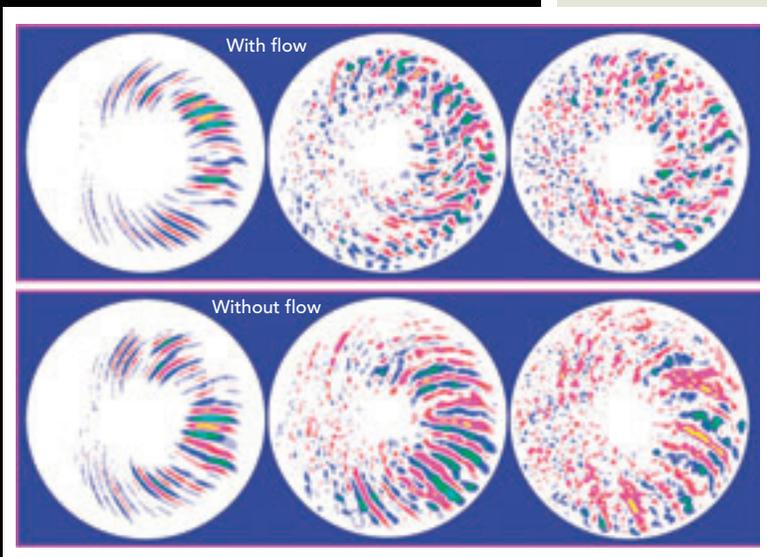


FIGURE 2. Turbulence reduction via sheared plasma flow compared to case with flow suppressed.

pressed by self-generated zonal flows within the plasma. The suppression of turbulence is caused by a shearing action that destroys the finger-like density contours which promote thermal transport. This dynamic process is depicted by the sequences shown in Figure 2, obtained using the Gyrokinetic

Toroidal Code (GTC) developed by the SciDAC-funded Gyrokinetic Particle Simulation Center (GPSC). The lower panels show the nonlinear evolution of the turbulence in the absence of flow, while the upper panels illustrate the turbulence decorrelation caused by the self-generated “E×B” flow, which arises from crossed electric and magnetic fields.

For the reactor-scale plasmas of the future, these simulations suggest that the relative level of heat loss driven by electrostatic ITG turbulence does not increase with plasma size (Figure 3). This transition from “Bohm” (linear) scaling to “gyro-Bohm” (quadratic) scaling is a positive trend, because simple empirical extrapolation of the smaller system findings would produce much more pessimistic predictions for energy confinement. Since neither experiments nor theory and simulations have previously been able to explore such trends in an ITER-sized plasma, these results represent a significant scientific discovery enabled by the SciDAC program.

Exploration of the underlying causes for the transition in the rate of heat loss that simulations show around the 400 gyroradii range has inspired the development of new nonlinear theoretical models based on the spreading of turbulence. Although this predicted trend is a very favorable one, the fidelity of the analysis needs to be further examined by investigating additional physics effects, such as kinetic electromagnetic dynamics, which might alter the present predictions.

The excellent scaling of the GTC code provides strong encouragement that future simulations will be able to capture the additional complexity and lead to greater scientific discoveries. GTC is currently involved in benchmark tests on a variety of supercomputing platforms—including the Earth Simulator in Japan, the Cray X1E and XT line at Oak Ridge National Laboratory, the IBM Power SP line, and the IBM Blue Gene line—and exhibits scaling properties which look to be readily extensible to the petascale regime. GTC’s performance and scaling on a variety of leading platforms is illustrated in Figure 4.

The GS2 and GYRO codes developed by the Plasma Microturbulence Project (the SciDAC predecessor to the current GPSC) have also contributed productively to the interpretation of turbulence-driven transport

This simulation is a good example of the effective use of powerful supercomputers (in this case, the 10 teraflop/s

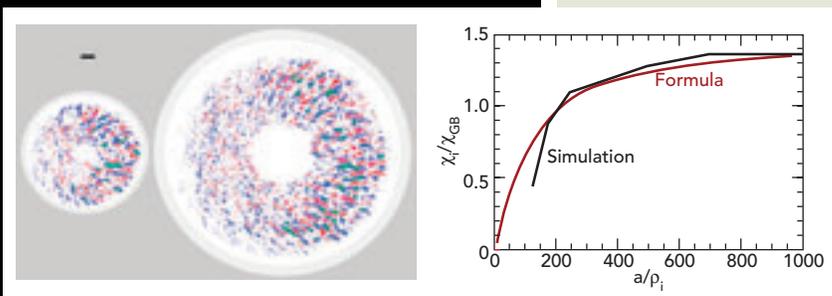


FIGURE 3. Full torus particle-in-cell gyrokinetic simulations (GTC) of turbulent transport scaling. (Left) The granular structures represent the scales of the turbulence in a typical plasma which need to be included in realistic plasma simulations. (Right) The horizontal axis expresses the plasma size, and the point at 1000 represents ITER’s size. The vertical axis represents the thermal diffusion, or heat loss.

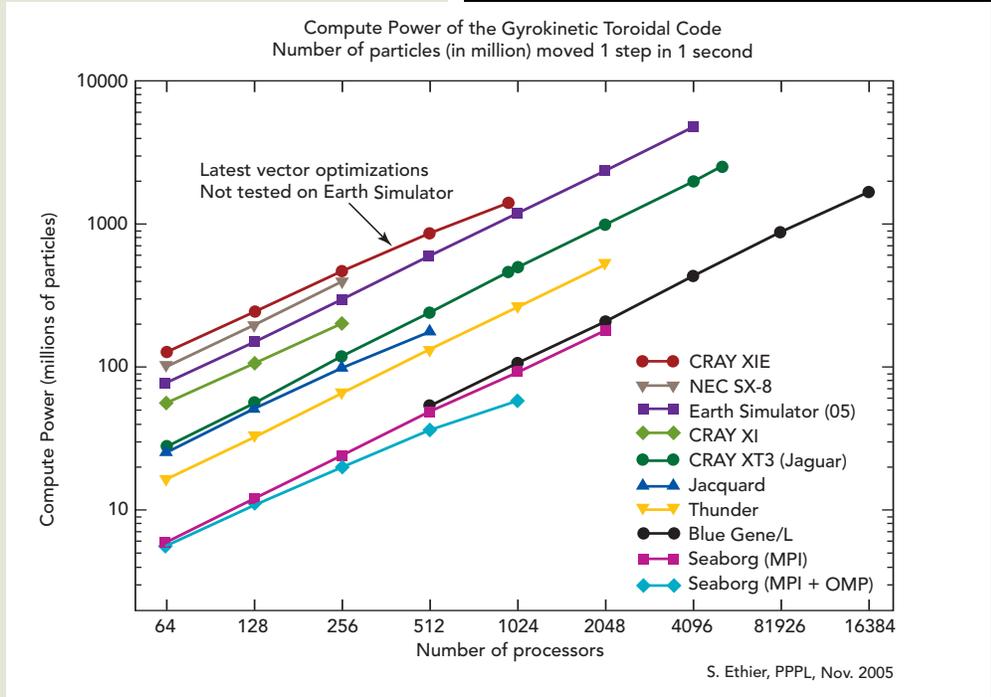
Seaborg IBM SP at NERSC). Typical global particle-in-cell simulations of this type have used one billion particles with 125 million grid points over 7000 time steps to produce significant physics results. Simulations of this size would not be feasible on smaller computers.

Large-scale simulations have also explored key consequences of scaling up from present-day experimental fusion devices (around 3 meters radius for the largest existing machines) to ITER-sized reactors (about 6 meters).

trends observed in experiments. These two SciDAC projects have involved researchers from Lawrence Livermore National Laboratory; Princeton Plasma Physics Laboratory (PPPL); Columbia University; the University of California at Los Angeles, Irvine, and Davis; the University of Colorado; the University of Maryland; the University of Tennessee at Knoxville; Oak Ridge National Laboratory; and General Atomics.

Bill Tang, Chief Scientist at PPPL, is encouraged by the progress in computational fusion research. "SciDAC has contributed strongly to the accelerated development of computational tools and techniques needed to develop predictive models for the analysis and design of magnetically confined plasmas," Tang commented. "Unraveling the complex behavior of strongly nonlinear plasma systems under realistic conditions is a key component of the next frontier of computational plasma physics in general and fusion research in particular. Accelerated progress in the development of the needed codes with higher physics fidelity has been greatly aided by the interdisciplinary alliances championed by the SciDAC Program, together with necessary access to the tremendous increase in compute cycles enabled by the rapid advances in supercomputer technology."

Ray Orbach, Director of the DOE Office of Science, expressed his hope for high-end computing's future contributions to fusion energy in an interview in SciDAC Review (Number 1,



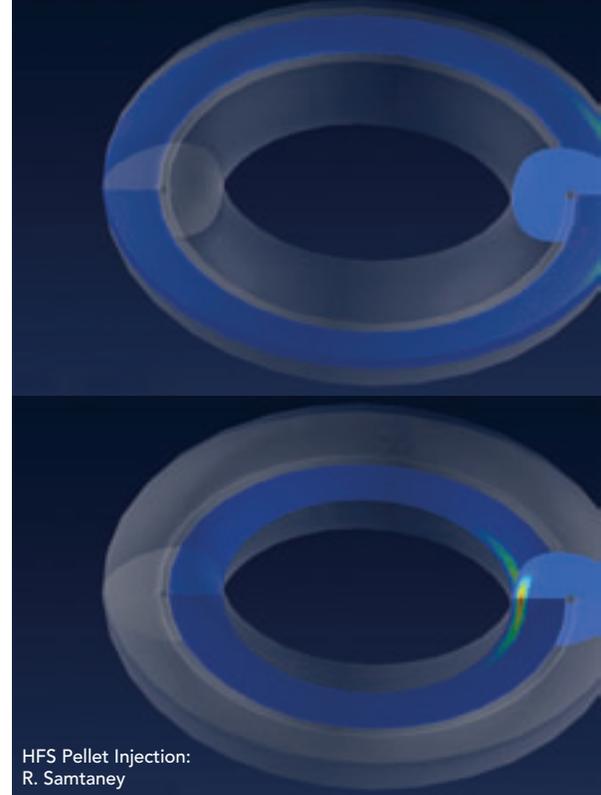
Spring 2006, p. 8): "At the speeds that we are talking about [50 teraflop/s], the electrons in a fusion device can be considered as real point particles and do not have to be treated in mean field approximations. So for the first time at 50 TF, one will be able to do simulations of high-density, high-temperature plasmas that were never possible before. This will have a significant impact on the treatment of these highly nonlinear systems. These systems are not subject to analytic examination and there may be instabilities that no one has thought about. This has already been found to be of profound importance for ITER [and] for fusion science in general."

FIGURE 4. Scaling study of the GTC code on multiple high performance computing platforms.

Fueling the Future

Simulations Examine the Behavior of Frozen Fuel Pellets in Fusion Reactors

What happens when you shoot one of the coldest materials into the hottest environment on earth? The answer may help solve the world's energy crisis.



HFS Pellet Injection:
R. Samtaney

Imagine that there is a large chamber in hell that's shaped like a doughnut, and that you'd like to shoot a series of hailstones into that chamber so that as they melt, the water vapor penetrates as deeply as possible and disperses as evenly as possible throughout the chamber. Setting aside for a moment the question of *why* you would want to do that, consider the multitude of *how* questions: Should you shoot in the hailstones from outside the ring of the doughnut or from inside the hole? What size hailstones should you use? At what speed and angle should they enter the chamber?

This strange scenario is actually an analogy for one of the questions facing fusion energy researchers: how to refuel a tokamak. A tokamak is a machine that produces a toroidal (doughnut-shaped) magnetic field. In that field, two isotopes of hydrogen – deuterium and tritium – are heated to about 100 million degrees Celsius (more than six times hotter than the interior of the sun), stripping the electrons from the nuclei. The magnetic field makes the electrically charged particles follow spiral paths around the magnetic field lines, so that they spin around the torus in a

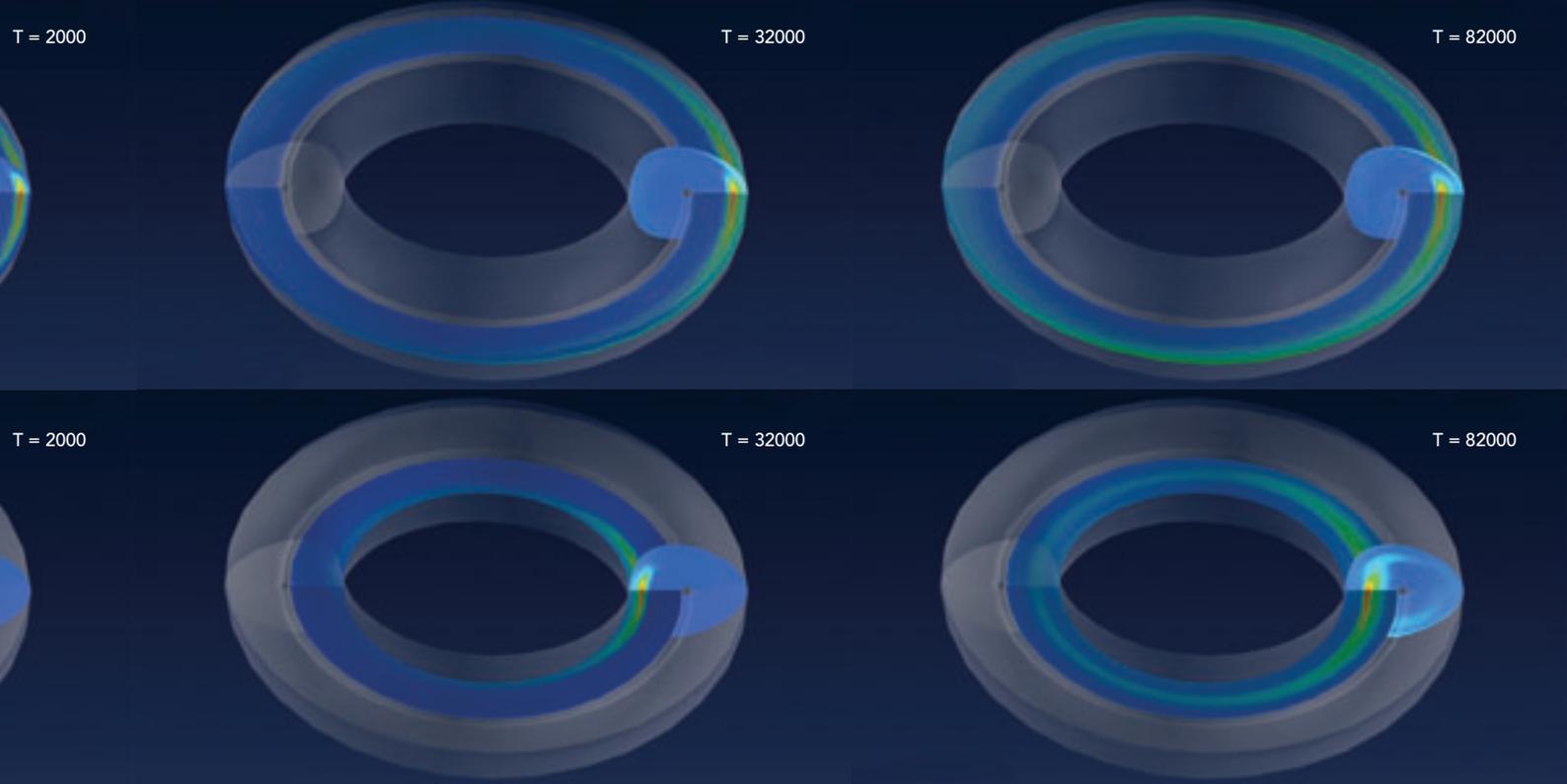


FIGURE 1. These simulations show the results of pellet injection into a tokamak fusion reactor from the low-field-side (LFS) and high-field-side (HFS). The top row shows a time sequence of the density in LFS injection while the bottom panel shows density evolution in HFS injection. The dominant motion of the ablated pellet mass is along field lines accompanied by transport of material across flux surfaces towards the low field side. This observation is qualitatively consistent with experimental observations leading to the conclusion that HFS pellet injection is a more efficient refueling technique than LFS injection. The MHD instabilities which cause the pellet material to move towards the low field side are currently under investigation.

fairly uniform flow and interact with each other, not with the walls of the tokamak. When the hydrogen ions (nuclei) collide at high speeds, they fuse, releasing energy. If the fusion reaction can be sustained long enough that the amount of energy released exceeds the amount needed to heat the plasma, researchers will have reached their goal: a viable energy source from abundant fuel that produces no greenhouse gases and no long-lived radioactive byproducts.

High-speed injection of frozen hydrogen pellets is an experimentally proven method of refueling a tokamak. These pellets are about the size of small hailstones (3–6 mm) and have a temperature of about 10 degrees Celsius above absolute zero. The goal is to have these pellets penetrate as deeply as possible into the plasma so that the fuel disperses evenly.

Pellet injection will be the primary fueling method used in ITER (Latin for “the way”), a multinational

tokamak experiment to be built at Cadarache in southern France. ITER, one of the highest strategic priorities of the DOE Office of Science, is expected to produce 500 million thermal watts of fusion power – 10 times more power than is needed to heat the plasma – when it reaches full operation around the year 2016. As the world’s first production-scale fusion reactor, ITER will help answer questions about the most efficient ways to configure and operate future commercial reactors.

However, designing a pellet injection system that can effectively deliver fuel to the interior of ITER represents a special challenge because of its unprecedented large size and high temperatures. Experiments have shown that a pellet’s penetration distance into the plasma depends strongly on how the injector is oriented in relation to the torus. For example, an “inside launch” (from inside the torus ring) results in better fuel distribution than an “outside launch” (from outside the ring).

In the past, progress in developing an efficient refueling strategy for ITER has required lengthy and expensive experiments. But thanks to a three-year, SciDAC-funded collaboration between the Computational Plasma Physics Theory Group at Princeton Plasma Physics Laboratory and the Advanced Numerical Algorithms Group at Lawrence Berkeley National Laboratory, computer codes have now reproduced some key experimental findings, resulting in significant progress toward the scientific goal of using simulations to predict the results of pellet injection in tokamaks.

“To understand refueling by pellet injection, we need to understand two phases of the physical process,” said Ravi Samtaney, the Princeton researcher who is leading the code development effort. “The first phase is the transition from a frozen pellet to gaseous hydrogen, and the second phase is the distribution of that gas in the existing plasma.”

The first phase is fairly well

understood from experiments and theoretical studies. In this phase, called ablation, the outer layer of the frozen hydrogen pellet is quickly heated, transforming it from a solid into an expanding cloud of dense hydrogen gas surrounding the pellet. This gas quickly heats up, is ionized, and merges into the plasma. As ablation continues, the pellet shrinks until all of it has been gasified and ionized.

The second phase – the distribution of the hydrogen gas in the plasma – is less well understood. Ideally, the injected fuel would simply follow the magnetic field lines and the “flux surfaces” that they define, maintaining a stable and uniform plasma pressure. But experiments have shown that the high-density region around the pellet

quickly heats up to form a local region of high pressure, higher than can be stably confined by the local magnetic field. A form of “local instability” (like a mini-tornado) then develops, causing the high-density region to rapidly move across, rather than along, the field lines and flux surfaces – a motion referred to as “anomalous” because it deviates from the large-scale motion of the plasma.

Fortunately, researchers have discovered that they can use this instability to their advantage by injecting the pellet from inside the torus ring, because from this starting point, the anomalous motion brings the fuel pellet closer to the center of the plasma, where it does the most good. This anomalous motion is one of the phe-

nomena that Samtaney and his colleagues want to quantify and examine in detail.

Figure 3 shows the fuel distribution phase as simulated by Samtaney and his colleagues in the first detailed 3D calculations of pellet injection. The inside launch (top row) distributes the fuel in the central region of the plasma, as desired, while the outside launch (bottom row) disperses the fuel near the plasma boundary, as shown in experiments.

Simulating pellet injection in 3D is difficult because the physical processes span several decades of time and space scales. The large disparity between pellet size and tokamak size, the large density differences between the pellet ablation cloud and

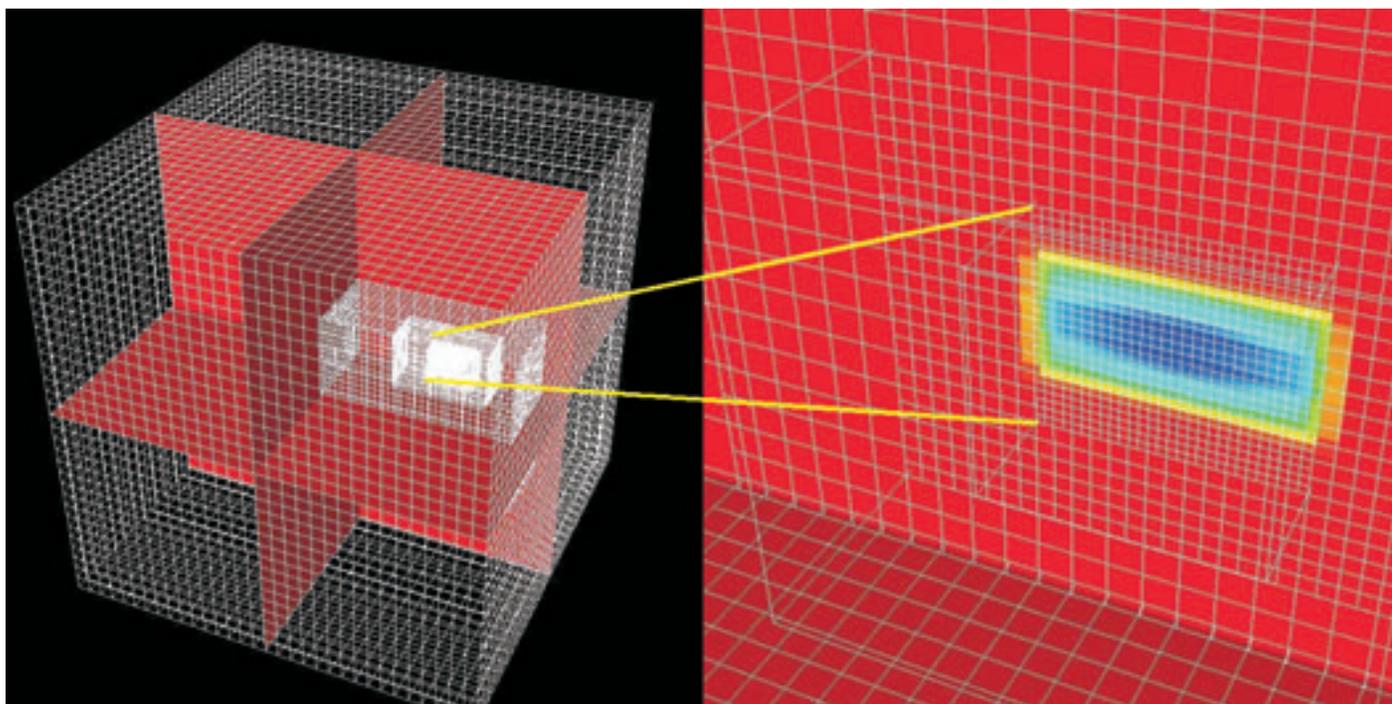


FIGURE 2. This illustration shows how computational problems are solved by dividing them into smaller pieces by covering them with a mesh. In this case, the fuel pellet for a fusion reactor is buried within the finest mesh which occupies less than 0.015 percent of the volume of the coarsest mesh. Using a technique known as adaptive mesh refinement, scientists can focus the power of a supercomputer on the most interesting part of a problem, such as the fuel pellet in a hot plasma.

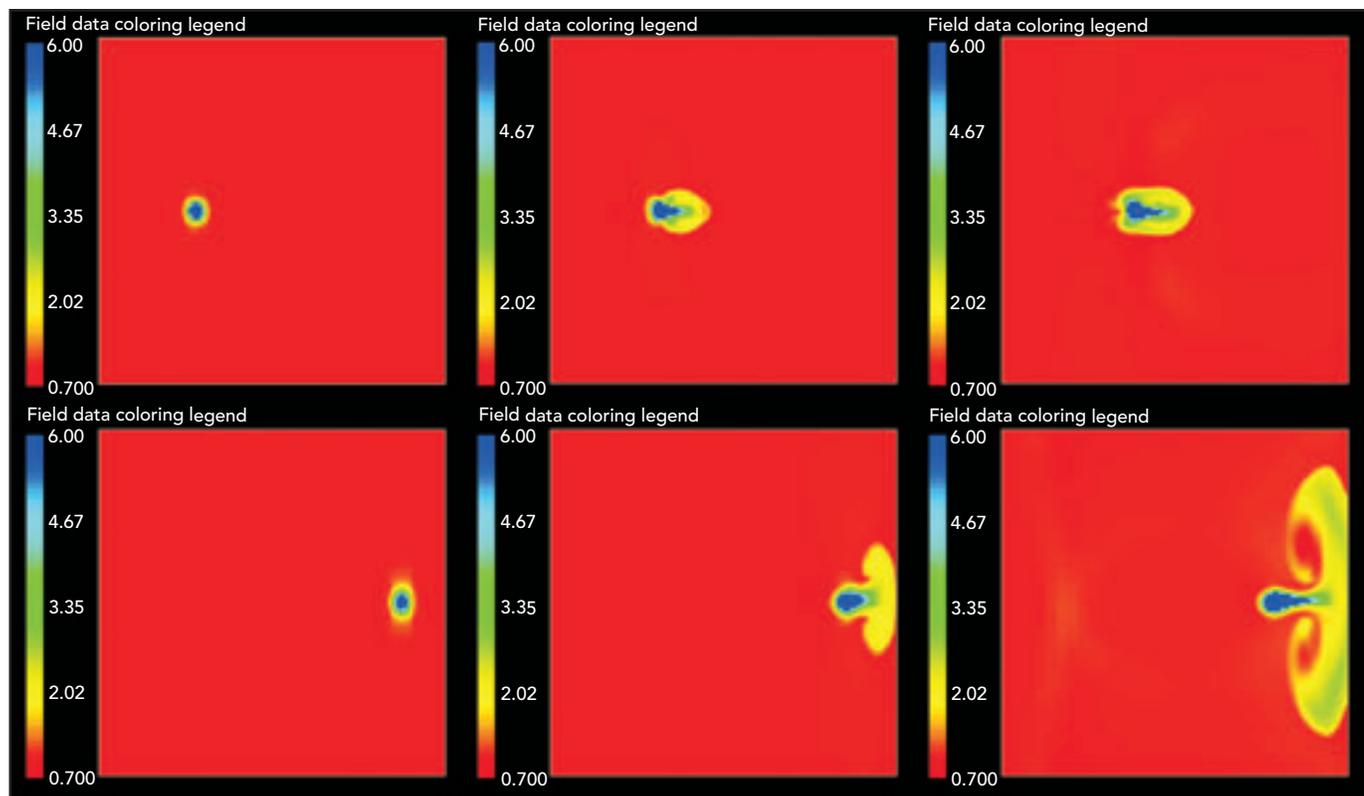


FIGURE 3. Time sequence of 2D slices from a 3D simulation of the injection of a fuel pellet into a tokamak plasma. Injection from outside the torus (bottom row, injection from right) results in the pellet stalling and fuel being dispersed near the plasma boundary. Injection from inside the torus (top row, injection from left) achieves fuel distribution in the hot central region as desired.

the ambient plasma, and the long-distance effects of electron heat transport all pose severe numerical challenges. To overcome these difficulties, Samtaney and his collaborators used an algorithmic method called adaptive mesh refinement (AMR), which incorporates a range of scales that change dynamically as the calculation progresses. AMR allowed this simulation to run more than a hundred times faster than a uniform-mesh simulation.

While these first calculations represent an important advance in

methodology, Samtaney's work on pellet injection is only beginning. "The results presented in this paper did not include all the detailed physical processes which we're starting to incorporate, along with more realistic physical parameters," he said. "For example, we plan to develop models that incorporate the perpendicular transport of the ablated mass. We also want to investigate other launch locations. And, of course, we'll have to validate all those results against existing experiments."

This pellet injection model will

eventually become part of a comprehensive predictive capability for ITER, which its supporters hope will bring fusion energy within reach as a commercial source of electrical power.

Burning Questions

**New 3D Simulations Are Closing in on the Holy Grail of Combustion Science:
Turbulence—Chemistry Interactions**

Controlling fire to provide heat and light was one of humankind's first great achievements, and the basic chemistry of combustion – what goes in and what comes out – was established long ago. But a complete quantitative understanding of what happens during the combustion process has remained as elusive as the ever-changing shape of a flame.



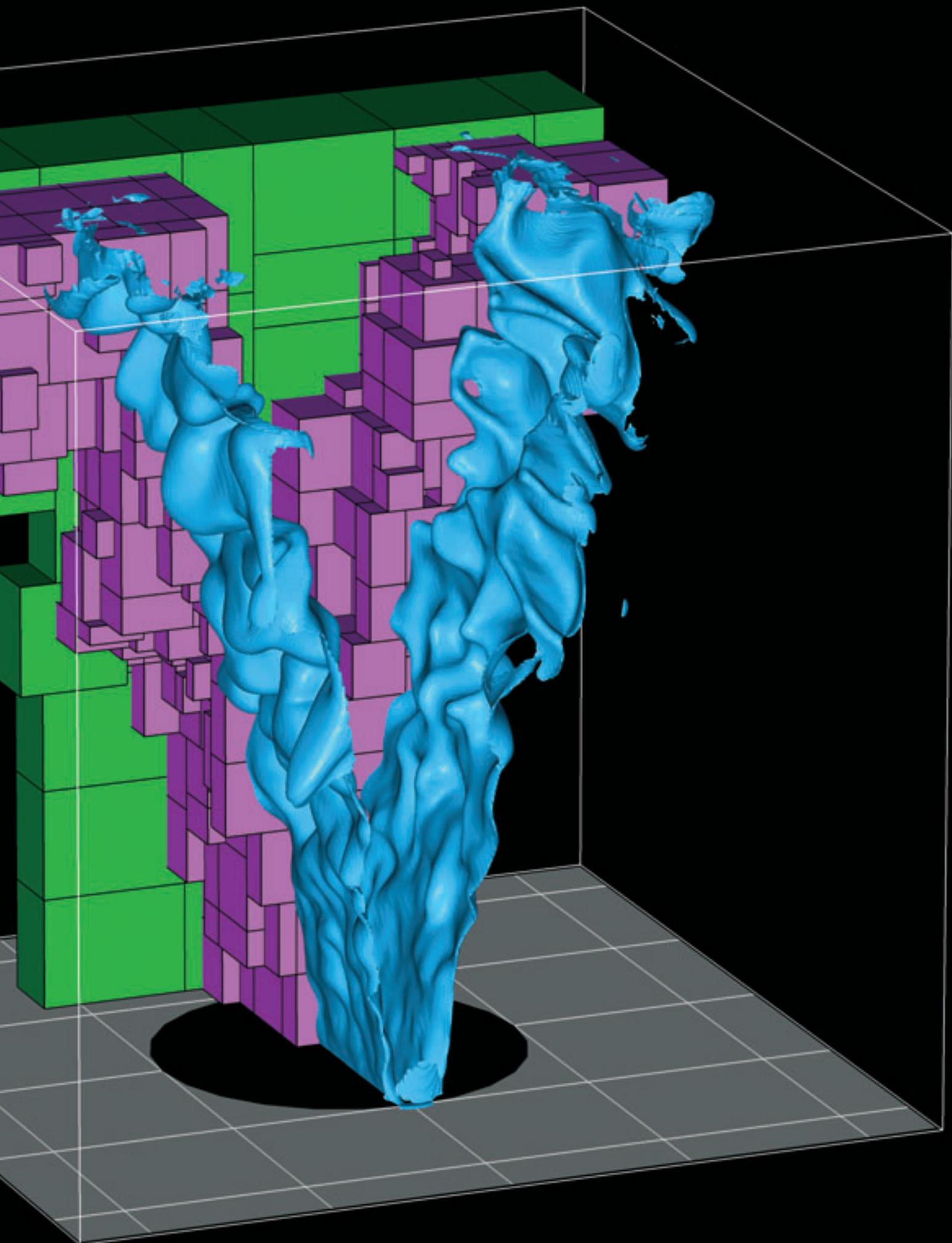




FIGURE 1. The calculated surface of a turbulent premixed laboratory methane flame.

Even a simple fuel like methane (CH_4), the principal component of natural gas, burns in a complex sequence of steps. Oxygen atoms gradually replace other atoms in the hydrocarbon molecules, ultimately leaving carbon dioxide and water. Methane oxidation involves about 20 chemical species for releasing energy, as well as many minor species that can become pollutants. Turbulence can distort the distribution of species and the redistribution of thermal energy which are required to keep the flame burning. These turbulence–chemistry interactions can cause the flame to burn faster or slower and to create more or less pollution.

The holy grail of combustion science has been to observe these turbulence–chemistry effects. Over the past few decades amazing progress has been made in observational techniques, including the use of lasers to excite molecules of a given species and produce a picture of the chemical distribution. But laser imaging is limited in the species and concentrations that can be reliably observed, and it is difficult to obtain simultaneous images to correlate different chemical species.

Because observing the details of combustion is so difficult, progress in combustion science has largely coincided with advances in scientific computing. For example, while basic concepts for solving one-dimensional flat flames originated in the 1950s, it only became possible to solve the 1D

flame equations some 30 years later using Cray-1 supercomputers. Those calculations, which are routine on personal computers today, enabled a renaissance in combustion science by allowing chemists to observe the interrelationships among the many hypothesized reaction processes in the flame.

Simulating three-dimensional turbulent flames took 20 more years of advances in applied mathematics, computer science, and computer hardware, particularly massively parallel systems. But the effort has been worth it. New 3D simulations are beginning to provide the kind of detailed information about the structure and dynamics of turbulent flames that will be needed to design new low-emission, fuel-efficient combustion systems.

The first 3D simulation of a laboratory-scale turbulent flame from first principles – the result of a SciDAC-funded collaboration between computational and experimental scientists at Berkeley Lab – was featured on the cover of the July 19, 2005 Proceedings of the National Academy of Sciences (Figure 1). The article, written by John Bell, Marc Day, Ian Shepherd, Matthew Johnson, Robert Cheng, Joseph Grcar, Vincent Beckner, and Michael Lijewski, describes the simulation of “a laboratory-scale turbulent rod-stabilized premixed methane V-flame.” This simulation was unprecedented in several aspects – the number of chemical species included, the number of chemical reactions modeled, and the overall size of the flame.

This simulation employed a different mathematical approach than has typically been used for combustion. Most combustion simulations designed for basic research use compressible flow equations that include sound waves, and are calculated with small time steps on very fine, uniform spatial grids – all of which makes them very computationally expensive. Because of limited computer time, such simulations often have been restricted to only two dimensions, to

scales less than a centimeter, or to just a few carbon species and reactions.

In contrast, the Center for Computational Sciences and Engineering (CCSE), under Bell's leadership, has developed an algorithmic approach that combines low Mach-number equations, which remove sound waves from the computation, with adaptive mesh refinement, which bridges the wide range of spatial scales relevant to a laboratory experiment. This combined methodology strips away relatively unimportant aspects of the simulation and focuses computing resources on the most important processes, thus slashing the computational cost of combustion simulations by a factor of 10,000.

Using this approach, the CCSE team has modeled turbulence and turbulence-chemistry interactions for a three-dimensional flame about 12 cm (4.7 in.) high, including 20 chemical species and 84 fundamental chemical reactions. The simulation was realistic

enough to be compared directly with experimental diagnostics.

The simulation captured with remarkable fidelity some major features of the experimental data, such as flame-generated outward deflection in the unburned gases, inward flow convergence, and a centerline flow acceleration in the burned gases (Figure 2). The simulation results were found to match the experimental results within a few percent. This agreement directly validated both the computational method and the chemical model of hydrocarbon reaction and transport kinetics in a turbulent flame.

The results demonstrate that it is possible to simulate a laboratory-scale flame in three dimensions without having to sacrifice a realistic representation of chemical and transport processes. This advance has the potential to greatly increase our understanding of how fuels behave in the complicated environments inside turbulent flames.

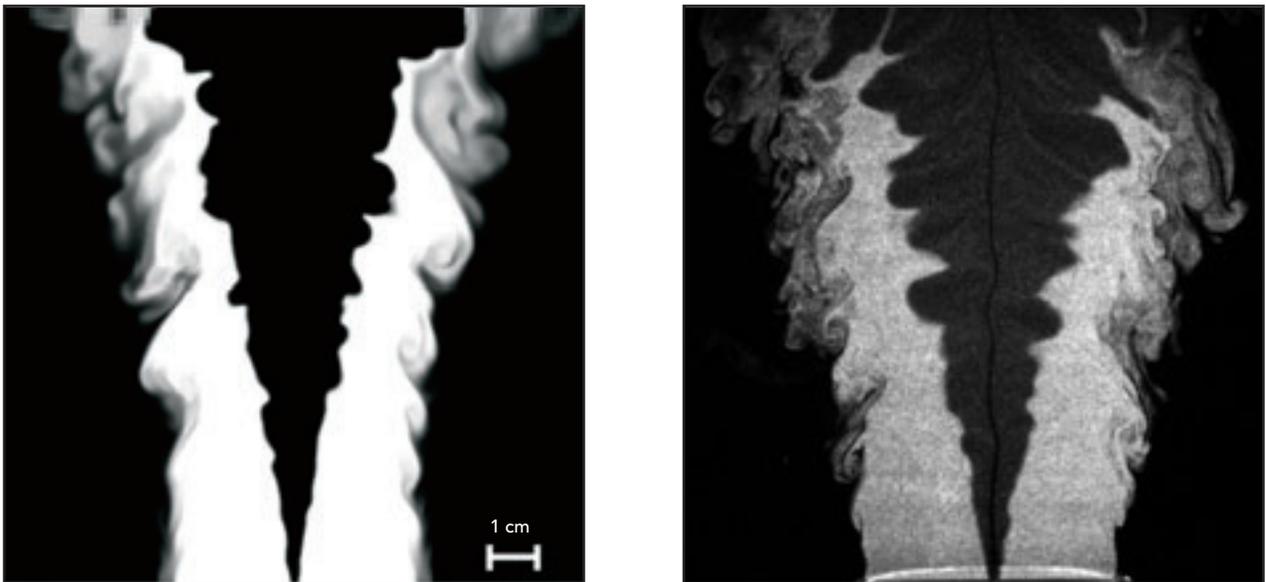


FIGURE 2. Left: A typical centerline slice of the methane concentration obtained from the simulation. Right: Experimentally, the instantaneous flame location is determined by using the large differences in Mie scattering intensities from the reactants and products to clearly outline the flame. The wrinkling of the flame in the computation and the experiment is of similar size and structure.

Pushing the Frontiers of Chemistry

Dirac's Vision Realized

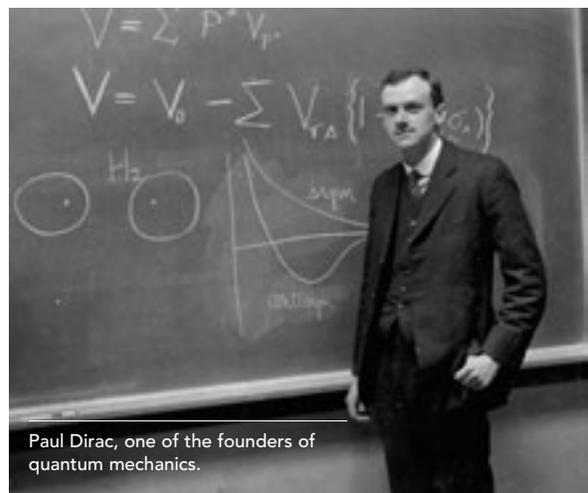
In most first-year chemistry classes at universities, students begin by learning how atoms of carbon and other elements form bonds with other atoms to form molecules, which in turn combine to form ourselves, our planet and our universe.

Understanding the properties and behavior of molecules, or better yet, being able to predict and control the behavior, is the driving force behind modern chemistry. The development of the theory of quantum mechanics in the 1920's made it possible that all the properties of molecules could be predicted. One of its founders, Paul Dirac, wrote in 1929 that "the difficulty is only that the exact application of these laws leads to equations much too complicated to be soluble."

Almost 80 years later this is still true, even using the most powerful supercomputers. For example, a scientist could solve for the behavior of a molecule with one electron moving in one dimension on a grid of roughly 100 points spread along that line. Since electrons in free space move in three dimensions, not one, the scientist would need 100^3 (one million)

calculations. Adding each extra electron then entails about a million times more work, a condition known as exponential scaling, so it is no wonder that the largest exact calculations still involve no more than about 15 electrons. Even with anticipated supercomputing advances over the next decade, only one or maybe two more electrons could be treated this way.

Quantum chemistry, however, provides scientists with the models to approximate these values without doing all the calculations. But there is a tradeoff, since researchers need to balance accuracy against feasibility. A very simple model can be applied to a giant molecular system, but the results will be uselessly inaccurate. As noted, a very accurate model, due to its complexity, may only be feasible for systems of up to 15 electrons. However, when scientists are



Paul Dirac, one of the founders of quantum mechanics.

studying molecular systems in a field such as nanotechnology, they are interested in systems with at least a few hundred electrons.

Realizing that a brute-force approach would not succeed, the Advanced Methods for Electronic Structure: Local Coupled Cluster Theory project was designed to develop new methods which strike novel compromises between accuracy and feasibility. The goal was to extend coupled cluster electronic structure theory to larger molecules. This was achieved by developing both new theory and new high-performance algorithms. The resulting method scales much better (at the 3rd power) than the old codes (to the 6th power), and can therefore be

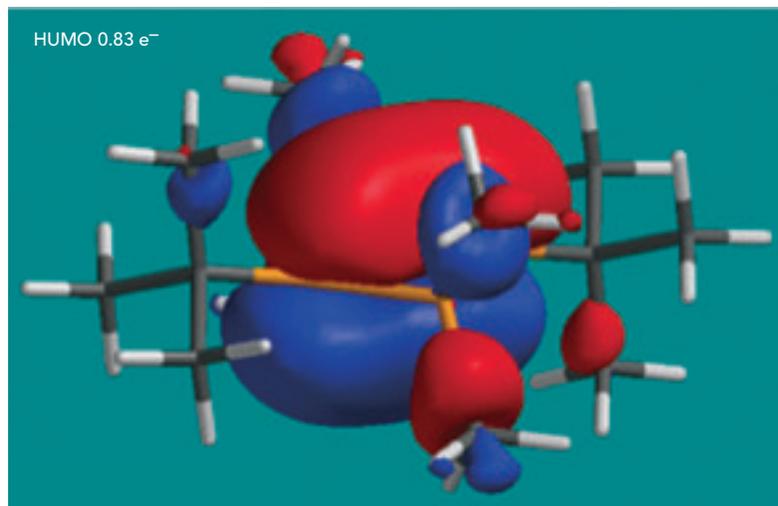
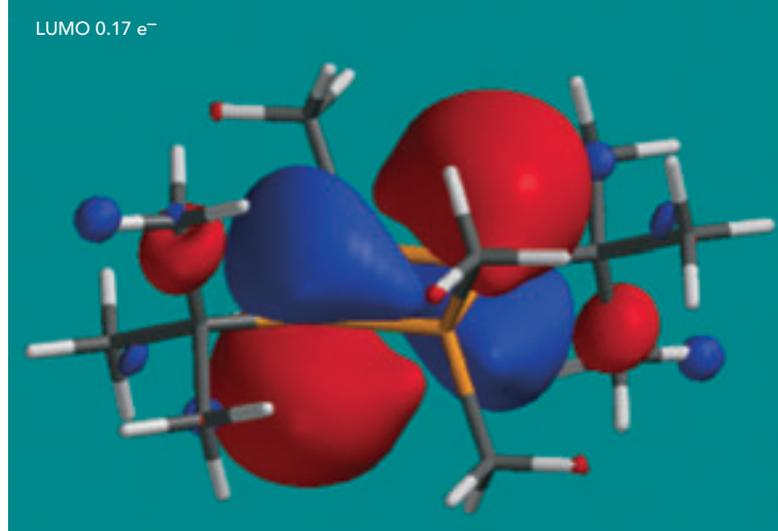


FIGURE 1. Graphical representation of the highest occupied molecular orbital (HOMO) and the lowest unoccupied level (LUMO) for the boron-phosphorus material. Instead of being fully empty, the LUMO contains 0.17 electrons, corresponding to 17 percent broken bond character.

applied to much larger molecules using existing supercomputers or even commodity computers.

One advantage the new method has over older methods is that it can be applied to systems where the electrons are highly correlated. In other words, the motions of the two electrons are closely tied to each other, much like two ballroom dancers or figure skaters who go through their motions almost like a single unit. Such a condition is mathematically much harder to describe than systems with weaker correlation, which can be imagined as two friends dancing separately at a rock concert. Being able to treat highly correlated electrons reliably allows scientists to study interesting molecules which would otherwise be difficult.

In particular, scientists have long been interested in studying molecules in which the chemical bonds are breaking, as opposed to the more common states where the molecule is either stable, with the bonds intact, or unstable, in which the bonds are broken. In bond-breaking, the electrons are strongly correlated. Scientists are interested in studying molecules as they are on their way to breaking into fragments because this may give insights that allow the design of useful materials in areas ranging from molecular electronics and spintronics to self-assembly of functional nanostructures.

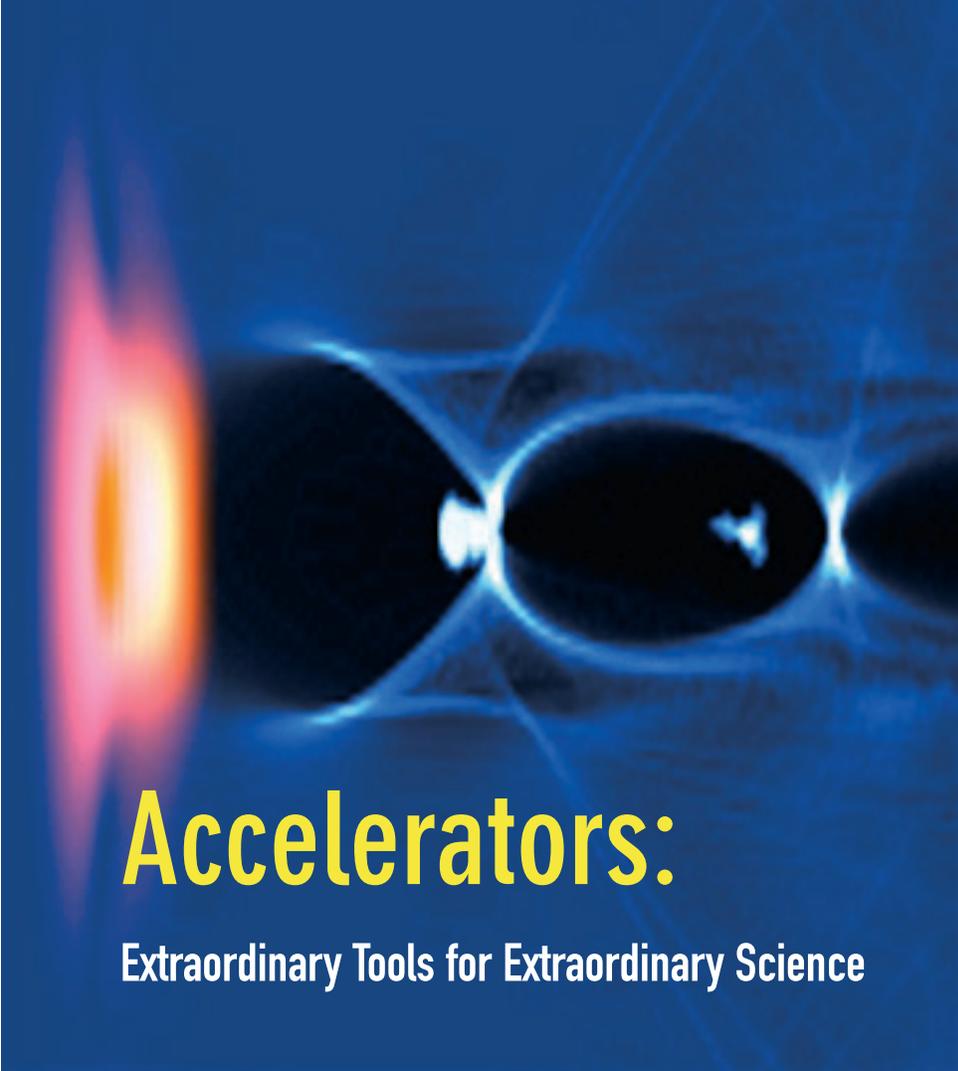
The project team studied a boron-phosphorus based material that caused excitement when it was reported as the first indefinitely stable

singlet diradical with the characteristics of broken bonds. The species was found computationally to have around 17 percent broken bond character, rather than the much higher fraction believed by experimentalists. This explained the stability of the material, and the origin of the 17 percent result could be understood from the role of neighboring groups, which shows the ability to tune the reactivity of the molecule by chemical design.

Project leader Martin Head-Gordon notes that the algorithms are currently not suitable for all problems, but that the team is looking to develop them into general purpose codes, and is working on improved successors. He predicts that the project's work will make its way into the mainstream within five years, bringing new computational chemistry capabilities to many of the 40,000 chemists who use such applications. Already, project members are getting inquiries from experimental chemists asking if they can use their codes to study specific molecules.

This has led to collaborations investigating the diradical character of what may be the world's longest carbon-carbon bonds as well as work in progress on triple bonds between heavier elements such as tin and germanium. Chemists are interested in whether these heavy elements can form triple bonds the way carbon does. Initial results indicate they do, but the behavior of the resulting molecules is very different.

Computation plays a key part in interpreting what experimentalists find and also in predicting what their experiments will produce, according to Head-Gordon, who has a joint appointment as a chemistry professor at the University of California, Berkeley, and DOE's Lawrence Berkeley National Laboratory. "High performance computing, together with new theory and algorithms, allows us to push the frontiers of chemistry and go where previous generations of chemists have not been able to go," Head-Gordon said.

A blue-toned simulation of a particle accelerator, showing a central beam pipe with two circular cross-sections. A bright, glowing orange and red beam enters from the left, passing through the first cross-section and continuing towards the second. The background is a dark blue field with faint, glowing lines representing the magnetic fields or particle paths.

Accelerators:

Extraordinary Tools for Extraordinary Science

Particle accelerators are some of the most powerful experimental tools available for basic science, providing researchers with insight into the basic building blocks of matter and enabling some of the most remarkable discoveries of the 20th century. They have led to substantial advances in applied science and technology, such as nuclear medicine, and are being studied for potential application to problems related to energy and the environment.

Given the importance of particle accelerators to the United States' scientific, industrial and economic competitiveness, bringing the most advanced high performance computing tools to bear on accelerator design, optimization and operation is in the national interest.

Within the DOE Office of Science, particle accelerators have enabled remarkable scientific discoveries and important technological advances that span several programs. In the High Energy Physics and Nuclear Physics programs, experiments associated with high-energy accelerators have led to important discoveries about elementary particles and the fundamental forces of nature, quark dynamics, and nuclear structure. In the Basic Energy Sciences and the Biological and Environmental Research programs, experiments with synchrotron light sources and spallation neutron sources have been crucial to advances in the materials, chemical and biological sciences. In the Fusion Energy Sciences program, great strides have been made in developing heavy-ion particle accelerators as drivers for high energy density physics research and ultimately inertial fusion energy. The importance of accelerators to the Office of Science mission is evident from an examination of the DOE "Facilities for the Future of Science: A Twenty-Year Outlook." Of the 28 facilities listed, 14 involve new or upgraded accelerator facilities.

The SciDAC Accelerator Science and Technology (AST) modeling project was a national research and development effort aimed at establishing a comprehensive terascale simulation environment needed to solve the most challenging problems in 21st century accelerator science and technology. The AST project had three focus areas: computational beam dynamics, computational electromagnetics, and modeling advanced accelerator concepts. The tools developed under this program are now being used by accelerator physicists and engineers across the country to solve the most challenging problems in accelerator design, analysis, and optimization. The following examples of scientific and engineering accomplishments achieved

by the AST project illustrate how the project software is already improving these extraordinary tools for extraordinary science.

Driving Discovery in Existing Accelerators

DOE operates some of the world's most productive particle accelerators, and one component of the AST project focused on developing scientific codes to help researchers get even more science out of these facilities.

As part of the SciDAC Accelerator Science and Technology project, the Fermilab Computational Accelerator Physics group has developed the Synergia framework. Integrating and extending existing accelerator physics codes in combination with new codes developed by the group, Synergia is designed to be a general-purpose framework with an interface that is accessible to accelerator physicists who are not experts in simulation.

In recent years, accurate modeling of beam dynamics in high-current, low-energy accelerators has become necessary because of new machines under consideration for future applications, such as the High Energy Physics neutrino program, and the need to optimize the performance of existing machines, such as the Spallation Neutron Source and the Fermilab Booster. These machines are characterized by high currents and require excellent control of beam losses. A common problem in accelerators is that the particles can stray from the beam core, creating what is known as a "halo," which can lead to beam loss. Understanding how the accelerator design and various physical phenomena (such as the space-charge effect) affect this halo formation is an essential component of accelerator modeling.

Several computer simulations of space-charge effects in circular accelerators using particle-in-cell techniques have been developed. Synergia is a

package for state-of-the-art simulation of linear and circular accelerators with a fully three-dimensional treatment of space charge. Space-charge calculations are computationally intensive, typically requiring the use of parallel computers.

Synergia was designed to be distributable to the particle accelerator community. Since compiling hybrid code can be a complicated task which is further complicated by the diverse set of existing parallel computing environments, Synergia includes a "build" system that allows it to be compiled and run on various platforms without requiring the user to modify the code.

When Beams Collide

While some accelerator research involves firing a beam at a stationary target, other accelerators are used to generate particle beams which are targeted to collide with each other, resulting in millions of particles flying out from the collision. High intensity, tightly focused beams result in high "luminosity," a parameter which is proportional to the number of events seen in an accelerator's detectors. But high luminosity also leads to increased beam disruption due to "beam-beam" effects, which consequently lowers the luminosity.

To better study these effects, a parallel three-dimensional, particle-in-cell code called BeamBeam3D was created to model beam-beam effects of colliding beams at high energy ring colliders. The resulting information can now be used to help accelerator operators determine the optimum parameters for colliders to maximize luminosity and hence maximize scientific output. Under SciDAC, BeamBeam3D was used to model colliding beams in several DOE accelerators including the Fermilab Tevatron, the Relativistic Heavy Ion Collider (RHIC) at Brookhaven National Laboratory, the PEP-II B-factory at Stanford Linear Accelerator Center, and the

soon-to-be-operating Large Hadron Collider (LHC) at CERN. In the future, BeamBeam3D is expected to be used to model beam-beam effects in the proposed International Linear Collider.

BeamBeam3D simulations of RHIC are relevant to both RHIC operations and to the future operation of the LHC (due to the similar beam-beam conditions). In regard to RHIC, the code was used to study beam-beam effects in three areas. The first are the coherent beam-beam effects. One problem is that beam-beam effects in hadron colliders, such as the one at RHIC, can create an oscillation under which the beams become unstable; this instability represents a roadblock to increasing the luminosity. Using BeamBeam3D, SciDAC researchers modeled the coherent beam-beam effects first during collisions of single bunches of particles, then during multiple bunch collisions. In the latter case, each beam has three bunches that are coupled to three bunches in the opposite beam at four interaction points. Using the BeamBeam3D code running on high performance computers, project members were able to calculate at which point the beams would become unstable. However, the group also found that by appropriately arranging the accelerator machine lattice parameters of the two rings at RHIC and generating sufficient tune separation between the two beams, oscillation could be controlled to the extent that there is no risk of instability.

Second, using BeamBeam3D, the team studied emittance growth (a reduction in beam quality that affects the luminosity) in beams which are offset from one another. The team found that, while static offsets do not cause significant emittance growth over short time periods, the impact of offsets is much larger when they are time-modulated (due, for example, to mechanical vibrations in the final-focus magnets.) This finding provides a potential mechanism to account for the extra emittance growth observed

during the machine operation.

Lastly, the team used BeamBeam3D to study long-range beam-beam effects at RHIC. Simulations performed at the energy level at which particles are injected into the accelerator showed a strong sensitivity of the long-range beam-beam effects to the machine tunes; this sensitivity was also observed in the experiments. The team subsequently modeled the long-range beam-beam effects at the higher energy levels at which the particles collide. Such simulations are providing insight and a means to numerically explore the parameter space for the future wire beam-beam compensation experiment planned at RHIC.

Shaping the Future of Accelerator Design

Although their scientific value is substantial, accelerators are very expensive and difficult to design, build and operate. As a result, the scientific community is now looking at international collaboration to develop the next generation of accelerators. One prime example is the International Linear Collider (ILC), the highest priority future project in the worldwide high energy physics community for probing into the fundamental nature of matter.

Presently, a large team of accelerator physicists and engineers from Europe, Asia and North America is working on the design of this tera-electronvolt-scale particle accelerator under a unified framework, called the Global Design Effort (GDE), to make the most of limited R&D resources. Under the AST project, modeling tools were developed and used to study the effectiveness of proposed designs and look for better solutions.

In the ILC design, two facing linear accelerators (linacs), each 20 kilometers long, hurl beams of electrons and positrons toward each other at nearly the speed of light. Each nanometer-scale size beam contain-

ing ten billion electrons or positrons is accelerated down the linac by superconducting accelerating cavities, which give them more and more energy till they meet in an intense crossfire of collisions at the interaction point. At these energies, researchers anticipate significant discoveries that will lead to a radically new understanding of what the universe is made of and how it works. The energy of the ILC beams can be adjusted to home in on elementary particle processes of interest.

A critical component of the ILC linac is the superconducting radio frequency (SRF) accelerating cavity. The design, called TESLA, was created in the early 1990s and has been the focus of more than a decade of experimental R&D effort. Recently, a new cavity design has been proposed which has a higher accelerating gradient and 20 percent less cryogenics loss over the TESLA design – an important consideration as the cavities must be cooled to extremely low temperatures (2 degrees Kelvin, or about -456 degrees Fahrenheit) to maintain superconductivity.

Accompanying intense experimental efforts at the KEK accelerator center in Japan and Jefferson Lab in Virginia, the development of this low-loss (LL) cavity has been greatly facilitated by the new parallel electromagnetic codes developed at the Stanford Linear Accelerator Center (SLAC) under the AST project. In collaboration with the SciDAC Terascale Optimal PDE Solvers Integrated Software Infrastructure Center (TOPS ISIC), the SLAC team created a nonlinear 3D parallel finite element eigensolver with which, for first time, one can directly solve for the damped dipole modes in the 3D LL cavity, complete with higher-order mode (HOM) couplers. HOM damping is essential for stable transport of “long bunch trains” of particles as they race down the ILC linacs and

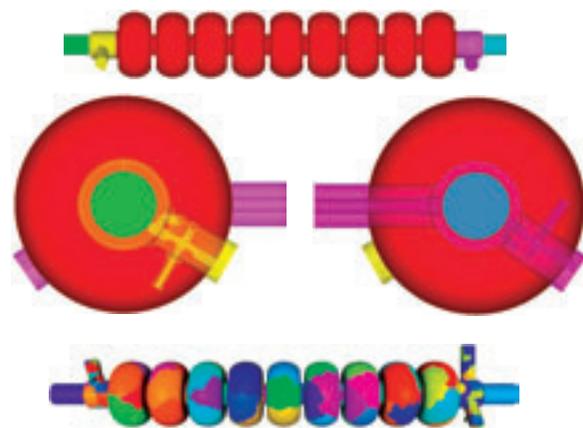


FIGURE 1. Using software developed under SciDAC, accelerator scientists developed this model of the new low-loss cavity for the proposed International Linear Collider. The different colors illustrate how various sections were modeled in parallel on multiple processors. The end perspectives show the geometry details in the couplers which could not previously be modeled because of the disparate length scales.

for preserving low emittance of particles from the beams, which can reduce the beam quality.

Previously, evaluating HOM damping in an SRF cavity took years to complete either experimentally or numerically. Today, using the tools developed under SciDAC, HOM calculations can be done in a matter of weeks with the new parallel solver running on NERSC's IBM SP3 and NCCS's Cray X1E supercomputers. In addition, the use of tetrahedral grid and higher order basis functions has enabled solutions in the complex cavity geometry to be obtained with unprecedented accuracy. As a result, costly and time-consuming prototyping by the trial-and-error approach is avoided as the LL cavity can be computationally designed as close to optimal as possible.

Advanced Accelerators: Smaller, Faster, Futuristic

DOE is also looking at future-generation accelerators, in which lasers and plasmas would be used instead of electromagnetic cavities to



FIGURE 2. SciDAC codes were used to analyze the experiments in two of three featured articles on compact particle accelerators in the September 30, 2004 issue of Nature. The cover image was created using the VORPAL code developed by Tech-X Corp. partly with SciDAC support.

accelerate particles, which could result in accelerators with 1,000 times the performance of current technology. The challenge is to control these high-gradient systems and then to string them together. Such technologies would enable the development of ultra-compact accelerators, which would be measured in meters, not kilometers. However, experiments to date have yielded acceleration over very short distances (millimeters to centimeters) resulting in beams of modest quality.

One approach being explored is a plasma-wakefield accelerator, in which a drive beam – either an intense particle beam or laser pulse – is sent through a uniform plasma. This creates a space-charge wake on which a trailing beam of particles can surf. SciDAC investigators in the Accelerator Science and Technology (AST) project have developed a suite of particle-in-cell (PIC) codes to model such devices.

SciDAC support has enabled the development of four independent, high-fidelity, particle-in-cell (PIC) codes: OSIRIS (fully explicit PIC), VORPAL (fully explicit PIC plus ponderomotive guiding center), QuickPIC (quasi-static PIC plus ponderomotive guiding center), and UPIC (a framework for rapid construction of new codes such as QuickPIC). Results from these codes were included in three articles in Nature and eight in Physical Review Letters. For example, members of the SciDAC AST project have performed large-scale simulations using the codes OSIRIS and VORPAL to help interpret laser- and plasma-wakefield accelerator experiments and to gain insight into the acceleration mechanism.

Researchers at Lawrence Berkeley National Laboratory took a giant step toward realizing the promise of laser wakefield acceleration, by guiding and controlling extremely intense laser beams over greater distances than ever before to produce high-quality, energetic electron beams. The experimental results were analyzed by running the VORPAL plasma simulation code, devel-

oped with SciDAC support, on supercomputers at DOE's NERSC. This allowed scientists to see details of the evolution of the experiment, including the laser pulse breakup and the injection of particles into the laser-plasma accelerator. This allows them to understand how the injection and acceleration occur in detail so that the experiment's designers can figure out how to optimize the process.

The results of the experiment and simulations were published as the cover article in the Sept. 30, 2004 issue of Nature (Figure 2).

Great progress has also been made in the development of reduced description models for laser/plasma simulation. One such example, QuickPIC, has been shown for some problems to provide answers as accurate as OSIRIS but with two to three orders of magnitude less computation time. Before the SciDAC effort, a full-scale simulation of a 1 TeV afterburner would have required 5 million node hours on a supercomputer (and thus was not done); after SciDAC, using QuickPIC, the simulation was done in 5,000 node hours on a cluster system.

Lattice QCD

Improved Formulations, Algorithms and Computers Result in Accurate Predictions of Strong Interaction Physics

The long-term goals of high energy and nuclear physicists are to identify the fundamental building blocks of matter and to determine the interactions among them that give rise to the physical world we observe. Major progress towards these goals has been made through the development of the Standard Model of high energy physics. The Standard Model consists of two quantum field theories: the Weinberg-Salam Theory of the electromagnetic and weak interactions, and quantum chromodynamics (QCD), the theory of the strong interactions.

The Standard Model has been enormously successful in explaining a wealth of data produced in accelerator and cosmic ray experiments over the past twenty-five years. However, our knowledge of the Standard Model is incomplete because it has been difficult to extract many of the most interesting predictions of QCD, those that depend on the strong coupling domain of the theory. The only way to extract these predictions from first principles and with controlled errors is through large-scale numerical simulations. These simulations are needed to obtain a quantitative understanding of the physical phenomena controlled by the strong interactions, to determine a number of the basic parameters of the Standard Model, and to make precise tests of the Standard Model's range of validity.

Despite the many successes of the Standard Model, it is believed that to understand physics at the shortest distances or highest energies, a more general theory, which unifies all four of the fundamental forces of nature, will be required. However, to determine where the Standard Model breaks down and new physics is required, one must first know what the Standard Model predicts.

Numerical simulations of QCD address problems that are at the heart of the Department of Energy's large experimental programs in high energy and nuclear physics. Among the major goals are (1) to calculate the effects of strong interactions on weak interaction processes to an accuracy needed to make precise tests of the Standard Model; (2) to determine the properties of strongly interacting matter under extreme conditions, such as those that existed in the very early development of the universe and are created today in relativistic heavy ion collisions; and (3) to calculate the masses of strongly interacting particles and obtain a quantitative understanding of their internal structure.

According to QCD, the funda-

mental building blocks of strongly interacting matter are quarks. Interactions among quarks are mediated by gluons, in a manner analogous to the way photons mediate the electromagnetic interactions. QCD interactions are so strong that one does not directly observe quarks and gluons under ordinary laboratory conditions. Instead, one observes strongly interacting particles, such as protons and neutrons, which are bound states of quarks and gluons. QCD was initially formulated in the four-dimensional space-time continuum; however, to carry out numerical simulations, one must reformulate the theory on a discrete four-dimensional lattice—hence the name *lattice QCD*.

Major progress has been made in the numerical study of QCD during the course of the SciDAC program through the introduction of improved formulations of QCD on the lattice, coupled with improved algorithms and major advances in the capabilities of high performance computers. These developments have allowed physicists to fully include the effects arising from the polarization of the QCD ground state due to the creation and annihilation of quark-antiquark pairs.

The inclusion of vacuum polarization effects has been the greatest challenge to performing accurate numerical calculations of QCD. Their importance is illustrated in Figure 1, where lattice QCD calculations of the masses of a number of strongly interacting particles, and the decay constants of the π and K mesons, are compared with their experimental values. In each case, agreement with experiment was found within statistical and systematic errors of 3% or less when vacuum polarizations were included (right panel), but this was not the case when the uncontrolled approximation of ignoring vacuum polarization was made (left panel). This work was described in a “News and Views” article in *Nature*, as well as in a “News Focus” article in *Science*. Within the high energy

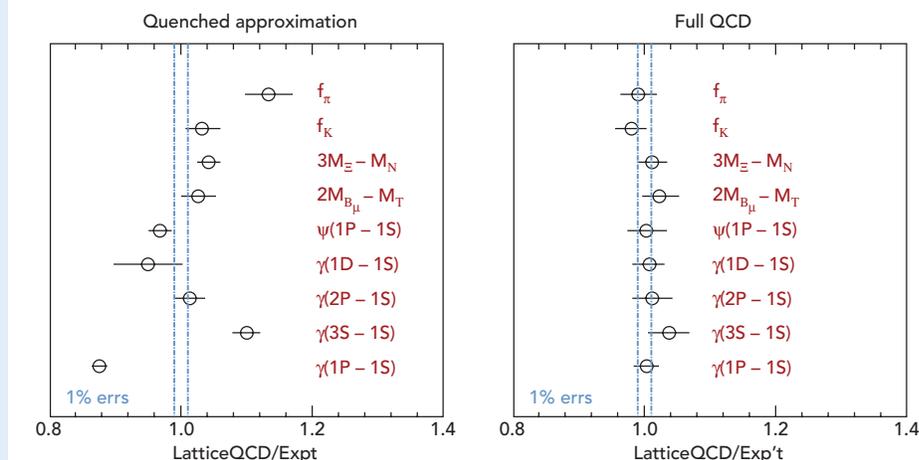


FIGURE 1. The ratio of several quantities calculated in lattice QCD to their experimental values. The panel on the left shows results from the quenched approximation, and that on the right from full QCD.

physics community, it has been featured in *Fermi News Today*, *Physics Today*, and in a cover article in the *CERN Courier*.

A number of other important validations of lattice QCD methods have also been enabled by the SciDAC Program. They include the determination of the strong interaction coupling constant in agreement with but with somewhat smaller errors than the world average from a number of different experimental determinations; the determination of the nucleon axial charge, which governs the β decay of a free neutron into a proton, electron, and neutrino; and the calculation of the Cabibbo-Kobayashi-Maskawa (CKM) matrix element V_{us} to an accuracy comparable with experiment. (The CKM matrix describes how quarks couple to the weak interactions; its elements are fundamental parameters of the Standard Model.)

During the course of the SciDAC Program, the lattice QCD community moved from the validation of techniques, through the calculation of quantities that are well known experimentally, to the successful prediction of quantities that had not yet been measured. One important example was the prediction of the mass of the B_c meson. This exotic particle, which consists of a bottom

quark and a charmed anti-quark, was first observed in 1998, but its mass was only poorly measured. With the aid of SciDAC-funded prototype clusters at Fermilab, the mass of the B_c meson was calculated to be 6304 ± 20 MeV, a dramatic improvement in accuracy over previous lattice calculations. Soon after this result was made public, the CDF experiment at Fermilab’s Tevatron finished a new, precise measurement of the mass: 6287 ± 5 MeV, confirming the prediction from lattice QCD. The fine agreement was covered in the *New Scientist*, *The Scotsman* newspaper, and a “News and Views” article in *Nature*. The success of the lattice calculation was named one of the top physics stories of 2005 by *Physics News Update*, which described it as “the best-yet prediction of hadron masses using lattice QCD.”

One of the major objectives of the field of lattice QCD is to determine the decay properties of pseudoscalar mesons with one light and one heavy quark. Strong interaction effects in leptonic decays are characterized by decay constants, while in semileptonic decays they are characterized by various form factors. The decay constants and form factors for B and B_s mesons, which contain heavy b quarks, play a critical role in

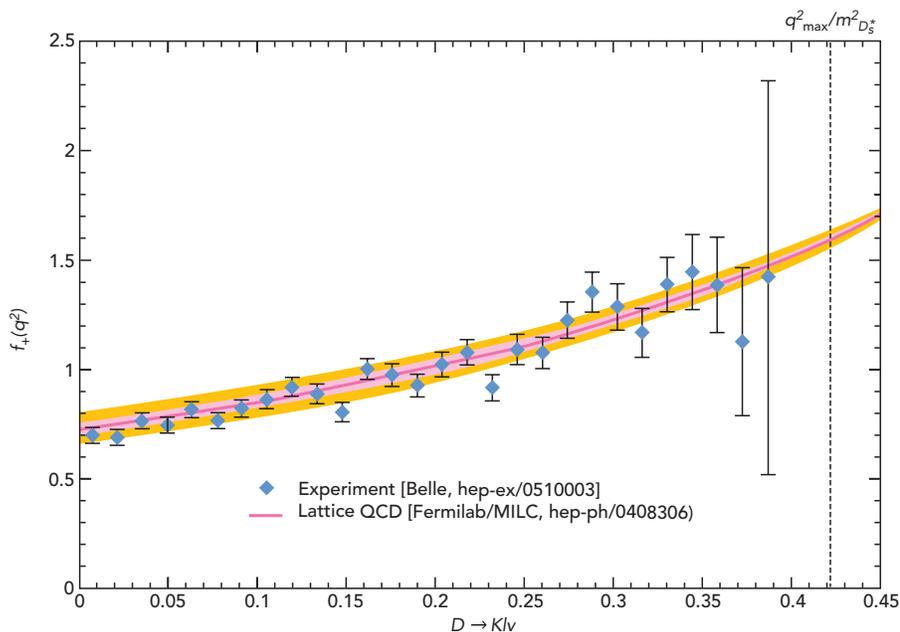


FIGURE 2. The semileptonic form factor $f_+(q^2)$ for the decay of a D meson into a K meson, a lepton, and a neutrino, as a function of the momentum transfer to the leptons q^2 . The orange curve is the lattice result, and the blue points are the experimental results of the Belle Collaboration.

tests of the Standard Model that are currently a major focus of the experimental program in high energy physics. These quantities are very difficult to measure experimentally, so accurate lattice calculations of them would be of great importance. On the other hand, the decay constants and form factors of D and D_s mesons, which

contain heavy c quarks, are being measured to high accuracy by the CLEO-c Collaboration. Since the lattice techniques for studying mesons with c and b quarks are identical, these experiments provide an excellent opportunity to validate the lattice approach being used in the study of D and B decays.

The first lattice QCD results for the leptonic decay constants of the D and D_s mesons that fully took into account vacuum polarization effects were announced in June 2005. Within a few days of these results being made public, the CLEO-c Collaboration announced its experimental result for the decay constant of the D meson; and in April 2006 the BaBar Collaboration announced its determination of the decay constant of the D_s meson. In both cases the experiments confirmed the lattice QCD calculations with comparable errors. The lattice and experimental results for the decay of the D meson were the subjects of cover articles in the CERN Courier and the New Scientist.

Finally, the form factors that characterize the decay of a D meson into a K meson and leptons were predicted with lattice QCD in work that was also enabled by the SciDAC Program. The results were subsequently confirmed in experiments by the Focus and Belle collaborations. The lattice results are compared with experimental ones from Belle in Figure 2. The excellent agreement between theory and experiment provides one more piece of evidence that lattice QCD calculations are able to make accurate predictions of strong interaction physics. This work too was featured in Fermi News Today.

Scientific Challenge Codes: Designing Scientific Software for Next-Generation Supercomputers

For more than 40 years, the power of computer processors has doubled about every 18 months, following a pattern known as Moore's Law. This constant increase has resulted in desktop computers which today are more powerful than the supercomputers of yesteryear. Today's supercomputers, at the same time, offer computing power which enables researchers to develop increasingly complex and accurate simulations to address the most challenging scientific problems. However, the software applications for studying these problems have not kept pace with the advances in computational hardware.

Throughout the Office of Science research program are major scientific challenges that can best be addressed through advances in scientific supercomputing. These challenges include designing materials with selected properties, understanding and predicting global climate change, understanding and controlling plasma turbulence for fusion energy, designing new particle accelerators, and exploring basic questions in astrophysics.

A key component of the SciDAC program is the development of scientific challenge codes – new applications aimed at addressing key research areas and designed to take advantage of the capabilities of

the most powerful supercomputers, and to run as efficiently as possible to make the most effective use of those systems.

This is a daunting problem. Current advances in computing technology are typically driven by market forces in the commercial sector, resulting in systems designed for commerce, not scientific computing. Harnessing commercial computing technology for scientific research poses problems unlike those encountered in previous supercomputers, both in magnitude as well as in kind. This problem will only be solved by increased investments in computer software – in research and development on scientific simu-

lation codes as well as on the mathematical and computing systems software that underlie these codes.

In the following pages, descriptions of many of the software projects will illustrate how SciDAC has supported software development to benefit the offices of Basic Energy Research, Biological and Environmental Research, Fusion Energy Sciences and High Energy and Nuclear Physics.

Basic Energy Sciences: Understanding Energy at the Atomic Level

As one of the world's leading sponsors of basic scientific research, the Department of Energy has long supported investigations into how atoms interact, how they form molecules and how groups of atoms and molecules react with one another. Understanding the forces at work among the most basic building blocks of matter can provide greater insight into how energy is released, how waste products are generated during chemical processes and how new materials can be developed.

Within the Office of Science, the Basic Energy Sciences (BES) program supports such fundamental research in materials sciences and engineering, chemistry, geosciences and molecular biosciences. Basic research supported by the BES program touches virtually every aspect of energy resources, production, conversion, efficiency, and waste mitigation. Under the SciDAC program, a number of projects were funded to support computational chemistry.

Research in chemistry leads to advances such as efficient combustion systems with reduced emissions of pollutants, new processes for converting solar energy, improved catalysts for the producing fuels and chemicals and better methods for

environmental remediation and waste management.

Over the past 50 years, molecular theory and modeling have advanced from merely helping explain the properties of molecules to the point where they provide exact predictive tools for describing the chemical reactions of three- and four-atom systems, the starting point for many of these chemical processes. This is increasingly important as researchers seek to understand more complex molecules and processes such as combustion, which involves complex interactions of chemistry with fluid dynamics.

The advances in computational chemistry in recent years in providing accurate descriptions of increasingly complex systems have come as much from improvements in theory and software as from improved computational hardware. However, the computational requirements often outweigh the scientific gains. For example, electronic structure theory provides the framework for modeling complex molecular-level systems (from atoms to thousands of atoms) with increasing accuracy. But accurate electronic structure theories have computational costs that rise with the 6th power (or higher) of molecular size, so that 10 times the computing resources translates into treating a system less than 1.5 times bigger. Predictive modeling of such processes is currently beyond the capabilities of existing computational resources and computational methods.

Under the SciDAC program, BES supported a number of projects aimed at developing computational approaches to solving problems in the modeling of chemical processes that exceed current computational capabilities. These projects were selected to increase the accuracy of models, increase the size of systems which could be simulated and expand the capabilities of key applications already being used in the research community.

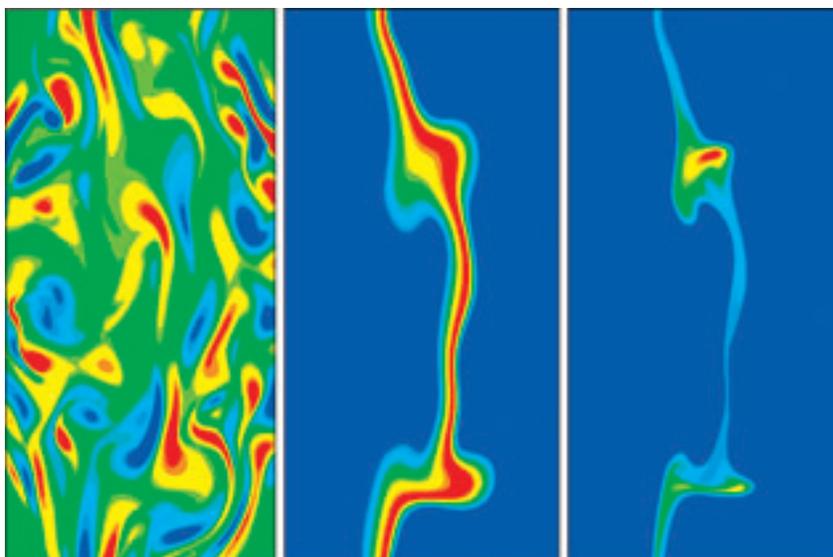


FIGURE 1: Simulations of turbulent nonpremixed ethylene-air flames reveal detailed structure of dynamics of the soot formation process. Shown from left to right are instantaneous images of the vorticity, temperature, and soot volume fraction.

Two of the projects focused on improving computational modeling of combustion, which is key to 80 percent of the energy production in the United States. Developing new methods to make combustion more efficient and cleaner will benefit the economy, the environment and the quality of life. With such a prominent role in energy production and use, combustion has long been an important research focus for the DOE.

But fully understanding combustion is an extremely complex problem involving turbulent flows, many chemical species and a continuing series of interdependent chemical processes and reactions. As computers have become more powerful, more detailed combustion simulations can be created to help scientists better understand the process.

The Terascale High-Fidelity Simulations of Turbulent Combustion with Detailed Chemistry project is a multi-university collaborative effort to develop a high-fidelity turbulent reacting flow simulation capability that can take advantage of the most powerful supercomputers available. The approach is based on direct numerical simulation

(DNS) to enable the highest accuracy and allow scientists to study the fine-scale physics found in turbulent reacting flows.

Under SciDAC, the simulation code named S3D has been enhanced with many new numerical algorithms and physical models to provide predictive capabilities for many of the detailed processes which occur during combustion, including thermal radiation, soot dynamics, spray injection and evaporation, and flame-wall interaction. The S3D code was used to perform detailed three-dimensional combustion simulations of flames in which fuel and oxygen are not premixed. This research could have applications in such areas as jet aircraft engines, where fuel and oxidizers are not premixed for safety reasons, and in direct-injection internal combustion engines. Under certain conditions, this type of combustion can suddenly and unexpectedly become extinguished, and this project is expected to lead to a better understanding of this problem, as well as re-ignition of extinguished flames. The team demonstrated the advanced DNS capability by undertaking several laboratory-scale simulations to high-

light fundamental aspects of turbulence-chemistry interaction occurring in many practical combustion systems.

Principal Investigators: Hong G. Im, University of Michigan; Arnaud Trouvé, University of Maryland; Christopher J. Rutland, University of Wisconsin; and Jacqueline H. Chen, Sandia National Laboratories

The Computational Facility for Reacting Flow Science project is aimed at advancing the state of the art in the understanding and prediction of chemical reaction processes, such as combustion, and their interactions with fluid flow. Such detailed reacting flow computations are computationally intensive, yet they provide information that is neces-

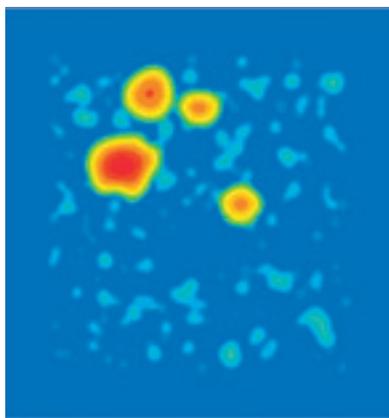


FIGURE 2. Reaction-diffusion high-order AMR computations of the propagation of random premixed hydrogen-oxygen ignition kernels in two dimensions. The temperature field is shown, indicating cold reactants (blue) and hot combustion products (red), separated by the propagating flame fronts.

sary for the understanding and prediction of turbulent combustion. The project's approach towards achieving this goal includes two broad focus areas. The first involves developing high-accuracy adaptive mesh refinement (AMR) algorithms and implementing them in a flexible software toolkit for reacting flow computations. The second involves the development of advanced

chemical analysis algorithms and software that enable both the extraction of enhanced physical understanding from reacting flow computations, and the development of chemical models of reduced complexity. The overall construction is being implemented in the context of the common component architecture (CCA) framework. The assembled software will be applied to targeted reacting flow problems, in two and three dimensions, and validated with respect to reacting flow databases at the Combustion Research Facility of Sandia National Laboratories.

Principal Investigator: Habib Najm, Sandia National Laboratories

Another set of projects focused on improving applications for modeling electronic structure. Electronic structure calculations, followed by dynamical and other types of molecular simulations, are recognized as a cornerstone for the understanding and successful modeling of chemical processes and properties that are relevant to combustion, catalysis, photochemistry, and photobiology.

The Advanced Methods for Electronic Structure project advanced the capabilities of quantum chemical methods to describe efficiently and with controllable accuracy the electronic structure, statistical mechanics, and dynamics of atoms, molecules and clusters. The project, which had goals of improving the

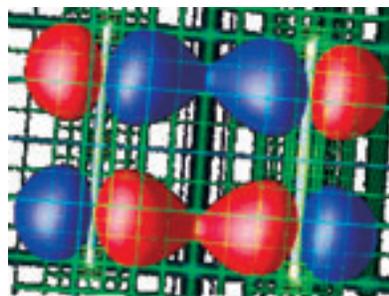


FIGURE 3. Molecular orbital of the benzene dimer with the adaptive grid and an isosurface.

speed, scalability and accuracy of electronic structure applications, had two thrusts.

(1) **The Multiresolution Quantum Chemistry – Guaranteed Precision and Speed** component worked to increase the accuracy of chemistry codes as they are scaled up to run on larger systems, while also reducing the amount of processing time needed to run the calculations. Additionally, computational chemistry is expected to benefit from the resulting computational framework that is substantially simpler than conventional atomic orbital methods, that has robust guarantees of both speed and precision, and that is applicable to large systems. In contrast, current conventional methods are severely limited in both the attainable precision and size of system that may be studied. The project achieved its goals for effective one-electron theories and is studying many-body theories. The electronic structure methods should be valuable in many disciplines, and the underlying numerical methods are broadly applicable.

Principal Investigator: Robert Harrison, Oak Ridge National Laboratory

(2) The component to develop **Advanced Methods for Electronic Structure: Local Coupled Cluster Theory** was designed to extend coupled cluster (CC) electronic structure theory to larger molecules (“coupled cluster” is a numerical technique used for describing many-body systems). This method was achieved by developing novel “fast” coupled cluster algorithms that are much more scalable with respect to system size to more fully realize the potential of high performance computing for treating larger molecular systems. CC methods are well established as the wave-function-based electronic structure method of choice. However, even the simplest CC method, incorporating just correlations between pairs of electrons (double substitutions), still requires computational costs that

scale with the 6th power of the size of the molecule. This research focused on defining new and powerful “local correlation” models that reduce scaling of coupled cluster calculations, and developing effective algorithms for their implementation.

Principal Investigator: Martin Head-Gordon, Lawrence Berkeley National Laboratory

The project for **Advancing Multi-Reference Methods in Electronic Structure Theory** researched and developed new models for investigating detailed mechanisms of chemical reactions for complex systems, including biomolecular processes, such as the interaction of protein molecules. Specifically, the project worked on the development of highly scalable electronic structure codes that are capable of predicting potential energy surfaces of very high accuracy, which is important since potential energy surfaces determine how atoms and chemical bonds rearrange during chemical reactions. The ability to study extended systems containing tens to thousands of atoms with reasonable accuracy is of paramount importance. The development of methods to adequately treat such systems requires highly scalable, highly correlated electronic structure methods interfaced with classical methods and mesoscale codes.

Principal Investigator: Mark S. Gordon, Ames Laboratory

The **Advanced Software for the Calculation of Thermochemistry, Kinetics and Dynamics** project consisted of two integrated programs to develop both scalable kinetics/dynamics software and infrastructure software. The overall thrust is to develop software to efficiently provide reliable thermochemistry (heat-related chemical reactions), kinetics, and dynamics for large molecular systems. Kinetics/dynamics studies compute how molecules move over a potential energy sur-

face. The first goal is to provide better thermochemistry for larger systems and improve the performance of related components in the Columbus computational chemistry software system. The second goal is to develop highly parallelized quantum dynamics and quantum kinetics software. Lastly, common component architecture techniques will be used to integrate kinetics and electronic structure software into a package that will allow users to compute kinetics information just by specifying the reactants. The infrastructure program’s focus is on preconditioners (tools which improve convergence rates of certain mathematical methods) and on potential energy surface fitting schemes that reduce the number of electronic structure calculations necessary for accurate kinetics and dynamics. The infrastructure software has benefited from Integrated Software Infrastructure Centers (ISIC) support.

Principal Investigators: Albert F. Wagner (2001-03), Ron Shepard (2004-05), Argonne National Laboratory

The **Theoretical Chemical Dynamics Studies of Elementary Combustion Reactions** project involved modeling the dynamics of chemical reactions of large polyatomic molecules and radicals important in combustion research, using a combination of theories and methods. These computationally challenging studies will test the accuracy of statistical theories for predicting reaction rates, which essentially neglect dynamical effects. This research is expected help extend theoretical chemical dynamics to the treatment of complex chemical reactions of large polyatomic molecules.

Principal Investigator: Donald L. Thompson, University of Missouri

Biological and Environmental Research: Advanced Software for Studying Global Climate Change

In many fields of scientific research, scientists conduct experiments to test theories, then analyze the results, and use the information to continue refining their theories and/or experiments as they gain additional insight. In the study of global climate change, however, the “experiment” takes place day after day, year after year, century after century, with the entire planet and its atmosphere as the “laboratory.” Climate change literally affects each and every person on Earth and because any measures taken to address the issue will have far-reaching social, economic and political implications, it’s critical that we have confidence that the decisions being made are based on the best possible information.

Scientists have thought for more than 100 years that increasing atmospheric concentrations of carbon dioxide (CO₂) and other greenhouse-gases from human activity would cause the atmospheric layer closest to the Earth’s surface to warm by several degrees. Because of the complex feedbacks within the Earth system, precisely estimating the magnitude and rate of the warming, or understanding its effects on other aspects of climate, such as precipitation, is a daunting scientific challenge. Thanks to a wealth of new observational data and advances in computing technology, current climate models are able to reproduce the global average temperature trends observed over the twentieth century and provide evidence of the effect of human activity on today’s climate.

The Climate Change Research

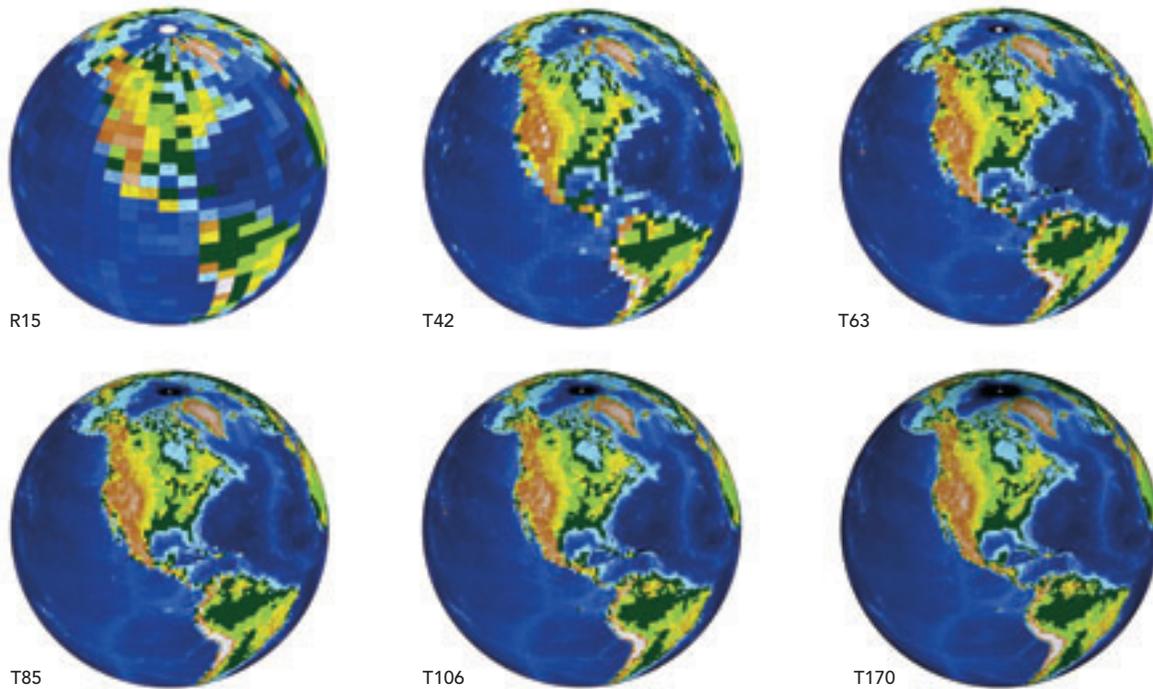


FIGURE 4. As supercomputers have become more powerful, climate researchers have developed codes with finer and finer resolution, allowing the models to incorporate more details which affect climate. The resulting climate models are more accurate. *Images by Gary Strand, NCAR*

Division in the DOE Office of Science was established "... to advance climate change science and improve climate change projections using state-of-the-science coupled climate models, on time scales of decades to centuries and space scales of regional to global."

Fortunately, the growing power and capabilities of high-performance computers are giving climate researchers more accurate tools for analyzing and predicting climate change. Climate, whether in our neighborhood, our region or continent, is determined by many factors, some local, others global. Geography, air temperature, wind patterns, ocean temperature and currents and sea ice are among the forces that shape our climate. As a result, comprehensive climate models which couple together simulations of different processes are among the most complex and sophisticated computer codes in existence. While they can simulate most of the conti-

ental-scale features of the observed climate, the models still cannot simulate, and therefore cannot predict, climate changes with the level of regional spatial accuracy desired for a complete understanding of the causes and effects of climate change.

As climate science advances and new knowledge is gained, there is a demand to incorporate even more factors into the models and to improve acknowledged shortcomings in existing models. These demands, which make the models more accurate, will unfortunately overwhelm even the most optimistic projections of computer power increases. So, a balance must be struck between the costs, benefits and tradeoffs required to allocate scarce computer and human resources to determine which improvements to include in the next generation of climate models.

Under SciDAC, teams of climate researchers worked to develop improved computational tools and

techniques for studying and modeling climate change, focusing on the numerical and computational aspects of climate modeling. In particular, SciDAC funding has enabled DOE researchers to participate with NSF and NASA researchers over an extended period in the design and implementation of new climate modeling capabilities. There are three major pieces, two of which pursued research in the academic arena that was longer term and more basic.

A National Consortium to Advance Global Climate Modeling

The major SciDAC effort in climate modeling was a multi-disciplinary project to accelerate development of the Community Climate System Model (CCSM), a computer model of the Earth's climate that combines component models for

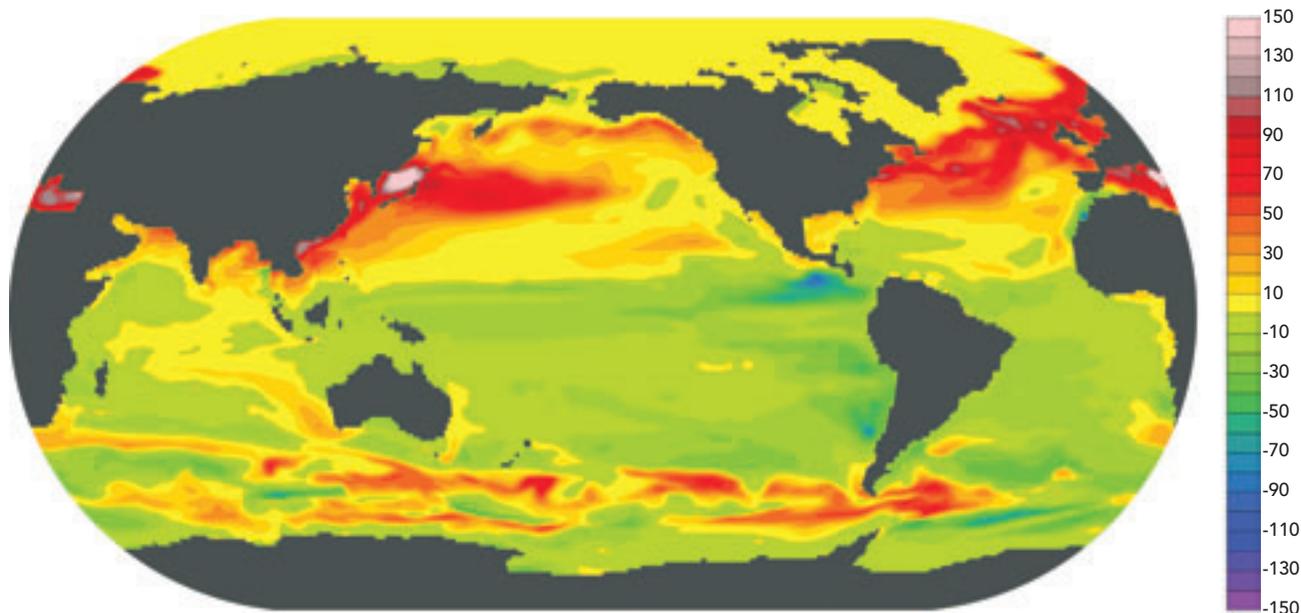


FIGURE 5. Using the Community Climate System Model, climate scientists ran a fully coupled simulation of the global carbon cycle to study how CO_2 is either taken up or released by the ocean. Understanding this transfer is important because a significant fraction of anthropogenic CO_2 emissions are currently absorbed by the ocean, thus slowing down the accumulation of CO_2 in the atmosphere. This image shows the average exchange (in tonnes of carbon per square kilometer per year) in the last month (December) of a nine-year simulation. Areas with positive values (shown ranging from yellow to pink) indicate where CO_2 is being taken up by the ocean, and negative values (shown in green ranging to purple) indicate where CO_2 is being released. When the CO_2 in the ocean is out of equilibrium with that in the atmosphere, CO_2 is exchanged between the two. Nonequilibrium can occur due to the temperature of the water, with the Northern Hemisphere (colder water in December) able to absorb more CO_2 than the relatively warm Southern Ocean. Biological activity can also alter the equilibrium when microscopic plants use CO_2 in the water for photosynthesis, which is then replaced by absorption from the atmosphere. The yellow and orange areas in the Southern Ocean demonstrate this phenomenon, showing uptake resulting from a seasonal phytoplankton bloom.

the atmosphere, ocean, land and sea ice. The CCSM is developed by researchers from NSF, DOE, NASA and NOAA laboratories, as well as universities. As one of the leading climate models in the United States, CCSM has a large user base of several hundred climate scientists. The CCSM community contributed a large number of climate change simulation results for the periodic Intergovernmental Panel on Climate Change (IPCC) climate assessment report. The **Collaborative Design and Development of the Community Climate System Model for Terascale Computers** project had two goals.

The first goal was to improve software design and engineering of the CCSM and its component models and improve performance

portability across the wide variety of computer architectures required for climate assessment simulations. The second goal was to accelerate the introduction of new numerical algorithms and new physical processes within CCSM models.

Under SciDAC, a consortium of six national laboratories and researchers from the National Center for Atmospheric Research (NCAR) and NOAA worked together on a variety of improvements to the CCSM. In the early years of the five-year project, attention was focused on improving the performance and portability of CCSM on the wide variety of vector and scalar computers available to the community.

The atmosphere and ocean models were improved with the introduction of new flexible data

decomposition schemes to enable tuning for each computational platform. The sea ice and land models were restructured to improve performance, particularly for vector computers, and new software was developed to improve the coupling of the four components into a fully coupled model.

These changes enabled the largest ensemble of simulations – 10,000 years worth – of any modeling group in the world for the recent IPCC assessment. These simulations were performed at relatively high resolution, generating more than 110 terabytes of climate model data, which were distributed internationally via the SciDAC-funded Earth System Grid. This important contribution to the international climate research effort was made pos-

sible by the critical software engineering work performed by the SciDAC team members on all the components of CCSM.

In addition to software expertise, the SciDAC CCSM consortium also contributed new model algorithms and new scientific capabilities. The consortium contributed to the introduction of a new finite-volume method for simulating atmospheric dynamics and developed new alternative schemes for ocean models as well.

Details of the consortium's software engineering of the CCSM were reported in articles featured in a special issue on climate modeling of the *International Journal of High Performance Computing and Applications* (Vol. 19, No. 3, 2005).

In the last years of the project, the focus of the SciDAC consortium was the development of new capabilities for simulating the carbon and sulfur cycles as part of the climate system. Until recently, most climate change scenarios specify a concentration of greenhouse gases and atmospheric aerosols. By adding the biological and chemical processes that govern the absorption and emission of greenhouse gases, better simulations of how the Earth system responds to human-caused emissions are possible (see Figure 1). A prototype carbon-climate-biogeochemistry model was assembled as a demonstration of the readiness to undertake coupled Earth system simulation at this dramatically new level of complexity.

The new model included a comprehensive atmospheric chemistry formulation as well as land and ocean ecosystem models. In this first step towards a comprehensive Earth system model, CO₂ fluxes were exchanged between components (see Figure 5) and the oceanic flux of dimethyl sulfide (DMS) was used in the atmospheric model to create sulfate aerosols. These aerosols then interacted and affected the physical climate system, the oceanic carbon cycle and terrestrial

ecosystems. The prototype model developed under SciDAC will form the basis for future work on a comprehensive Earth system model and help the climate community enter a new phase of climate change research.

Principal Investigators: John Drake, Oak Ridge National Laboratory, Phil Jones, Los Alamos National Laboratory

A Brand New Model

Under SciDAC, DOE's Office of Biological and Environmental Research pursued the novel concept of a five-year cooperative agreement with one or two universities to build the prototype climate model of the future. The concept was to develop

ocean components and alternative numerical methods.

The project, **A Geodesic Climate Model with Quasi-Lagrangian Vertical Coordinates**, was created to develop a new, comprehensive model of the Earth's climate system that includes model components for the atmosphere, ocean, sea-ice, and land surface, along with a model coupler to physically link the components. A multi-institutional collaboration of universities and government laboratories in an integrated program of climate science, applied mathematics and computational science, this project built upon capabilities in climate modeling, advanced mathematical

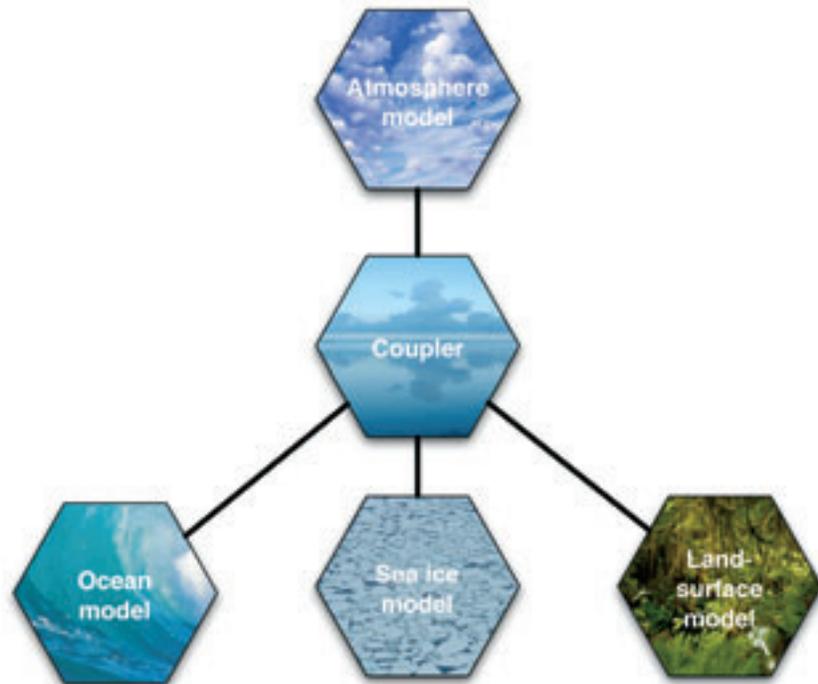


FIGURE 6. By coupling different climate modeling components, climate researchers can take advantage of advancements in the separate codes to improve overall accuracy.

a new climate model without the legacies of past approaches, which may be scientifically or computationally outdated. Only one qualified proposal was submitted, that from a group centered at Colorado State University, which is developing a model using a radically different grid design, alternative formulations for both the atmosphere and

research, and high-end computer architectures to provide useful projections of climate variability and change at regional to global scales. The project team developed components for modeling the atmosphere, oceans and sea ice, and a coupler component to links the physical model components. For example, the coupler computes the

latent heat flux exchanged between the ocean surface and atmosphere. This physical coupling enables the coupled system to evolve in a coherent manner. As the pieces come together the integrated system is being thoroughly tested.

Principal Investigator: David A. Randall, Colorado State University

University-Led Climate Modeling Projects

SciDAC supported two rounds of university grants to individual researchers to address the 5- to 15-year needs and opportunities to advance climate modeling. These grants supported research into new methodologies and numerical methods. It is this basic research that explores new ideas and concepts that tie climate science and computational science together as both fields advance.

A major factor in climate change between decades is the effect of winds on ocean circulation. The project on **Predictive Understanding of the Oceans: Wind-Driven Circulation on Interdecadal Time Scales** was aimed at developing and

applying advanced computational and statistical methods to help climate models better predict these effects. The computational aspect of the project was aimed at developing efficient multi-level methods to simulate these ocean flows and study their dependence on physically relevant parameters. The oceanographic and climate work consists in applying these methods to study the bifurcations in the wind-driven circulation and their relevance to the flows observed at present and those that might occur in a warmer climate. Both aspects of the work are crucial for the efficient treatment of large-scale, eddy-resolving numerical simulations of the oceans and an increased understanding of climate change.

Principal Investigators: Michael Ghil, UCLA; and Roger Temam, Indiana University

With the advent of more powerful supercomputers and modeling applications, climate models, resolution of the globe has become increasingly precise – to the extent that some models can focus on an area as small as 10 kilometers square. While this is useful for large-scale climate

conditions, understanding smaller phenomena such as tropical storms require 1-km resolution. **The Continuous Dynamic Grid Adaptation in a Global Atmospheric Model** is aimed at providing a capability to break down the larger grid structure to target areas of interest. While climate models are not expected to provide a uniform 1-km resolution within the next 15 years, this grid adaptation capability is a promising approach to providing such resolution for specific conditions.

Principal Investigators: J.M. Prusa and W.J. Gutowski, Iowa State University

While much discussion of climate change focuses on the global scale, there is also increasing demand for climate modeling on the regional or mesoscale covering areas from 50 to several hundred miles in size. The project to develop **Decadal Climate Studies with Enhanced Variable and Uniform Resolution GCMs Using Advanced Numerical Techniques** focused on using developed and evolving state-of-the-art general circulation models (GCMs) with enhanced variable and uniform resolution to run on parallel-processing terascale supercomputers. The major accomplishment of the project was completion of the international SGMIP-1 (Stretched-Grid Model Intercomparison Project, phase-1), which allows smaller regional models to be computed as part of global modeling systems. The “stretched grid” approach provides an efficient down-scaling to mesoscales over the area of interest and computes the interactions of this area with the larger model. This capability is expected to improve our understanding of climate effects on floods, droughts, and monsoons.

Collaboration with J. Côté of MSC/RPN and his group is a strong integral part of the joint effort. The companion study with the members of another SciDAC group, F. Baer and J. Tribbia, is devoted to developing a stretched-grid GCM using the advanced spectral-element tech-

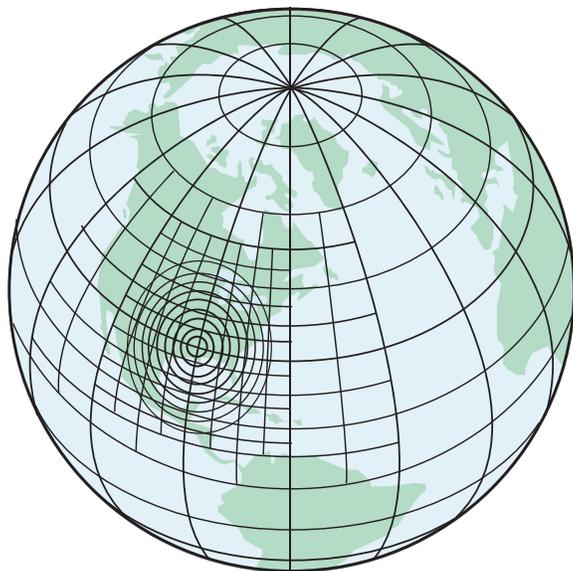


FIGURE 7. Adapted grid with block-data structure.

nique with variable resolution, and the NCAR CAM physics.

Principal Investigator: Michael S. Fox-Rabinovitz, University of Maryland

The project for **Development of an Atmospheric Climate Model with Self-Adapting Grid and Physics** was funded to develop adaptive grid techniques for future climate model and weather predictions. This approach will lead to new insights into small-scale and large-scale flow interactions that cannot be modeled by current uniform-grid simulations. This project will result in a climate model that self-adjusts the horizontal grid resolution and the complexity of the physics module to the atmospheric flow conditions. Using an approach called adaptive mesh refinement to model atmospheric motion will improve horizontal resolution in a limited region without requiring a fine grid resolution throughout the entire model domain. Therefore, the model domain to be resolved with higher resolution is kept at a minimum, greatly reducing computer memory and speed requirements. Several tests show that these modeling procedures are stable and accurate.

Principal Investigator: Joyce E. Penner, University of Michigan

The credibility of ocean models for climate research depends on their ability to simulate the observed state and natural variations of the oceans as indicated by actual measurements. Existing models, which can be used for simulating the long time scales of climate change of the order of centuries, still do not provide a very satisfactory treatment of key climatic processes such as water mass formation in the subpolar oceans. Ocean models based on Cartesian coordinates have been well tested and their drawbacks are well known. Models based on a moving vertical coordinate have the potential to provide a much more accurate simulation, but are not “mature” enough at present to gain widespread accept-

ance in the climate modeling community. The project to develop **A Vertical Structure Module for Isopycnal Ocean Circulation Models** aimed at providing a module for representing key processes in such a model and organizing the vertical structure. The module can then be inserted in the ‘dynamic core’ of existing models and used by the modeling community.

Principal Investigator: Kirk Bryan, Princeton University

The advent of terascale supercomputers has advanced climate modeling by allowing the use of higher-resolution atmospheric and ocean models. However, more dramatic improvements are likely through development of improved descriptions of physical processes treated by the models, especially those that can be observationally characterized in much finer details than are conventionally included in the climate models. The project for **Improving the Processes of Land-Atmosphere Interaction in CCSM 2.0 at High Resolution** has contributed to this goal by advancing the treatment of land with much improved details in the climatically most important processes, using primarily the Community Land Model (CLM).

Principal Investigator: Robert E. Dickinson, Georgia Institute of Technology

As part of the effort to develop climate models for making reliable climate predictions, one need is for an efficient and accurate method for producing regional climate predictions and the development of computing methodology which uses the latest in computing hardware (massively parallel processing or MPP) most efficiently and economically, to produce the best prediction results with minimal expenditure of resources. To meet this goal, the **Multi-Resolution Climate Modeling** project developed a Spectral Element Atmospheric Model (SEAM), a fairly recent concept using spectral ele-

ments (a method of approximating mathematical solutions). The Earth’s spherical domain is tiled with spectral elements that can be arbitrarily sized to meet local scaling requirements, allowing the model to create predictions over a range of scales on the entire global domain without user involvement in the computational process. The method also takes optimum advantage of state-of-the-art MPP by minimizing communication among elements and thereby amongst processors. This procedure has yielded dramatic speedup, making the production of multiple realizations more feasible. In addition to serving as a research and training tool, the model is expected to provide more accurate climate predictions.

Principal Investigator: Ferdinand Baer, University of Maryland

The **Decadal Variability in the Coupled Ocean-Atmosphere System** project researched the slowly changing circulations in the oceans and their influence on long-term global climate variability. There were two main themes in the project. The first concerned the decadal variability of upper ocean circulation and its role in the long period fluctuations of the coupled ocean/atmosphere system. The second was the role of mesoscale eddies – the oceanic flows on scales of 10 to 100 km – in the large-scale heat and energy budget of the ocean. The model was developed and tested on a smaller computer and will be adapted for use on larger massively parallel supercomputers.

Principal Investigator: Paola Cessi, Scripps Institution of Oceanography – University of California, San Diego

A project called **Towards the Prediction of Decadal to Multi-Century Processes in a High-Throughput Climate System Model** pursued interdisciplinary research to simulate decadal to multi-century global variability and change in an earth system model that couples climate to the terrestri-

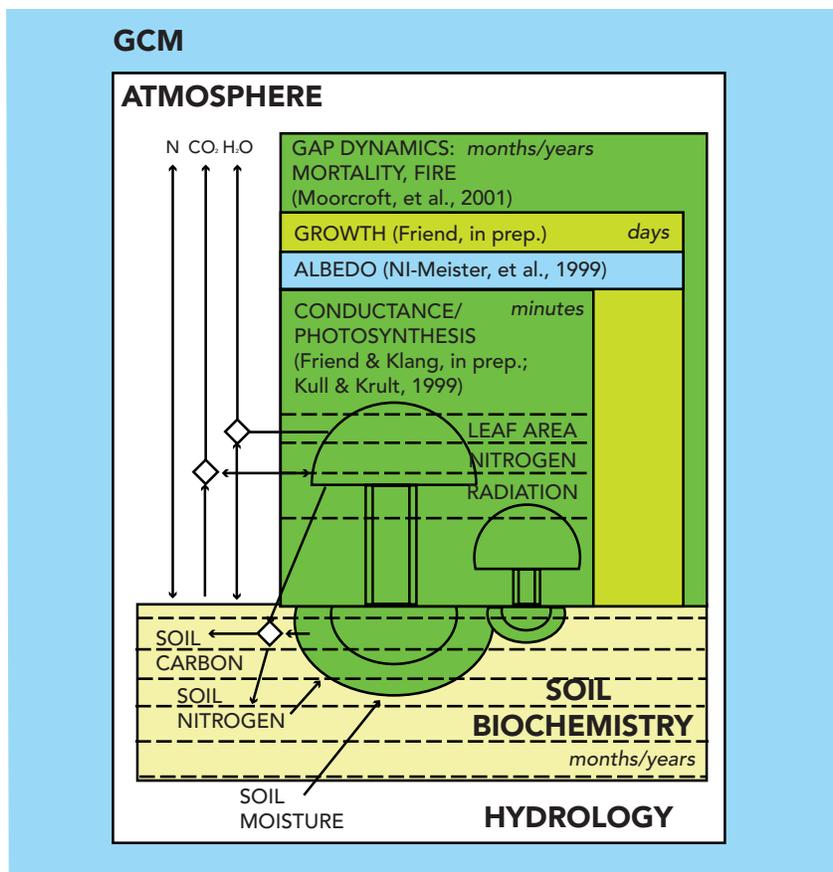


FIGURE 8. Diagram of the dynamic vegetation embedded within the atmospheric and land surface components of a GCM.

al ecosystem. The primary tool will be a high-throughput climate-ecosystem model called FOAM that enables rapid turnaround on long climate runs. FOAM is used to continue the study of the mechanisms of decadal climate variability, the dynamics of global warming, and the interaction of climate and the land biosphere. The project also linked climate scientists with ecosystem modelers in building a fully coupled ocean-atmosphere-land biosphere model, which enabled the study of the interaction of land vegetation changes and climate changes for past, present, and future scenarios.

Principal Investigator: Zhengyu Liu,
University of Wisconsin-Madison

Vegetation is an important component of the global climate system through its control of the fluxes of

energy, water, carbon dioxide and nitrogen over land surfaces. The aim of the project for **Modeling Dynamic Vegetation for Decadal to Century Climate Change Studies** is to develop, evaluate, utilize and make available a model of vegetation/soil dynamics to improve the ability of general circulation models (GCMs) to make predictions of future climate change. The model is being developed by combining treatments of carbon and nitrogen fluxes, and vegetation community dynamics, as a standalone module within the NASA Goddard Institute for Space Studies GCM. This model will be a tool for answering questions about past climate and vegetation distributions, as well as for predicting global changes due to rising atmospheric CO₂ in the coming decades and century. In one sce-

nario, in which the concentration of CO₂ in the atmosphere was doubled, the model showed that vegetation increased the uptake of CO₂ by 48 percent and that surface temperatures in some regions increased by up to 2° C due to stomatal closure.

Principal Investigator: Nancy Y. Kiang,
Columbia University

Gaining a better understanding of the relationship between climate change and the transport of water vapor and chemicals in the atmosphere was the aim of the **Modeling and Analysis of the Earth's Hydrologic Cycle** project. This research addresses fundamental issues underlying the understanding and modeling of hydrologic processes, including surface evaporation, long-range transport of water, and the release of latent heating through evaporation. These processes play a central role in climate. The goals of the project are to advance climate change modeling by developing a hybrid isentropic model (in which the flow variables change gradually) for global and regional climate simulations, to advance the understanding of physical processes involving water substances and the transport of trace constituents, and to examine the limits of global and regional climate predictability.

Principal Investigator: Donald R. Johnson,
University of Wisconsin – Madison

Fueling the Future: Plasma Physics and Fusion Energy

Plasmas, or very hot ionized gases, make up more than 99 percent of the visible universe. In fact, the earth's sun and the stars we see at night are masses of plasma burning at millions of degrees Celsius. Inside these stars, the incredibly high temperatures and pressures result in atoms fusing together and

releasing more energy. For much of the 20th century, scientists have investigated whether similar conditions could be created and used on Earth as source of energy.

Over the last 30 years, computing has evolved as a powerful tool in scientific research, including fusion. High performance computers are now routinely being used to advance our understanding of fusion energy, with the goal of harnessing the same power source of the sun in specially built reactors to provide clean and almost unlimited energy. But just as fusion energy is filled with huge potential, high performance computing also presents daunting scientific challenges.

Most of DOE's current research, both in experimental facilities and in computational science, focuses on magnetic fusion research. One approach centers on reactors with doughnut-shaped chambers called tokamaks which would be used to heat ionized gas to about 100 million degrees centigrade, then use powerful magnetic fields to confine and compress the plasma until hydrogen atoms fuse together and release helium and energy. This sustained fusion reaction is known as a "burning plasma." As envisioned, such reactors would generate more energy than they consume. Sustained fusion power generation requires that we understand the detailed physics of the fluid of ions and electrons, or plasma, in a fusion reactor.

With recent advances in super-computer hardware and numerical algorithm efficiency, large-scale computational modeling can play an important role in the design and analysis of fusion devices. Among the nations engaged in developing magnetic confinement fusion, the U.S. remains the world leader in plasma simulation.

Computational techniques for simulating the growth and saturation of turbulent instabilities in this plasma are critical for developing this understanding of how to control

the plasma and hence to develop a successful fusion reactor. Achieving such a goal is a long-term objective and many steps are needed before efficient fusion reactors can be designed and built. A number of smaller experimental reactors have been built in the U.S. and around the world. These facilities are expensive to build and operate, but give scientists valuable information for future designs.

Computational physics research has also helped give fusion scientists important insights into many aspects of plasma confinement, reactor fueling and the effects of turbulence on plasmas. In fact, turbulence has emerged as one of the biggest challenges in magnetic fusion – as the plasmas are heated and confined, turbulence occurs and introduces instabilities, which disrupt the confinement. If the plasmas cannot be confined, the plasmas come into contact with the reactor walls, losing temperature and making fusion impossible.

Under the SciDAC program, computational scientists, applied mathematicians and fusion researchers worked in teams to develop new tools and techniques for advancing fusion research, then using these methods to run more detailed simulations. The result is ever increasing knowledge and understanding about the forces at work within a fusion reactor. This information is being directly applied to the design and construction of ITER, a multinational facility to be built in France.

Since construction costs for ITER have been estimated at \$12 billion and once completed, operating costs could be up to \$1 million per day, it's critical that the newest, most detailed research findings be applied before the system is completed. Not only will this help minimize the number of costly modifications, but it will also mean a greater likelihood of success.

The results can also be applied on a broader scale to increase our

understanding of the many complex physical phenomena in the universe, which we are only starting to understand. Being able to capture and reproduce the phenomenon of fusion on Earth would solve the world's energy problems in an economically and environmentally sustainable manner.

Like other fields, fusion science has developed subfields such as plasma microturbulence theory, magnetohydrodynamics (MHD) and transport theory to study the phenomena which occur, and computational science has greatly advanced research in these areas. Under SciDAC, multi-institution projects were created to advance understanding in specific areas to help advance fusion research around the world.

Center for Extended Magnetohydrodynamic Modeling

The **Center for Extended Magnetohydrodynamic Modeling** project was established to enable a realistic assessment of the mechanisms leading to disruptive and other stability limits in the present and next generation of fusion devices. Rather than starting from scratch, the project built on the work of fusion research teams which developed the NIMROD and M3D codes for magnetohydrodynamic (MHD) modeling. This discipline studies the dynamics of electrically conducting fluids such as plasmas. The goal was to develop these codes to enable a realistic assessment of the mechanisms leading to disruptive and other stability limits in the present and next generation of fusion devices. The main work involves extending and improving the realism of the leading 3D nonlinear magneto-fluid based models of hot, magnetized fusion plasmas, increasing their efficiency, and using this improved capability

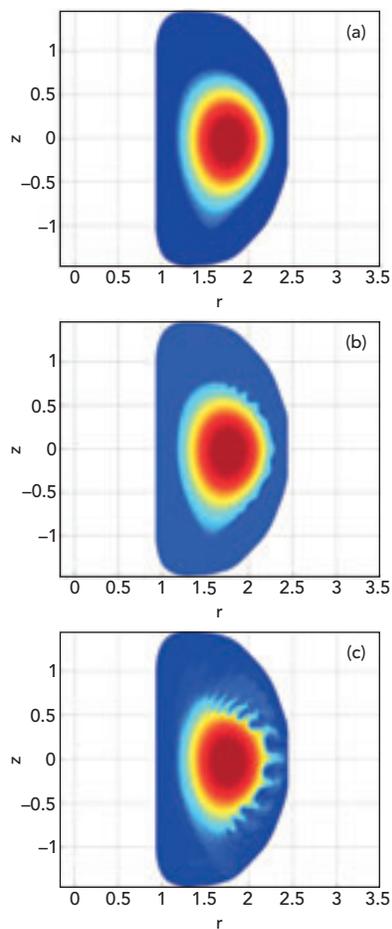


FIGURE 9. Evolution of the temperature during the nonlinear evolution of an edge localized mode in shot 113207.

to pioneer new terascale simulations of unprecedented realism and resolution. As a result, scientists gained new insights into low frequency, long-wavelength nonlinear dynamics in hot magnetized plasmas, some of the most critical and complex phenomena in plasma and fusion science. The underlying models are validated through comparisons with experimental results and other fusion codes.

An example of an important fusion problem that is being addressed by the center is the onset and nonlinear evolution of “edge localized modes” (ELMs) and their effect on plasma confinement and the reactor walls. By using the two MHD codes and data from the DIII-D fusion device operated by General Atomics in California, the center

was able to achieve a series of scientific milestones in 2005 and 2006. These modes shed thermal energy from the edge of the confinement region and, in their most virulent form, could overheat the walls near the plasma, and may also affect the core plasma. Using the codes advanced under this project, the center team was able to create more extensive simulations of the effects of ELMs than previously possible. The evolution of the temperature during the nonlinear evolution of an ELM in shot 113207 is shown in Figure 9, illustrating the formation of finger-like structures near the plasma edge with increasingly fine structure as the calculation progresses.

The project also made a number of modifications to the M3D code to improve the accurate representation of a number of conditions in tokamaks.

Principal Investigator: Steve Jardin,
Princeton Plasma Physics Laboratory

The Plasma Microturbulence Project

A key goal of magnetic fusion programs worldwide is the construction and operation of a burning plasma experiment. The performance of such an experiment is determined by the rate at which energy is transported out of the hot core (where fusion reactions take place) to the colder edge plasma (which is in contact with material surfaces). The dominant mechanism for this transport of thermal energy is plasma microturbulence.

The development of terascale supercomputers and of efficient simulation algorithms provides a new means of studying plasma microturbulence – direct numerical simulation. Direct numerical simulation complements analytic theory by extending its reach beyond simplified limits. Simulation complements experiment because non-perturbative diagnostics measuring quantities of immediate

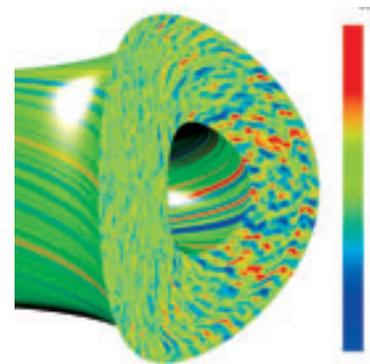


FIGURE 10. The (turbulent) electrostatic potential from a GYRO simulation of plasma microturbulence in the DIII-D tokamak.

theoretical interest are easily implemented in simulations, while similar measurements in the laboratory are difficult or prohibitively expensive.

The development of tools in this project will significantly advance the interpretation of experimental confinement data and will be used to test theoretical ideas about electrostatic and electromagnetic turbulence. By fulfilling the objective of enabling direct comparisons between theory and experiment, direct numerical simulation will lead to improvements in confinement theory and increased confidence in theoretical confinement predictions. **The Plasma Microturbulence Project (PMP)** is addressing this opportunity through a program of code development, code validation, and expansion of the user community.

The project has implemented a number of code improvements for modeling different geometries and multi-species plasma models to simulate both the turbulent electric and magnetic fields in a realistic equilibrium geometry (see Figure 10). The project’s global codes are able to simulate plasmas as large as those in present experiments, and the even larger plasmas foreseen in burning plasma experiments.

The project’s algorithms scale nearly linearly with processor number to ~1000 processors, which is important as fusion codes are significant users of supercomputing resources at DOE centers. To help

the fusion research community benefit from this investment in the GS2 and GYRO codes, the project has engaged researchers in helping to validate the codes against experimental results. Additionally, project members have conducted training workshops and presented talks on the codes at key fusion research meetings.

Principal Investigator: Bill Nevins,
Lawrence Livermore National Laboratory

Center for Gyrokinetic Particle Simulations of Turbulent Transport in Burning Plasmas

The **Center for Gyrokinetic Particle Simulations of Turbulent Transport in Burning Plasmas** consortium was formed in 2004 to develop codes to simulate turbulent transport of particles and energy, and improve the confinement of burning plasmas in fusion reactors. In particular, the project is aimed at developing the capabilities for simulating burning plasma experiments at the scale of ITER, the international thermonuclear experimental reactor which is expected to be the first fusion reactor capable of sustaining a burning plasma when the machine goes on line in about 10 years.

Although fusion scientists have been creating simulation codes to study fusion for more than 30 years, ITER presents two unique challenges, according to Wei-li Lee, head of the center.

First, the size of the reactor and

the necessary simulations are large. The ITER will have a major radius of 6.2 meters and a minor radius of two meters, and a correspondingly large confined plasma must be simulated. Second, the temperature inside ITER will be higher than any other fusion reactor. This higher temperature means that new types of physics will be encountered. As a result, new numerical simulation models must be developed.

The new codes must also be created to perform as efficiently as possible so that the larger-scale simulations can be run on the nation's most powerful supercomputers.

The major challenge for the project is to use simulations to better understand and minimize the problem of turbulence in the reactor. At the core of the reactor, the temperatures are at their highest. At the outside edges, the temperatures are lower. As with weather, when there are two regions with different temperatures, the area between is subject to turbulence. This turbulence provides a means for the charged particles in the plasma to move toward the outer edges of the reactor rather than fusing with other particles. If enough particles (and their energy) come into contact with the reactor wall, the particles lose temperature and the fusion reaction cannot be sustained.

Scientists understand that the difference in temperatures as well as densities is what causes the turbulence. But what is still not fully understood is the rate at which par-

ticles are transported through the plasma by the turbulence. Experiments show that the particles are transported quite differently than theory suggests. One of the objectives of the center's simulations is to bridge this gap between experiment and theory.

In the simulations, particles will fly around in the plasma according to Newton's laws of motion, although in a much smaller number than will actually be present in the ITER plasma. Algorithms will be developed to follow the path of each simulated particle as it interacts with other particles and is affected by turbulence transport. The results of the simulations will then be compared with experimental results.

To model the particles, the group is using the gyrokinetic particle-in-cell (PIC) method, first developed in the early 1980s and widely adopted in the 1990s. Gyrokinetic refers to the motion of the particles in a strong magnetic field. In such a field, particles spiral along, with electrons spinning in one direction and ions in the opposite directions. By simplifying the spiraling motion as a charged ring, researchers are able to increase the time steps by several orders of magnitude without sacrificing scientific validity.

The project team also collaborated with Terascale Optimal PDE Simulations Integrated Software Infrastructure Center (TOPS ISICs) established under SciDAC. By integrating other software libraries for solving the equations describing the interaction between the particles, the project team was able to solve a number of physics problems and integrate more realistic physics modules in their simulations. The result is that the simulations are expected to be able to model ITER-scale plasmas by solving a system of equations with millions of unknowns by following billions of particles in the computer.

Principal Investigator: W.W. Lee, Princeton Plasma Physics Laboratory

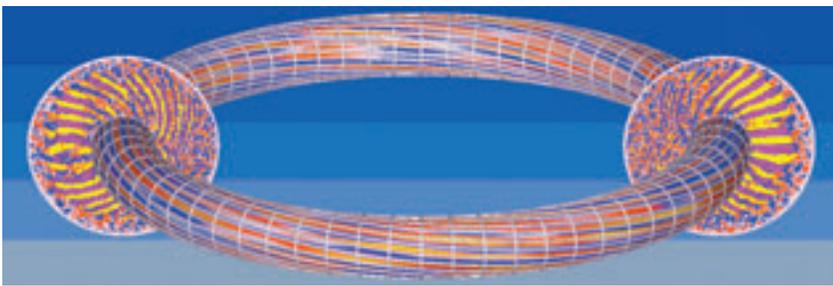


FIGURE 11. Potential contours of microturbulence for a magnetically confined plasma. The finger-like perturbations (streamers) stretch along the weak field side of the poloidal plane as they follow the magnetic field lines around the torus.

Understanding Magnetic Reconnection in Plasmas

The **Center for Magnetic Reconnection Studies (CMRS)** is a multi-university consortium dedicated to physical problems involving magnetic reconnection in fusion, space, and astrophysical plasmas. Understanding magnetic reconnection is one of the principal challenges in plasma physics. Reconnection is a process by which magnetic fields reconfigure themselves, releasing energy that can be converted to particle energies.

The goal of the Magnetic Reconnection project was to produce a unique high performance code to study magnetic reconnection in astrophysical plasmas, in smaller-scale laboratory experiments and in fusion devices. The principal computational product of the CMRS – the **Magnetic Reconnection Code (MRC)** – is a state-of-the-art, large-scale MHD code for carrying out magnetic reconnection research with accuracy and completeness in 2D and 3D. The MRC is massively parallel and modular, has the flexibility to change algorithms when necessary, and uses adaptive mesh refinement (AMR) (Figure 12). The

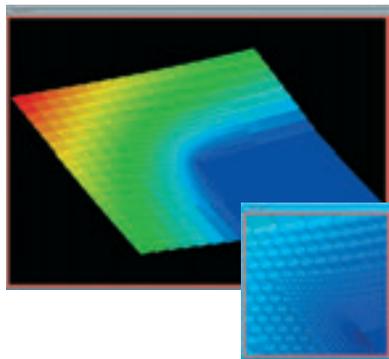


FIGURE 12. AMR allows the MRC code to place refined numerical grids automatically where the fine spatial structures require them, as shown in the figure where sharp spatial gradients at the center are resolved by placing refined grids (see inset for zoom). The MRC code has the capability to refine not only in space but also in time, which is more efficient since one does not need to take unnecessarily small time steps in regions where the fields are smooth.

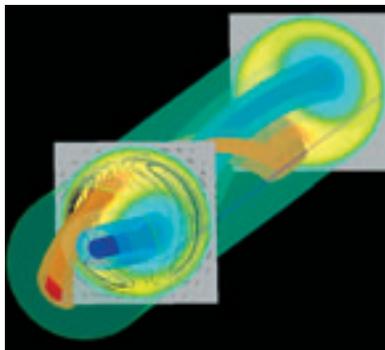


FIGURE 13. A simulation of the so-called kink-tearing instability in tokamaks that can cause internal current disruption and core temperature collapse. The two intertwining helical flux tubes are formed as a result of reconnection, and their 2D projections show magnetic island structure.

project team has applied the MRC to a wide spectrum of physical problems: internal disruption (Figure 13) and error-field induced islands in tokamaks, storms in the magnetosphere, solar/stellar flares, and vortex singularity formation in fluids.

Principal Investigator: Amitava Bhattacharjee, University of New Hampshire

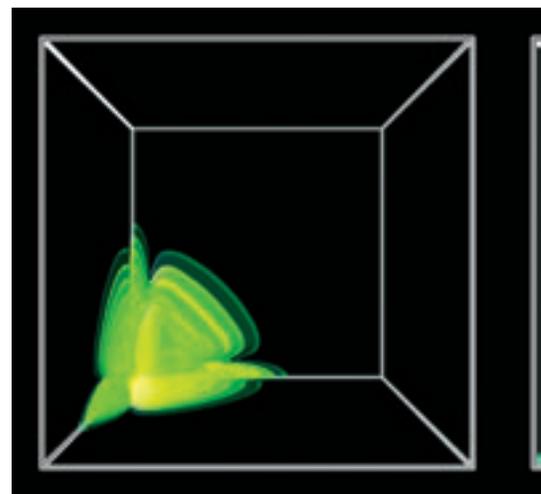
Terascale Computational Atomic Physics for the Edge Region in Controlled Fusion Plasmas

Atomic physics plays a central role in many of the high temperature and high density plasmas found in magnetic and inertial confinement fusion experiments, which are crucial to our national energy and defense interests, as well as in technological plasmas important to the U.S. economic base. In turn, the development of the necessary atomic physics knowledge depends on advances in both experimental and computational approaches. The **Terascale Computational Atomic Physics for the Edge Region in Controlled Fusion Plasmas** project was established to develop a new generation of scientific simulation codes, through a broad ranging collaboration of atomic physicists and

computer scientists, which will take full advantage of national terascale computational facilities to address the present and future needs for atomic-scale information for fusion and other plasma environments.

Especially in regard to the needs of fusion, atomic and molecular collisions significantly influence the transport and energy balance in regions crucial for the next step of magnetic confinement fusion development, the divertor and edge of tokamaks. For example, the microscopic modeling of turbulence and transport in magnetic fusion edge plasmas relies heavily on an accurate knowledge of the underlying atomic processes, such as elastic scattering, vibrational energy transfer, mutual neutralization, and dissociative recombination. The transport and atomic conversion of radiation is also at the heart of inertial fusion experiments, where a vast array of electron and photon collision processes, as well as of radiative and electron cascade relaxations, is required to model, diagnose, guide, and understand experiments. In addition, most magnetic fusion diagnostics involve interpretation of observations of the conversion of

FIGURE 14. Accurate modeling of plasmas requires accurate knowledge of atomic processes, such as scattering. This series is from a simulation capturing all scattering processes for the scattering of a wavepacket from a helium atom.



fast electron energy to atomic ion light emission, such as in the well known charge exchange recombination spectroscopy. Development of new scientific simulation codes to address these needs will also benefit research involving a huge range of technical, astrophysical, and atmospheric plasmas that also depend on accurate and large scale databases of atomic processes.

Many of the project's atomic collision codes are being implemented on supercomputers at DOE computing centers in California and Tennessee. Results of these code developments will be disseminated openly to the relevant plasma science and atomic physics communities, and will enable new regimes of computation by taking advantage of terascale and successor facilities. All of the atomic collision codes have important applications in a variety of research areas outside controlled fusion plasma science.

Principal Investigator: Michael Pindzola,
Auburn University

Numerical Computation of Wave-Plasma Interactions in Multi-Dimensional Systems

In order to achieve the extremely high temperatures – six times hotter than the core of the sun – needed to drive fusion reactions in a toka-

mak, scientists are studying different approaches to heating. One method involves radio waves, similar to the microwaves used to heat food. As the waves propagate through the plasma, they interact with the ions spinning in the plasma. When the ions' gyro-frequency resonates at the same frequency as the wave, the ions spin up to a very high energy, further heating the plasma.

The goal of **Numerical Computation of Wave-Plasma Interactions in Multi-Dimensional Systems** is to create two-dimensional and three-dimensional simulations of the interactions between the plasma and the waves to make accurate predictions of the processes. The results of these simulations will be

used to design and build fusion reactors, including ITER. During the first three years of the project, the main focus was on wave interaction and conversion. What researchers discovered was that under certain conditions, the relatively long waves propagated from an antenna inside the tokamak could be converted to much shorter waves, change characteristics and become much shorter in length. This process is known as "mode conversion." One form of the shorter wave, known as an "ion cyclotron wave," interacts with both the ions and the electrons in the plasma. The project is using two plasma modeling codes – AORSA and TORIC – to study the effects of the

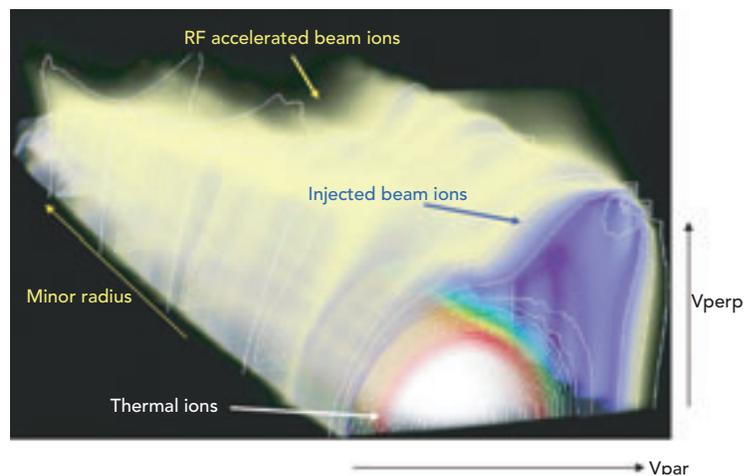
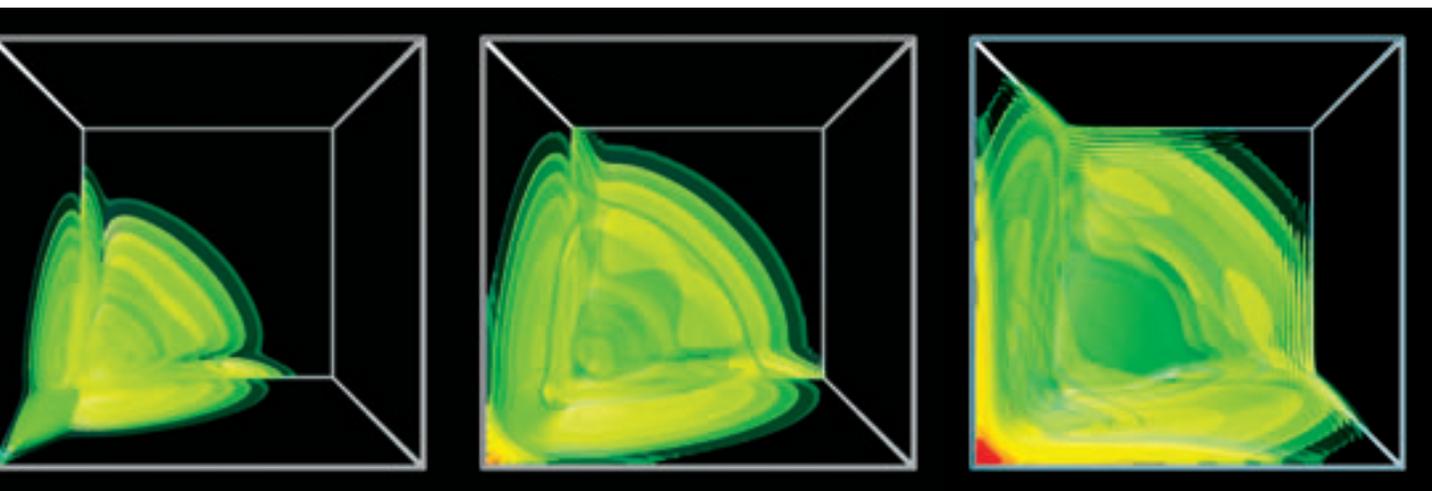


FIGURE 15. Under SciDAC, fusion energy researchers were allocated time on the Cray XT-3 massively parallel supercomputer "Jaguar" at Oak Ridge National Laboratory. By using thousands of processors, the team was able for the first time to run this calculation of how non-thermal ions are distributed in a heated fusion plasma.



ion cyclotron wave. One effect appears to be that the wave drives flows in the plasma which help break up turbulence, thereby improving confinement of the plasma.

In summer 2005, the AORSA global-wave solver was ported to the new Cray XT3 supercomputer at Oak Ridge National Laboratory. The code demonstrated excellent scaling to thousands of processors. Preliminary calculations using 4,096 processors have allowed the first simulations of mode conversion in ITER. Mode conversion from the fast wave to the ion cyclotron wave has been identified in ITER using mixtures of deuterium, tritium and helium-3 at 53 MHz.

A second focus of the project has been on how energies are distributed in the plasma. The simplest type of distribution is known as a Maxwellian distribution, which is used to calculate how energy is distributed in all gases, such as the air in a room. This distribution function was previously the only method in which fusion researchers could calculate wave propagation and absorption. But, when plasmas are heated to very high temperatures, a long tail of particles can form, which represents a non-Maxwellian distribution. Using another code called CQL3D, project team members were able to calculate the new energy distributions. The code was also installed on the Cray XT3 supercomputer along with the AORSA code, and both were programmed to automatically iterate with one another. So, once the new energy distributions were calculated, this data was fed back into the AORSA code to recalculate the heating of the plasma. Once the reheating was simulated, the CQL3D code then calculated the energy distribution. The iterations were repeated until a steady state was reached in the simulation.

The project results, which make key contributions to understanding processes in hot plasmas, were

achieved through the scaling of simulation codes and the availability of terascale computing systems.

Principal Investigator: Don Batchelor, Oak Ridge National Laboratory (2001-05), Paul Bonoli, Massachusetts Institute of Technology, (2006-present)

High Energy and Nuclear Physics: Accelerating Discovery from Subatomic Particles to Supernovae

The Office of Science supports a program of research into the fundamental nature of matter and energy through the offices of High Energy Physics and Nuclear Physics. In carrying out this mission, the Office of Science:

- builds and operates large, world-class charged-particle accelerator facilities for the nation and for the international scientific research community;
- builds detectors and instruments designed to answer fundamental questions about the nature of matter and energy; and
- carries out a program of scientific research based on experimental data, theoretical studies, and scientific simulation.

The scale of research ranges from the search for subatomic particles, which are one of the most basic building blocks of matter, to studying supernovae, massive exploding stars which also serve as an astrophysical laboratory in which unique conditions exist that are not achievable on Earth. Under the SciDAC program, projects were launched to advance research in the areas of accelerator science and simulation, quantum chromodynamics and supernovae science.

Advanced Computing for 21st Century Accelerator Science & Technology

Accelerators underpin many of the research efforts of the Office of Science and physics research around the world. From biology to medicine, from materials to metallurgy, from elementary particles to the cosmos, accelerators provide the microscopic information that forms the basis for scientific understanding and applications. Though tremendous progress has been made, our present theory of the physical world is not complete. By tapping the capabilities of the different types of accelerators, including colliders, spallation neutron sources and light sources, scientists are making advances in many areas of fundamental science.

Much of our knowledge of the fundamental nature of matter results from probing it with directed beams of particles such as electrons, protons, neutrons, heavy ions and photons. The resulting ability to “see” the building blocks of matter has had an immense impact on society and our standard of living. Over the last century, particle accelerators have changed the way we look at nature and the universe we live in and have become an integral part of the nation’s technical infrastructure. Today, particle accelerators are essential tools of modern science and technology.

For example, about 10,000 cancer patients are treated every day in the United States with electron beams from linear accelerators. Accelerators produce short-lived radioisotopes that are used in over 10 million diagnostic medical procedures and 100 million laboratory tests every year in the United States.

The SciDAC **Accelerator Science and Technology (AST)** modeling project was a national research and development effort aimed at establishing a comprehensive terascale simulation environ-

ment needed to solve the most challenging problems in 21st century accelerator science and technology. The AST project had three focus areas: computational beam dynamics, computational electromagnetics, and modeling advanced accelerator concepts. The newly developed tools are now being used by accelerator physicists and engineers across the country to solve the most challenging problems in accelerator design, analysis, and optimization.

An accelerator generates and collects billions of elementary particles, such as electrons, positrons or protons, into a limited space, then accelerates these particles in a beam toward a target – another beam of particles or devices for producing radiation or other particles. In the process of acceleration, the energy of every particle in the beam is increased tremendously. In order to advance elementary particle physics into regions beyond our present knowledge, accelerators with larger final beam energies approaching the tera electron volt (TeV) scale are required.

The Department of Energy operates some of the world's most powerful accelerators, including a three-kilometer-long linear accelerator at the Stanford Linear Accelerator Center (SLAC) in California, the Tandem Linac Accelerator System at Argonne National Laboratory (ANL) in Illinois, the Tevatron at Fermilab in Illinois, the Relativistic Heavy Ion Collider (RHIC) at Brookhaven National Laboratory (BNL) in New York, the Continuous Electron Beam Accelerator Facility in Virginia and the Holifield Radioactive Ion Beam Facility at Oak Ridge National Laboratory (ORNL) in Tennessee. DOE-operated light sources are the Advanced Light Source at Lawrence Berkeley National Laboratory in California, the Advanced Photon Source at ANL, the National Synchrotron Light Source at BNL and the Stanford Synchrotron Radiation Laboratory

at SLAC. Neutron sources include the Spallation Neutron Source at ORNL and the Los Alamos Neutron Science Center at Los Alamos National Laboratory in New Mexico. While accelerators can be of different configurations with different capabilities, most use radio frequencies (RF) to accelerate the particles. In order to experimentally pursue the quest for the grand unified theory ever more powerful accelerators are needed. Such massive facilities are costly to build and require specialized infrastructure.

The long-term future of experimental high-energy physics research using accelerators depends on the successful development of new acceleration methods which can accelerate particles to even higher energy levels per acceleration length (known as the accelerating gradient). In the past, this was achieved by building a longer accelerator, an approach which faces practical limits in space and cost.

Since building such accelerators is time-consuming and expensive, scientists are using computer simulations to increasing our understanding of the physics involved and

to help improve the design for more efficient acceleration. High-resolution, system-scale simulation utilizing terascale computers such as the IBM SP supercomputer at NERSC, has been made possible by SciDAC-supported code development efforts and collaborations with the SciDAC ISICs. This team-oriented approach to computing is beginning to make a qualitative difference in the R&D of major DOE accelerators, existing or planned.

Beam Dynamics: Maximizing Scientific Productivity

The AST project has had a major impact on computational beam dynamics and the design of particle accelerators. Thanks to SciDAC, accelerator design calculations that were once thought impossible are now carried out routinely. SciDAC accelerator modeling codes are being used to get the most science out of existing facilities, to produce optimal designs for future facilities, and to explore advanced accelerator concepts that may hold the key to qualitatively new ways of accelerating charged particle beams. Here are some high-

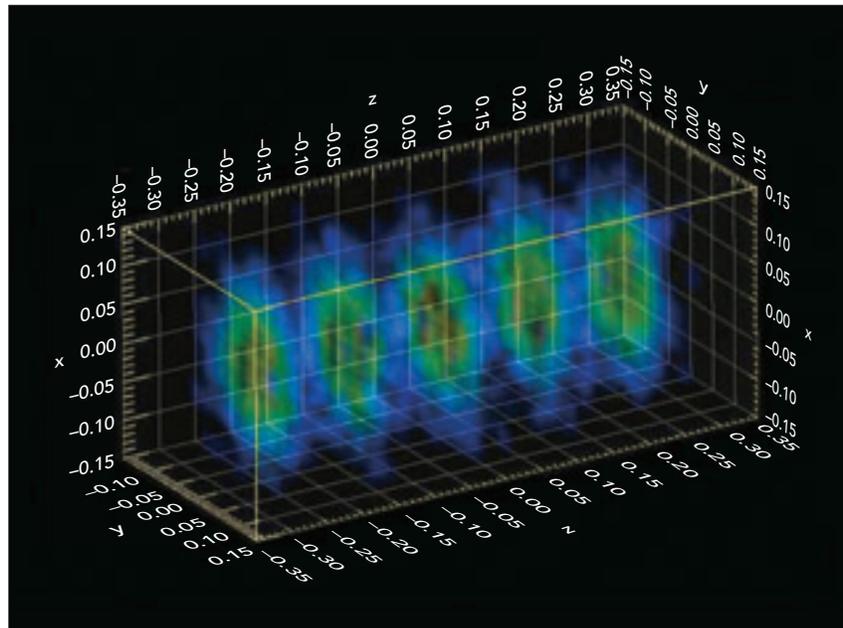


FIGURE 16. This simulation of the Fermi Booster, a synchrotron accelerator ring, was created using the Synergia software developed under SciDAC.

lights from the AST project Beam Dynamics focus area in regard to algorithm development, software development, and applications.

Particle simulation methods have been among the most successful and widely used methods in computational beam physics, plasma physics and astrophysics. Under the AST project a comprehensive, state-of-the-art set of parallel particle-in-cell (PIC) capabilities has been developed, including:

- **IMPACT:** An integrated suite of codes originally developed to model high intensity ion linacs, IMPACT's functionality has been greatly enhanced so that it is now able to model high brightness electron beam dynamics, ion beam dynamics and multi-species transport through a wide variety of systems.
- **BeamBeam3D:** A code for modeling beam-beam effects in colliders. This code contains multiple models and multiple collision geometries and has been used to model the Tevatron, Positron-Electron Project (PEP)-II, Relativistic Heavy Ion Collider (RHIC), and Large Hadron Collider (LHC) accelerators.
- **MaryLie/IMPACT:** A code that combines the high-order optics modeling capabilities of the MaryLie Lie algebraic beam transport code with the parallel PIC capabilities of IMPACT. It is used to model space-charge effects in large circular machines such as the ILC damping rings.
- **Synergia:** A parallel beam dynamics simulation framework, Synergia combines multiple functionality, such as the space-charge capabilities of IMPACT and the high-order optics capabilities of MXYZPLT, along with a "humane" user interface and standard problem description.

Crucial to the development of the AST project's codes has been the collaboration with the SciDAC

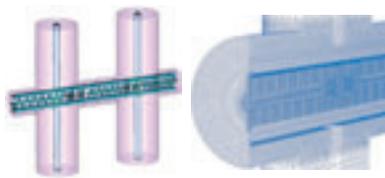


FIGURE 17. (Left) A model of RIA's hybrid RFQ and (Right) an enlarged portion of the corresponding Omega3P mesh.

ISICs and with other DOE-supported activities. In some cases, the use of advanced algorithms has led to applications running up to 100 times faster.

SciDAC AST beam dynamics codes have been applied to several important projects within the DOE Office of Science. Examples include:

- existing colliders, including the Tevatron, RHIC and PEP-II
- future colliders such as the LHC currently under construction
- proposed linear colliders such as the International Linear Collider (ILC)
- high intensity machines, including the Fermilab booster and the Spallation Neutron Source (SNS) ring under construction
- linacs for radioactive ion beams, such as the proposed Rare Isotope Accelerator (RIA)
- electron linacs for fourth-generation light sources, such as the Linac Coherent Light Source now under construction.

Electromagnetic Systems Simulation: Prototyping through Computation

The AST project has supported a well-integrated, multi-institutional, multi-disciplinary team to focus on the large-scale simulations necessary for the design and optimization of electromagnetic systems essential to accelerator facilities. As a result, an increasing number of challenging design and analysis problems are being solved through large-scale simulations, benefiting such projects as the PEP-II, Next Linear Collider (NLC),

and the proposed ILC and RIA.

Central to these efforts is a suite of 3D parallel electromagnetic design codes: Omega3P, Tau3P, Track3P, S3P and T3P. Collaborations with SciDAC ISICs led to a number of improvements in the codes which have not only increased the speed and accuracy of the codes up to tenfold, but have expanded their capabilities for solving more complicated problems of importance to present and future accelerators. The codes have been used to advance a number of DOE accelerator projects, including the ILC, PEP-II, NLC, RIA and LCSS.

For example, the AST electromagnetic codes were applied to the PEP-II facility, which now operates with twice the luminosity of the original design and is aiming for another twofold increase. However, beam heating in the interaction region (IR) could limit the accelerator from meeting that goal. The Tau3P code has been used to calculate the heat load in the present design. Higher currents and shorter bunches of particles will require modifications to the IR to reduce the increased heating. T3P will be able to model the entire IR geometry and to simulate the actual curved beam paths in any new IR design.

The proposed RIA, which is ranked third in the Office of Science's 20-year Science Facility Plan, requires design of a variety of radio frequency quadrupole (RFQ) cavities (Figure 17) for its low-frequency linacs. Due to lack of accurate predictions, tuners are designed to cover frequency deviations of about 1 percent. With Omega3P, frequency accuracy of 0.1 percent can be reached, significantly reducing the number of tuners and their tuning range. Parallel computing will play an important role in prototyping RIA's linacs and help reduce their costs.

Modeling Advanced Accelerators: Miniaturizing Accelerators from Kilometers to Meters

The long-term future of experimental high-energy physics research using accelerators is partly dependent on the successful development of novel ultra high-gradient acceleration methods. New acceleration techniques using lasers and plasmas have already been shown to exhibit gradients (or acceleration) and focusing forces more than 1000 times greater than conventional technology. The challenge is to control these high-gradient systems and then to string them together. Such technologies would not only enable a cost-effective path to the outermost reaches of the high-energy frontier, but could also lead to the development of ultra-compact accelerators. Such compact accelerators could benefit science, industry and medicine by shrinking large facilities to a much reduced size and allowing them to be built nearer research organizations, high-tech businesses and medical centers.

Under the AST Project, the Advanced Accelerator effort has emphasized developing a suite of fully parallel 3D electromagnetic PIC codes. The codes have been benchmarked against each other and their underlying algorithms, as well as against experiments. The resulting codes provide more realistic models, which are being applied to advanced accelerators as well as more mainstream problems in accelerator physics. Furthermore, the effort has included running these codes to plan and interpret experiments and to study the key physics that must be understood before a 100+ GeV collider based on plasma techniques can be designed and tested.

In some advanced accelerator concepts a drive beam, either an intense particle beam or laser pulse, is sent through a uniform plasma. The space charge or radiation pres-

sure creates a space-charge wake on which a trailing beam of particles can surf. To model such devices accurately usually requires following the trajectories of individual plasma particles. Therefore, the software tools developed fully or partially under this project – OSIRIS, VORPAL, OOPIC, QuickPIC and UPIC – rely on the PIC techniques.

Among the major accomplishments are:

The development of UPIC: Using highly optimized legacy code, a modern framework for writing all types of parallelized PIC codes including electrostatic, gyrokinetic, electromagnetic, and quasi-static has been developed. The UPIC Framework has obtained 30 percent of peak speed on a single processor and 80 percent efficiency on well over 1,000 processors. It is used across both the accelerator and fusion SciDAC projects.

The rapid construction of QuickPIC: The development of QuickPIC is a success story for the rapid construction of a new code using reusable parallel code via a SciDAC team approach. The basic equations and algorithms were developed from a deep understanding of the underlying physics involved in plasma and/or laser wakefield

acceleration while the code was constructed rapidly using the UPIC Framework. In some cases, QuickPIC can completely reproduce the results from previous algorithms with a factor of 50 to 500 savings in computer resources. It is being used to study beam-based science in regimes not accessible before.

Rapidly adding realism into the computer models: The large electromagnetic fields from the intense drive beams can field ionize a gas, forming a plasma. Early SciDAC research using two-dimensional codes revealed that this self-consistent formation of the plasma from the drive beam needs to be included in many cases. Via the SciDAC team approach, ionization models have been added and benchmarked against each other in the fully three-dimensional PIC codes, VORPAL and OSIRIS.

Extending the plasma codes to model the electron-cloud instability: The electron cloud instability is one of the major obstacles for obtaining the design luminosity in circular accelerators, storage rings and damping rings. A set of modules has been written for QuickPIC that models circular orbits under the influence of external focusing elements. This new software tool has already modeled 100,000 km of beam

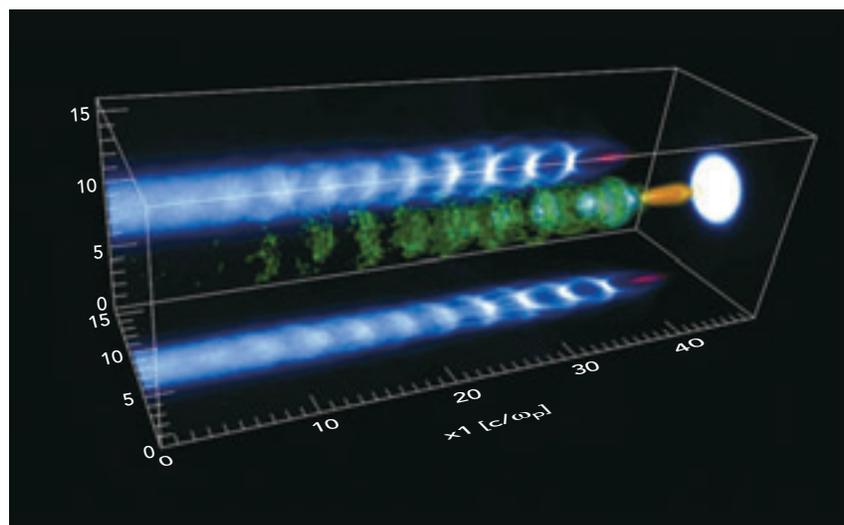


FIGURE 18. 3D afterburner simulation results.

propagation of the SPS machine at CERN. It is a major improvement over previously existing tools for modeling electron cloud interactions.

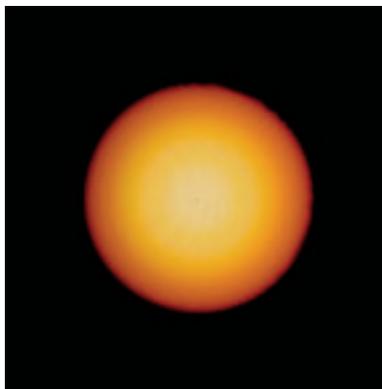
Applying the suite of codes to discover new accelerator science: The suite of PIC codes has been used to model particle beam accelerator experiments at the Stanford Linear Accelerator Center and the laser-plasma experiments at the L'OASIS lab at Lawrence Berkeley National Laboratory. They have also been used to study key physics issues related to the afterburner concept, in which the energy of an existing beam from an accelerator is doubled with gradients near 10 GeV/m. An example is shown in Figure 18 where the beam and plasma density is shown from a three-dimensional simulation for the afterburner, including field ionization.

Principal Investigators: Robert Ryne, Lawrence Berkeley National Laboratory, and Kwok Ko, Stanford Linear Accelerator Center

Supernovae Science: Getting the Inside Story on How Stars Explode

While supernovae are awesome in their own right – these exploding stars expire in flashes brighter than whole galaxies – they can also provide critical information about our universe. Because of their regular brightness, Type Ia supernovae are used as astronomical “standard candles” for cosmic distances and the expansion of the universe. Supernovae are also helping scientists gain a better understanding of the dark energy thought to make up 75 percent of the universe. And finally, supernovae are also the source of the heavy elements in our universe, as well as prolific producers of neutrinos, the most elusive particles so far discovered.

The best perspective from which to observe supernovae is from a telescope mounted on a



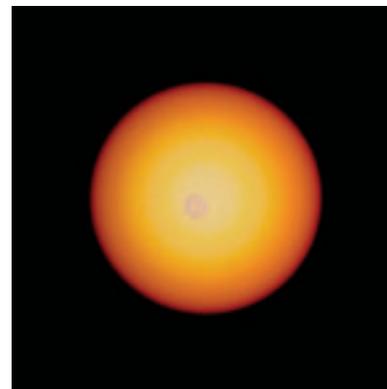
satellite. But earth-bound supercomputers are also proving to be powerful tools for studying the forces behind the most powerful explosions in the universe. To help understand the conditions which lead to supernovae explosions, the SciDAC program launched two projects – the Supernova Science Center and the Terascale Supernova Initiative.

The Supernova Science Center

The **Supernova Science Center (SNSC)** was established with the objective of using numerical simulations to gain a full understanding of how supernovae of all types explode and how the elements have been created in nature. These computational results are then compared with astronomical observations, including the abundance patterns of elements seen in our own sun.

The explosions both of massive stars as “core-collapse” supernovae and of white dwarfs as “thermonuclear” supernovae (also called Type Ia) pose problems in computational astrophysics that have challenged the community for decades. Core-collapse supernovae require at least a two- (and ultimately three-) dimensional treatment of multi-energy-group neutrinos coupled to multi-dimensional hydrodynamics. Type Ia explosions are simulation problems in turbulent (nuclear) combustion.

During its first five years, the SNSC focused on forging the interdisciplinary collaborations necessary to address cutting-edge problems in



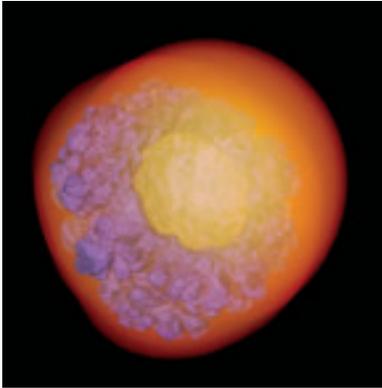
a field that couples astrophysics, particle physics, nuclear physics, turbulence theory, combustion theory and radiation transport. The project team also worked to develop and modify the necessary computer codes to incorporate the physics efficiently and do exploratory calculations.

Examples of these alliances are the chemical combustion groups at Lawrence Berkeley and Sandia National Laboratories; the NSF's Joint Institute for Nuclear Astrophysics (JINA); and radiation transport and nuclear physics experts at Los Alamos and Lawrence Livermore National Laboratories. With LBNL, the team applied codes previously optimized to study chemical combustion on large parallel computers to novel problems in nuclear combustion in white dwarf stars, carrying out for the first time fully resolved studies of the flame in both the “flamlet” and “distributed” regimes.

The project team worked with JINA to develop a standardized library of nuclear data for application to the study of nucleosynthesis in stars, supernovae and X-ray bursts on neutron stars.

The SNSC is also tapping the expertise of the SciDAC ISICs. For example, the TOPS ISIC is working with members of the team to produce more efficient solvers, with a goal of increasing the performance of the multi-dimensional codes by a factor of 10.

Other SNSC achievements include:



- Carrying out the first 3D, full-star simulation of core collapse and explosion. The project also carried out simulations in 2D with multi-group neutrino transport of the collapse, bounce, and early post-bounce evolution. Recently a new supernova mechanism was discovered, powered by neutron star vibrations.
- Developing and using 3D adaptive mesh relativistic hydrocodes to model the production of gamma-ray bursts in massive stars, which led to the prediction of a new kind of high energy transient – cosmological x-ray flashes, which were later discovered. (An image from this model appears on the cover of this report.)
- Carrying out the world's best simulations of Type I X-ray bursts on neutron stars. The nuclear physics of these explosions is a primary science goal of DOE's proposed Rare Isotope Accelerator.
- Calculating the ignition of the nuclear runaway in a white dwarf star becoming a Type Ia supernova, which showed that the ignition occurs off-center. Other research has shown that the resulting supernova depends critically upon whether it is ignited centrally or at a point off center. This issue of how the white dwarf ignites has become the most debated and interesting issue in Type Ia supernova modeling today.

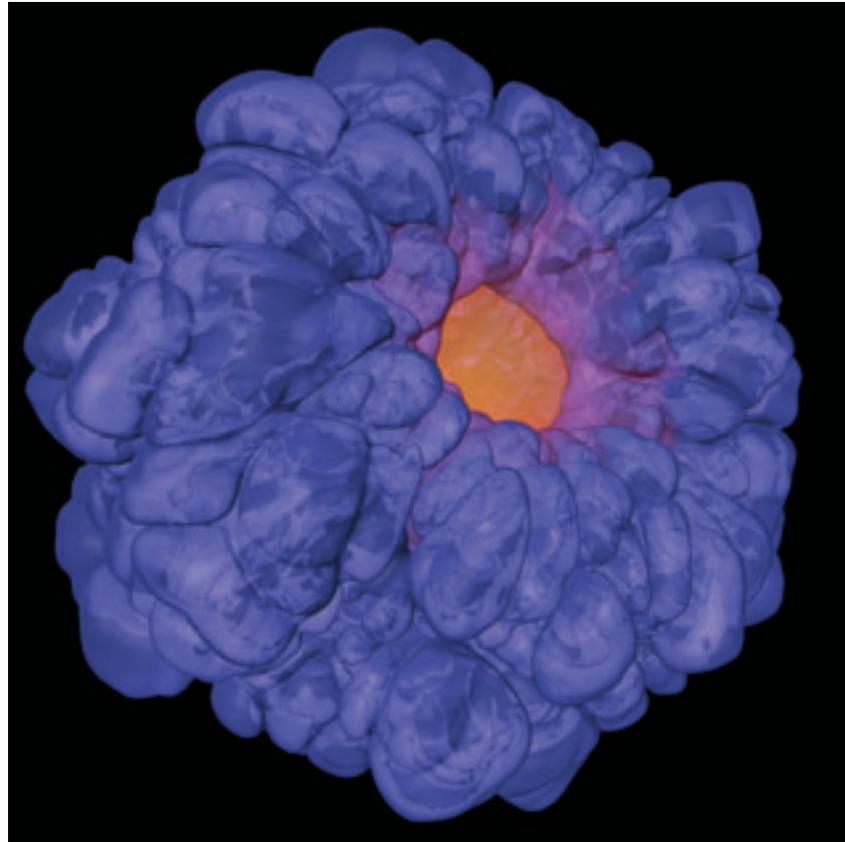


FIGURE 19. Snapshots of a thermonuclear supernova simulation with asymmetric ignition. After igniting a sub-sonic thermonuclear flame slightly off-center of the white dwarf progenitor star, it propagates towards the surface. The interaction of buoyancy and flame propagation changes the morphology of the initially spherical burning bubble (see 1st snapshot, where the flame corresponds to the blue isosurface and the white dwarf star is indicated by the volume rendering of the logarithm of the density) to a toroidal shape perturbed by instabilities (2nd snapshot, cf. Zingale et al., SciDAC 2005 proceedings). The flame interacts with turbulence and is accelerated, but due to the asymmetric ignition configuration, only a small part of the star is burned. Thus the white dwarf remains gravitationally bound and when the ash bubble reaches its surface it starts to sweep around it (3rd snapshot), as also noticed by the ASC Flash center at the University of Chicago. The question addressed in the present study was whether the collision of ash on the far side of the star's surface (4th snapshot) would compress fuel in this region strong enough to trigger a spontaneous detonation ("GCD scenario", Plewa et al., 2004). Three-dimensional simulations like those shown in the images disfavor this possibility.

The Terascale Supernova Initiative: Shedding New Light on Exploding Stars

Established to "shed new light on exploding stars," SciDAC's **Terascale Supernova Initiative (TSI)** is a multidisciplinary collaboration which will use modeling of integrated complex systems to search for the explosion mechanism of core-collapse supernovae – one of the most important and challenging problems in nuclear astrophysics. The project team, which includes scientists, computer scientists and mathematicians, will develop mod-

els for core collapse supernovae and enabling technologies in radiation transport, radiation hydrodynamics, nuclear structure, linear systems and eigenvalue solution and collaborative visualization. These calculations, when carried out on terascale machines, will provide important theoretical insights and support for the large experimental efforts in high energy and nuclear physics.

A key focus of the TSI project is understanding what causes the core of a large star to collapse. When the star runs out of fuel, pressures in the core build and fuse elements together, leading to accelerating combustion.

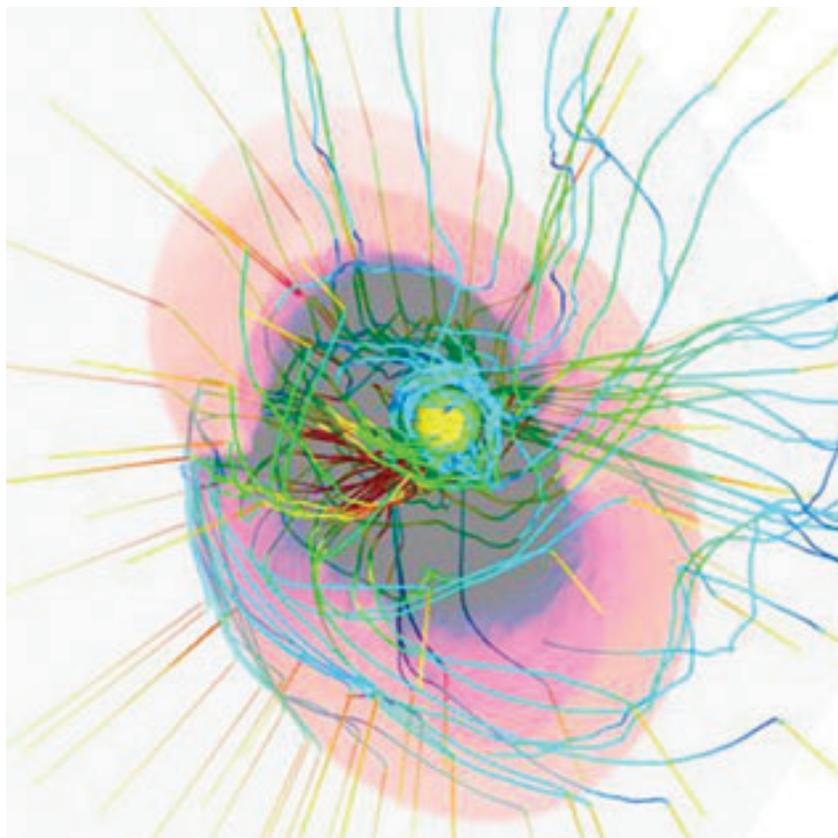


FIGURE 20. This image illustrates the stable rotational flow found in recent 3D simulations of the SASI, computed on the Cray X1 at NCCS. The streamlines show the inward flow of the core being deflected by the nascent supernova shock and ultimately wrapping around the inner core.

At the end, only iron is left at the core and fusion stops. As the iron core grows, gravitational forces eventually cause the core to implode, or collapse, leading to an explosion. This explosion creates both heavy elements and neutrinos. Once the core can no longer collapse any more, it “bounces” back and sends a shock wave through the external layers of the star. If the shock wave kept going, it would lead to the supernova blast. But, for reasons still not understood, the shock wave stalls before it can trigger the explosion in current simulations. TSI researchers are formulating new ways to examine critical events in the supernova core within the first second after the core bounces. One theory is that neutrino heating helps re-energize the stalled shock.

The TSI project created a series of 3D hydrodynamic simulations

showing the flow in a stellar explosion developing into a strong, stable, rotational flow (streamlines wrapped around the inner core). The flow deposits enough angular momentum on the inner core to produce a core spinning within a period of only a few milliseconds. This new ingredient of core-collapse supernovae, which can only be modeled in full 3D simulations, may have a dramatic impact on both the supernova mechanism itself and on the neutron star left behind by the supernova event. The strong rotational flow below the supernova shock leads to an asymmetric density distribution that can dramatically alter the propagation of neutrinos, which may in turn enhance the efficiency of a neutrino-driven shock. This rotational flow will also provide a significant source of energy for amplifying any pre-existing magnetic fields through the magneto-

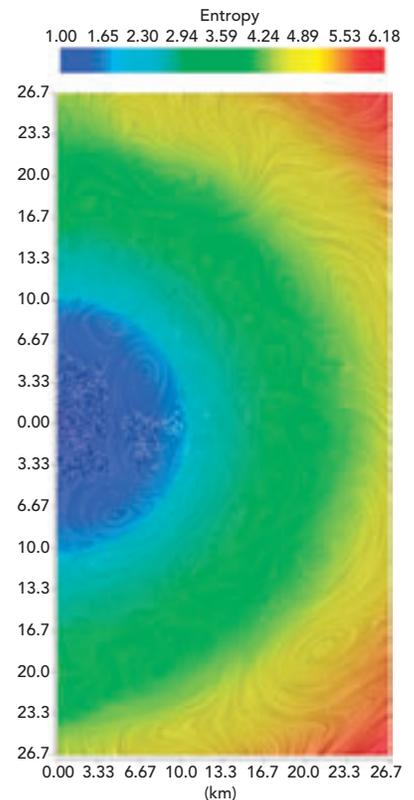


FIGURE 21. A closeup look at a proto-neutron star model, shown at about 32 ms following core bounce, as calculated by the 2D radiation-hydrodynamic code, V2D. The distance scale is given in kilometers. This view focuses on the inner 25 km of a model that extends nearly 6000 km into the oxygen shell. The figure is color mapped by entropy. Velocity direction is shown using an LEA texture map. Convective motion can be seen to encompass the entire inner core. However, even at this early stage of evolution, vigorous motion has ceased, leaving this region moving at only a few percent of the sound speed.

rotational instability. Finally, the accretion of the rotational flow onto the central core can spin up the proto-neutron star to periods of order tens of milliseconds, consistent with the inferred rotational period of radio pulsars at birth.

Other team members were able to carry out the most physically detailed 2D core-collapse supernova simulations to date using preconditioners, which are approximating techniques for speeding up linear solvers. Earlier simulations used a gray approximation, which assumes a shape of the energy distribution spectrum for neutrinos. In contrast,

the TSI simulations using the V2D code computed the energy spectra of the neutrinos, which change with time and with position as the supernova evolves. That's far more realistic, but requires extensive computing power. With SciDAC providing access to terascale supercomputers, the team was able to solve the problem.

Using V2D, initial studies have shown a number of new results and confirmed some results reported by others. These include a much earlier onset of post-bounce radiative-hydrodynamic instabilities and development of a hydrodynamically unstable region that quickly grows to engulf the entire region inside the core-bounce shock wave, including both neutrino-transparent regions as well as the proto-neutron star where material is opaque to neutrinos. The most detailed simulations have proceeded to about 35 milliseconds following core bounce, and the team is working to extend these simulations to significant fractions of a second and beyond.

Interaction between astrophysicists and nuclear physicists on the TSI team resulted in the first stellar core collapse simulations to implement a sophisticated model for the stellar core nuclei. The increased sophistication in the nuclear models led to significant quantitative changes in the supernova models, demonstrating that both the microphysics and the macrophysics of core collapse supernovae must be computed with care. The significance of this work led to the publication of two associated Physical Review Letters, one focusing on the astrophysics and one focusing on the nuclear physics. The work has also motivated nuclear experiments to measure nuclear cross-sections as a way to validate the nuclear models used in supernova simulations.

In the last few years experiments and observations have revealed that neutrinos have mass and can change their flavors (electron, muon and tau). This development could turn out to

be significant for the core collapse supernova explosion problem, for associated nucleosynthesis and for a neutrino signal detection. Although some experimental results indicate that neutrinos would not be able to transform to the point where they could affect the shock wave, the TSI team has mapped out scenarios in which the behavior of neutrinos could in principle affect the dynamics and nucleosynthesis deep inside the supernova environment.

National Computational Infrastructure for Lattice Gauge Theory: Accomplishments and Opportunities

The long-term goals of high energy and nuclear physicists are to identify the fundamental building blocks of matter, and to determine the interactions among them that lead to the physical world we observe. The fundamental theory of the strong interactions between elementary particles is known as quantum chromodynamics, or QCD. The strong force is one of four fundamental forces in nature, and provides the force to bind quarks to construct protons and neutrons, which account for about 98 percent of the matter in nature.

The objective of the SciDAC **National Computational Infrastructure for Lattice Gauge Theory** project is to construct the computational infrastructure needed to study QCD. Nearly all high energy and nuclear physicists in the United States working on the numerical study of QCD are involved in this project, and the infrastructure created will be available to all. The project includes the development of community software for the effective use of terascale computers, and research and development of specialized computers for the study of QCD.

The Department of Energy supports major experimental, theoretical and computational programs aimed

at reaching these goals. Remarkable progress has been made through the development of the Standard Model of high energy and nuclear physics, which provides fundamental theories of the strong, electromagnetic and weak interactions. This progress has been recognized through the award of Nobel Prizes in Physics for the development of each of the components of the Standard Model: the unified theory of weak and electromagnetic interactions in 1979, and QCD, the theory of the strong interactions, in 1999 and 2004. However, our understanding of the Standard Model is incomplete because it has proven extremely difficult to determine many of the most interesting predictions of QCD, those that involve the strong coupling regime of the theory. To do so requires large scale numerical simulations within the framework of lattice gauge theory.

The scientific objectives of lattice QCD simulations are to understand the physical phenomena encompassed by QCD, and to make precise calculations of the theory's predictions. Lattice QCD simulations are necessary to solve fundamental problems in high energy and nuclear physics that are at the heart of DOE's large experimental efforts in these fields. Major goals of the experimental programs in high energy and nuclear physics on which lattice QCD simulations can have an important impact are to: (1) verify the Standard Model or discover its limits; (2) understand the internal structure of nucleons and other strongly interacting particles and (3) determine the properties of strongly interacting matter under extreme conditions, such as those that existed immediately after the Big Bang and are produced today in relativistic heavy-ion experiments. Lattice QCD calculations are essential to research in all of these areas.

The numerical study of QCD requires very large computational resources, and has been recognized as

one of the grand challenges of computational science. The advent of terascale computing, coupled with recent improvements in the formulation of QCD on the lattice, provide an unprecedented opportunity to make major advances in QCD calculations. The infrastructure created under this SciDAC grant will play a critical role in enabling the U.S. lattice QCD community to take advantage of these opportunities. In particular, the hardware research and development work provides the groundwork for the construction of dedicated computers for the study of QCD, and the community software will enable highly efficient use of these computers and the custom designed 12,228-node QCDOC computer recently constructed at Brookhaven National Laboratory (BNL).

Software Development

The project's software effort has created a unified programming environment, the QCD Application Programming Interface (QCD API), that enables members of the U.S. lattice gauge theory community to achieve high efficiency on terascale computers, including the QCDOC,

commodity clusters optimized for QCD, and commercial supercomputers. Design goals included enabling users to quickly adapt codes to new architectures, easily develop new applications and preserve their large investment in existing codes.

The QCD API is an example of an application-specific code base serving a national research community. It exploits the special features of QCD calculations that make them particularly well suited to massively parallel computers. All the fundamental components have been implemented and are in use on the QCDOC hardware at BNL, and on both the switched and mesh architecture Pentium 4 clusters at Fermi National Accelerator Laboratory (FNAL) and Thomas Jefferson National Accelerator Facility (JLab). The software code and documentation are publicly available via the Internet.

Hardware Research and Development

The second major activity under the Lattice QCD SciDAC grant has been the design, construction and operation of commodity clusters optimized for the study of QCD. This

work has taken place at FNAL and JLab. The objective has been to provide computing platforms to test the QCD API, and to determine optimal configurations for the terascale clusters planned for FY 2006 and beyond.

The clusters that have been built are being used to carry out important research in QCD. The bottleneck in QCD calculations on clusters, as on commercial supercomputers, is data movement, not CPU performance. QCD calculations take place on four-dimensional space-time grids, or lattices. To update variables on a lattice site, one only needs data from that site and a few neighboring ones. The standard strategy is to assign identical sub-lattices to each processor. Then, one can update lattice points on the interior of the sub-lattices, for which all the relevant neighbors are on the same processor, while data is being collected from a few neighboring processors to update lattice sites on the sub-lattice boundaries. This strategy leads to perfect load balancing among the processors, and, if the computer and code are properly tuned, to overlap of computation and communications.



FIGURE 22. Under SciDAC's Lattice QCD project, two computer clusters were developed for use by the lattice gauge theory community. The system on the left is the 6n cluster at DOE's Jefferson Lab. At the right is the QCDOC computer, constructed at Brookhaven National Laboratory by a group centered at Columbia University. The RBRC computer on the right side of the photo was funded by the Riken Research Institute in Japan. Together, the two systems have a peak speed of 20 teraflop/s.

Building a Better Software Infrastructure for Scientific Computing

Infrastructure usually brings to mind images of physical structures such as roads, bridges, power grids and utilities. But computers also rely on an infrastructure of processors, software, interconnects, memory and storage, all working together to process data and accomplish specific goals. While personal computers arrive ready to use with easy-to-install software, scientists who rely on supercomputers have typically developed their own software, often building on codes that were developed 10 or 20 years ago.

As supercomputers become more powerful and faster, computational scientists are seeking to adapt their codes to utilize this computing power to develop ever more detailed simulations of complex problems. Other scientists are using supercomputing centers to store, analyze and share data from large experimental facilities, such as particle colliders or telescopes searching deep space for supernovae.

But scaling up applications to run on the newest terascale supercomputers, which can perform tens of trillions of calculations per second, presents numerous challenges. First, many of these scientific applications were written by researchers over the course of years or decades for older computers with tens or hundreds of processors and may not be easily adapted to run on thousands or tens of thousands of processors. Second, in some cases, scaling an existing application to run on more processors can actually slow down the run time as the communication between processors

bogs down. Because these “legacy codes” often represent significant intellectual investment aimed at solving very specific problems, scientists are reluctant to start anew in developing new codes. Finally, even when applications can be adapted to run more efficiently on one computer architecture, they may not achieve similar performance on another type of supercomputer, an issue of growing importance as supercomputing centers increasingly share their computers to meet the growing demand for greater computing capability. Understanding computer performance and developing tools to enhance this performance enables scientists to scale up their codes to study problems in greater detail for better understanding.

To help bridge the gap between these legacy codes and terascale computing systems, SciDAC established seven Integrated Software Infrastructure Centers, or ISICs, which brought together experts from government computing centers, scientific disciplines and indus-

try to approach the problems from different perspectives to create broadly applicable solutions. In many cases, computing tools originally developed and refined for one type or problem were found to be useful in solving problems in other scientific areas, too. As a result of these focused efforts, scientists can now draw on a wide range of tools and techniques to ensure their codes perform efficiently on supercomputers of varying sizes, with different architectures and at various computing centers.

An important emphasis of SciDAC is that the ISIC teams worked with other projects to enable the effective use of terascale systems by SciDAC applications.

Specifically, the ISICs addressed the need for:

- new algorithms which scale to parallel systems having thousands of processors
- methodology for achieving portability and interoperability of complex high performance scientific software packages
- operating systems tools and support for the effective management of terascale and beyond systems, and
- effective tools for feature identification, data management and visualization of petabyte-scale scientific datasets.

The result is a comprehensive, integrated, scalable, and robust high performance software infrastructure for developing scientific applications, improving the performance of existing applications on different supercomputers and giving scientists new techniques for sharing and analyzing data.

Applied Mathematics ISICs

Three applied mathematics ISICs were created to develop algorithms, methods and mathematical

libraries that are fully scalable to many thousands of processors with full performance portability.

An Algorithmic and Software Framework for Applied PDEs (APDEC)

Many important DOE applications can be described mathematically as solutions to partial differential equations (PDEs) at multiple scales. For example, combustion for energy production and transportation is dominated by the interaction of fluid dynamics and chemistry in localized flame fronts. Fueling of magnetic fusion devices involves the dispersion of material from small injected fuel pellets. The successful design of high-intensity particle accelerators relies on the ability to accurately predict the space-charge fields of localized beams in order to control the beams, preserve the beam emittance and minimize particle loss.

The goal of the **Algorithmic and Software Framework for Applied PDEs (APDEC) ISIC**

project has been to develop a high performance algorithmic and software framework for multiscale problems based on the use of block-structured adaptive mesh refinement (AMR) for representing multiple scales. In this approach, the physical variables are discretized on a spatial grid consisting of nested rectangles of varying resolution, organized into blocks. This hierarchical discretization of space can adapt to changes in the solution to maintain a uniform level of accuracy throughout the simulation, leading to a reduction in the time-to-solution by orders of magnitude compared to traditional fixed-grid calculations with the same accuracy. In short, AMR serves as a computational microscope, allowing scientists to focus their computing resources on the most interesting aspects of a problem. The resulting algorithms have enabled important scientific discoveries in a number of disciplines.

The APDEC center has been involved with the development of applications in the areas of combustion simulations, magnetohydrody-

namic (MHD) modeling for fusion energy reactors called tokamaks, and modeling particle accelerators. In combustion, the center developed new, more efficient simulation methods for solving chemically reacting fluid flow problems. These algorithms and software have been used to study a variety of problems in the combustion of hydrocarbon fuels, as well as providing the starting point for the development of a new capability for simulating nuclear burning in Type 1a supernovae.

In the area of fusion research, the center developed AMR codes to investigate a variety of scientific problems, including tokamak fueling (Figure 1) and magnetic reconnection, which can help in the design of fusion reactors for future energy production.

In the area of accelerator modeling, two new capabilities were developed: an AMR-particle-in-cell (AMR-PIC) algorithm that has been incorporated into the MLI and Impact beam dynamics codes (Figure 2); and a prototype capability for use in simulating gas dynamics

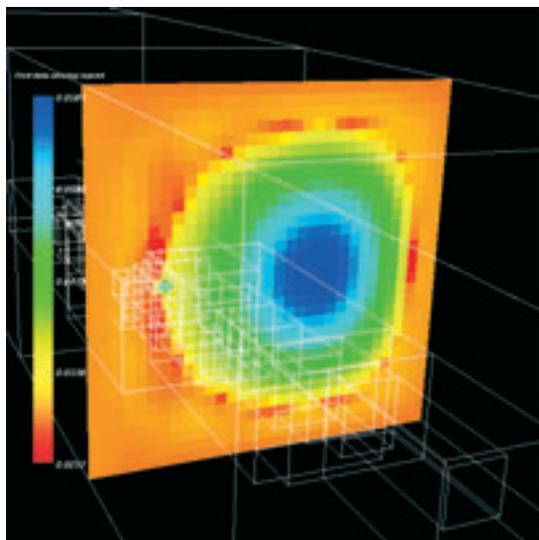


FIGURE 1. AMR simulation of a fuel pellet being injected into a tokamak. Image shows the pressure along a slice as well as the outlines of the various refined grid patches (in collaboration with the SciDAC CEMM project).

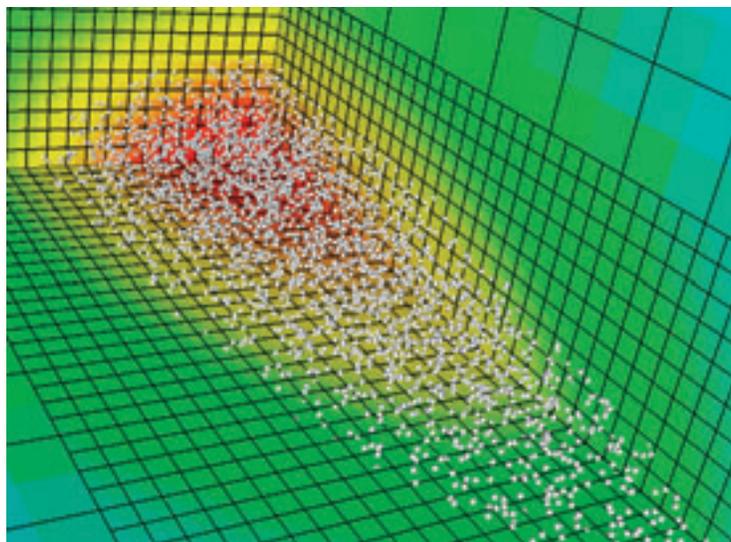


FIGURE 2. AMR particle-in-cell calculation, showing the electrostatic field induced by the particles (in collaboration with the SciDAC AST project).

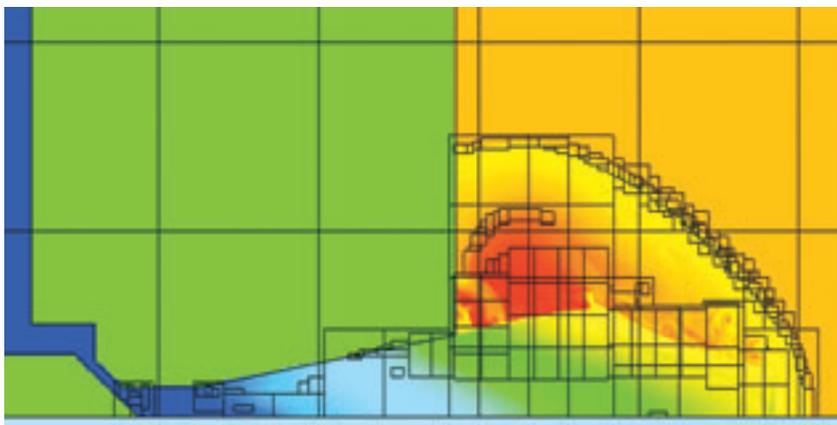


FIGURE 3. Axisymmetric gas jet expanding into a vacuum, with the axis of symmetry along the bottom of the figure, in a laser-driven plasma-wakefield accelerator (in collaboration with the SciDAC AST project).

problems in jets and capillary tubes arising in the design of plasma-wakefield accelerators (Figure 3).

To manage these and other complex AMR algorithms and software, the APDEC team used a collection of libraries written in a combination of programming languages. This software architecture, which maps the mathematical structure of the algorithm space onto a hierarchy of software layers, enables the codes to be more easily adapted and reused across multiple applications. The project team has also developed new grid-generation tools for calculations using mathematical representations from microscopy, medical image data and geophysical data.

Since the APDEC was established, a guiding principle has been to develop a software infrastructure that performs as efficiently as possible. By creating algorithms to efficiently use the number of available processors and combining these with tools for load balancing, the result is software which has a computational cost per grid point and scalability comparable to that of uniform-grid calculations using the same algorithms, giving scientists more detailed results without requiring additional computing resources.

For AMR algorithms for simple compressible flow models, the APDEC project has measured 75 percent efficiency up to 1,024 processors on IBM SP and HP Alpha systems. For incompressible flow, they have observed 75 percent efficiency up to 256 processors. For low-Mach-number combustion, production calculations on IBM SP and SGI systems are typically done using 128–512 processors, and the efficiency of the APDEC algorithms has enabled significant scientific achievements in combustion and astrophysics problems.

For many problems, such as combustion applications, a principal barrier to scaling past a few hundred processors is the performance of iterative methods for solving elliptic equations. To address this problem, the APDEC Center is developing a new class of AMR solvers in three dimensions. The goal is to create a solver with a far lower communications-to-computation ratio than traditional iterative methods, which means the solution

can be calculated much more quickly. A preliminary implementation of this algorithm has obtained 85 percent efficiency on 1,024 processors, with less than 4 percent of the execution time spent in communications. The APDEC adaptive calculation used 3 billion grid points, as compared to the 2 trillion grid points that would have been required using a uniform grid algorithm at the same resolution, leading to a reduction in computing time of at least two orders of magnitude. This solver is the basis for the AMR-PIC method developed for the beam dynamics codes mentioned above.

Principal Investigator: Phil Colella,
Lawrence Berkeley National Laboratory

Terascale Optimal PDE Solvers (TOPS) ISIC

In many areas of science, physical experimentation is impossible, such as with cosmology; dangerous, as with manipulating the climate; or simply expensive, as with fusion reactor design. Large-scale simulations, validated by comparison with related experiments in well-understood laboratory contexts, are used by scientists to gain insight and confirmation of existing theories in such areas, without benefit of full experimental verification. But today's high-end computers, such as those at DOE's supercomputing centers, are one-of-a-kind, and come without all of the scientific software libraries that scientists expect to find on desktop workstations.

The **Terascale Optimal PDE Solvers (TOPS) ISIC** was created to develop and implement algo-

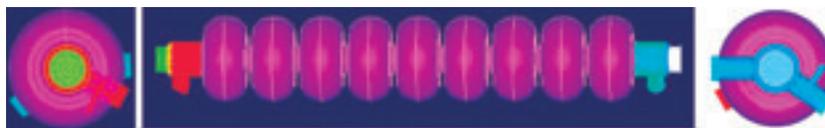


FIGURE 4. Model of the ILC low-loss cavity

gorithms and support scientific investigations performed by DOE-sponsored researchers. These simulations often involve the solution of partial differential equations (PDEs) on terascale computers. The TOPS Center researched, developed and deployed an integrated toolkit of open-source, optimal complexity solvers for the nonlinear partial differential equations that arise in many DOE application areas, including fusion, accelerator design, global climate change and reactive chemistry. The algorithms created as part of this project were also designed to reduce current computational bottlenecks by orders of magnitude on terascale computers, enabling scientific simulation on a scale heretofore impossible.

Nonlinear PDEs give mathematical expression to many core DOE mission applications. PDE simulation codes require implicit solvers for the *multirate*, *multiscale*, *multicomponent*, *multiphysics* phenomena of hydrodynamics, electromagnetism, chemical reaction and radiation transport. Currently, such problems typically reach into the tens of millions of unknowns – and this size is expected to increase 100-fold in just a few years.

Unfortunately, the algorithms traditionally used to solve PDEs were designed to address smaller problems and become much less efficient as the size of the system being studied increases. This creates a double jeopardy for applications, particularly in the case when the algorithms are based on iterations – as it takes more computing resources to solve each step of the problem, and the number of steps also goes up. Fortunately, the physical origin of PDE problems often allows them to be addressed by using a sequence of approximations, each of which is smaller than the one before. The solutions to the approximations, which may be obtained more efficiently, are combined judiciously to provide the solution to the original

problems. One well-known example of this approach is called the multi-grid method, which solves a problem by tackling a coarse approximation and then using the solution to generate the solution of a better approximation, and so on. It can be shown that the performance of multigrid method is optimal for certain classes of problems. The underlying philosophy of this and other similar approaches is to make the majority of progress towards a high quality result through less complex intermediate steps.

The efforts defined for TOPS and its collaborations with other projects have been chosen to revolutionize large-scale simulation through incorporation of existing and new optimal algorithms and code interoperability. TOPS provides support for the software packages Hypre, PARPACK, PETSc, ScaLAPACK, Sundials, SuperLU and TAO, some of which are in the hands of thousands of users, who have created a valuable experience base on thousands of different computer systems.

Software developed and supported by the TOPS project is being used by scientists around the globe. In the past few years, researchers outside the SciDAC community authored more than two dozen scientific papers reporting results achieved using TOPS software. The papers cover many disciplines, from astronomy to chemistry to materials science to nanotechnology to optics.

TOPS solver software has also been incorporated into numerous other packages, some commercial and some freely available. Among the widely distributed packages maintained outside of the SciDAC program that employ or interface to TOPS software “under the hood” are Dspice, EMSolve, FEMLAB, FIDAP, Global Arrays, HP Mathematical Library, libMesh, Magpar, Mathematica, NIKE, Prometheus, SCIRun, SLEPc, Snark and Trilinos.

Finally, TOPS software is taught in many courses in the U.S. and abroad, and forms a core of the annual training workshop sponsored by DOE to introduce researchers to the Advanced Computational Software (ACTS) Collection. TOPS software is also regularly featured in short courses at professional meetings.

Principal Investigator: David Keyes,
Cornell University

The Terascale Simulation Tools and Technology (TSTT) Center

Terascale computing provides an unprecedented opportunity to achieve numerical simulations at levels of detail and accuracy previously unattainable. DOE scientists in many different application areas can reach new levels of understanding through the use of high-fidelity calculations based on multiple coupled physical processes and multiple interacting physical scales. The best way to achieve detailed and accurate simulations in many application areas, and frequently the only way to obtain useful answers, is to use adaptive, composite, hybrid approaches. However, the best strategy for a particular simulation is not always clear and the only way to find the best method is to experiment with various options. This is both time-consuming and difficult due to the lack of easy-to-apply, interoperable meshing, discretization and adaptive technologies.

The **Terascale Simulation Tools and Technologies (TSTT) Center** was established to address the technical and human barriers preventing the effective use of powerful adaptive, composite, and hybrid methods. The primary objective was to develop technologies that enable application scientists to easily use multiple mesh and discretization strategies within a single simulation on terascale computers. A key aspect of the project is to develop

services that can be used interoperably to enable mesh generation for representing complex and possibly evolving domains, high-order discretization techniques for improved numerical solutions, and adaptive strategies for automatically optimizing the mesh to follow moving fronts or to capture important solution features.

The center encapsulated its research into software components with well-defined interfaces that enable different mesh types, discretization strategies and adaptive techniques to interoperate in a “plug and play” fashion. All software was designed for terascale computing environments with particular emphasis on scalable algorithms for hybrid, adaptive computations and single processor performance optimization.

To ensure the relevance of the project’s research and software developments to the SciDAC goals, the TSTT team collaborated closely with both SciDAC application researchers and other ISICs. Specifically, TSTT technologies were integrated into fusion, accelerator design, climate modeling and biology applications to both provide near-term benefits for those efforts as well as receive the feedback for further improvements.

Working with the Accelerator Modeling project, TSTT researchers generated high-quality meshes from CAD (computer-aided design) models for accurate modeling of the interaction region of the PEP II accelerator. This led to the first successful simulation with a transit beam, using the Tau3P accelerator modeling application. This success led to a decision to continue to use Tau3P for PEP-II computations for reaching higher luminosities. Similar mesh generation efforts by TSTT researchers were used to analyze the wakefield in the advanced Damped Detuned Structure and verify important performance characteristics of the system.

TSTT also partnered with SLAC

and the TOPS ISIC to create an automatic design optimization loop to provide automatic tuning of accelerator geometries to significantly increase the speed and decrease the cost by which new accelerators can be designed.

In the field of plasma physics, TSTT researchers helped the Center for Extended MHD Modeling (CEMM) solve a number of problems, particularly the case of anisotropic diffusion, after determining that using fewer higher-order elements significantly decreased the total solution time needed to obtain a given accuracy. This demonstration resulted in a new effort by scientists at Princeton Plasma Physics Laboratory (PPPL) to develop fifth-order finite elements for their primary application code, M3D. TSTT researchers also worked with scientists at PPPL to insert adaptive mesh refinement technologies and error estimators directly into the new high-order M3D code.

TSTT researchers are collaborating with climate modeling scientists to develop enhanced mesh generation and discretization capabilities for anisotropic planar and geodesic surface meshes. The initial mesh is adapted and optimized to capture land surface orographic or topographic height fields. This technology has improved the prediction of rainfall, snowfall and cloud cover in regional weather models in prototype simulations.

TSTT researchers are collaborating with computational biologists to develop the Virtual Microbial Cell Simulator (VMCS), which is targeting DOE bioremediation problems to clean up heavy metal waste using microbes. Methods developed by TSTT were used to study how microbe communities aggregate in certain environments, providing new insight into the behavior of these microbes.

In partnership with combustion scientists, TSTT researchers have developed a new tool that combines

adaptive mesh refinement with front tracking and used this technology to develop a new capability for diesel engine design. This effort has emphasized the modeling of the instability and breakup of a diesel jet into spray, thought to be the first such effort to provide a predictive capability in this arena.

Principal Investigators: Jim Glimm, State University of New York – Stony Brook; Lori Freitag Diachin, Lawrence Livermore National Laboratory

Computer Science ISICs

The software infrastructure vision of SciDAC is for a comprehensive, portable, and fully integrated suite of systems software and tools for the effective management and utilization of terascale computational resources by SciDAC applications. The following four Computer Science ISIC activities addressed critical issues in high performance component software technology, large-scale scientific data management, understanding application/architecture relationships for improved sustained performance, and scalable system software tools for improved management and utility of systems with thousands of processors.

Center for Component Technology for Terascale Simulation Software

The **Center for Component Technology for Terascale Simulation Software (CCTSS)** was created to help accelerate computational science by bringing a “plug and play” style of programming to high performance computing. Through a programming model called the Common Component Architecture (CCA), scientists can dramatically reduce the time and effort required to compose independently created software libraries

into new terascale applications. This approach has already been adopted by researchers in the application areas of combustion, quantum chemistry and climate modeling, with new efforts beginning in fusion and nanoscale simulations.

While the CCA approach has been more common in industry, the goal of the CCTTSS was to bring a similar approach to scientific computing. Unlike more narrowly focused commercial applications, the scientific CCA effort faced such challenges as maintaining high performance, working with a broad spectrum of scientific programming languages and computer architectures, and preserving DOE investments in legacy codes.

An example of the CCA approach in action is a prototype application being developed within the Community Climate System Model (CCSM) project. The CCA is used at the level of system integration to connect skeleton components for the atmosphere, ocean, sea ice, land surface, river routing, and flux coupler. Prototype applications within the Earth System Modeling Framework (ESMF), the infrastructure targeted for the future CCSM, also employ the CCA, including reusable components for visualization and connectivity.

In addition, CCTTSS researchers are collaborating closely with application scientists to create high-performance simulations in quantum chemistry and combustion. Moreover, new externally funded projects incorporate CCA concepts in applications involving nanotechnology, fusion, and underground transport modeling, and proposals have been recently submitted involving biotechnology and fusion.

As a part of its mission, the CCTTSS has developed production components that are used in scientific applications as well as prototype components that aid in teaching CCA concepts. These freely available components include vari-

ous service capabilities, tools for mesh management, discretization, linear algebra, integration, optimization, parallel data description, parallel data redistribution, visualization, and performance evaluation. The CCTTSS is also collaborating with the APDEC, TSTT and TOPS SciDAC centers to define common interfaces for mesh-based scientific data management as well as linear, nonlinear, and optimization solvers.

SciDAC funding has accelerated both CCA technology development and the insertion of this technology into massively parallel scientific applications, including major SciDAC applications in quantum chemistry and combustion. Additionally, numerous CCA-compliant application components have been developed and deployed, and the underlying infrastructure is maturing and establishing itself in the scientific community.

Principal Investigator: Rob Armstrong,
Sandia National Laboratories

High-End Computer System Performance: Science and Engineering ISIC

As supercomputers become ever more powerful tools of scientific discovery, demand for access to such resources also increases. As a result, many scientists receive less time on supercomputers than they would like. Therefore, it's critical that they make the most of their time allocations. One way to achieve this optimal efficiency is to analyze the performance of both scientific codes and the computer architectures on which the codes are run. The resulting improvements can be dramatic. In one instance, an astrophysics code which had been running at about 10 percent of a supercomputer's theoretical peak speed (which is typical) was able to run at more than 50 percent efficiency after careful performance analysis and tuning.

SciDAC's **High-End Computer**

System Performance: Science and Engineering ISIC, also known as the Performance ISIC or PERC, focused on developing tools and techniques to help scientists determine how they can best execute a specific application on a given computer platform. The research was aimed at determining what level of achievable performance is realistic, how scientific applications can be accelerated toward these levels, and how this information can drive the design of future applications and high-performance computing systems.

In addition to helping scientists better understand performance behavior of their codes, the major improvements to computer performance modeling technology made by PERC researchers are also helping supercomputer centers to better select systems, thereby ensuring higher scientific productivity on these publicly funded systems. Such supercomputer procurements run to millions of dollars, and PERC research is being used by centers operated by DOE, the National Science Foundation and the Department of Defense around the country. PERC researchers have also collaborated with several SciDAC scientific application efforts, resulting in major speedups for several high-profile computer codes used in these projects.

For the **Community Climate System Model (CCSM)**, one of the United States' primary applications for studying climate change, PERC researchers helped determine optimal algorithmic settings for the Community Atmosphere Model (CAM), a key component of the CCSM, when running the Intergovernmental Panel on Climate Change scenario runs on the IBM p690 cluster computer at ORNL, accelerating the completion of this milestone. Similar studies are ongoing on the SGI Altix, Cray XD1, and Cray XT3 supercomputers. The resulting data from these runs will contribute to an international report on global climate change.

Working with the **Plasma Microturbulence Project (PMP)**, PERC researchers applied Active Harmony, a software tool supporting distributed execution of computational objects, to GS2, a gyrokinetic turbulence simulator used by the fusion energy research community. The result was a 2.3 to 3.4 times speedup of GS2 for a common configuration used in production runs.

PERC researchers worked with the **Terascale Supernovae Initiative (TSI)** project to port and optimize the EVH1 hydrodynamics code on the Cray X1 supercomputer, achieving excellent performance for large problems. The EVH1 performance analysis was completed for up to 256 processors on all current target platforms.

Working with two other ISICs, the **Terascale PDE Simulations (TOPS)** and **Terascale Simulation Tools and Technology (TSTT)** centers, PERC researchers analyzed the performance of a TOPS-TSTT mesh smoothing application and found that the sparse matrix-vector multiply achieves 90 percent of the peak performance imposed by the memory bandwidth limit.

PERC researchers collaborated closely with SciDAC's **Lattice Gauge Theory** project, conducting performance analyses of the MILC (MIMD Lattice Computation) application on Pentium-3 Linux clusters. The PERC team collected detailed performance data on different aspects of the code and identified the key fragments that affect the code performance, then presented their findings to the lattice gauge theory community.

In the **Accelerator Science and Technology** program, the PERC team has shown how to improve the performance of the post-processing phase within the Omega3P application at the Stanford Linear Accelerator Center. Post-processing can consume 40 percent or more of total execution time in a full run of

Omega3P, where most of this time is due to a complex and redundant computation. By exchanging this redundant computation for cached lookup to pre-computed data, researchers were able to achieve 4.5 times speedups of the post-processing phase, and 1.5 time speedups of a full Omega3P run.

In addition to working directly with scientists on the performance of codes, PERC researchers have made major improvements in the usability and effectiveness of several performance tools, enabling researchers to analyze their codes more easily. The result is an integrated suite of measurement, analysis and optimization tools to simplify the collection and analysis of performance data and help users in optimizing their codes.

Principal Investigator: David Bailey,
Lawrence Berkeley National Laboratory

Scalable Systems Software for Terascale Computer Centers

As terascale supercomputers become the standard environment for scientific computing, the nation's premiere scientific computing centers are at a crossroads. Having built the operations of their systems around home-grown systems software, system administrators and managers now face the prospect of rewriting their software, as these incompatible, ad hoc sets of systems tools were not designed to scale to the multi-teraflop systems that are being installed in these centers today. One solution would be for each computer center to take its home-grown software and rewrite it to be scalable. But this would incur a tremendous duplication of effort and delay the availability of terascale computers for scientific discovery.

The goal of the SciDAC's **Scalable Systems Software** project is to provide a more timely and cost-effective solution by pulling together representatives from the major computer centers and indus-

try and collectively defining standardized interfaces between system components. At the same time, this group is producing a fully integrated suite of systems software components that can be used by the nation's largest scientific computing centers.

The Scalable Systems Software suite is being designed to support computers that scale to very large physical sizes without requiring that the number of support staff scale along with the machine. But this research goes beyond just creating a collection of separate scalable components. By defining a software architecture and the interfaces between system components, the Scalable Systems Software research is creating an interoperable framework for the components. This makes it much easier and cost effective for supercomputer centers to adapt, update, and maintain the components in order to keep up with new hardware and software. A well-defined interface allows a site to replace or customize individual components as needed. Defining the interfaces between components across the entire system software architecture provides an integrating force between the system components as a whole and improves the long-term usability and manageability of terascale systems at supercomputer centers across the country.

The Scalable Systems Software project is a catalyst for fundamentally changing the way future high-end systems software is developed and distributed. The project is expected to reduce facility management costs by reducing the need to support home-grown software while making higher quality systems tools available. The project will also facilitate more effective use of machines by scientific applications by providing scalable job launch, standardized job monitoring and management software, and allocation tools for the cost-effective management and utilization of terascale and petascale computer resources.

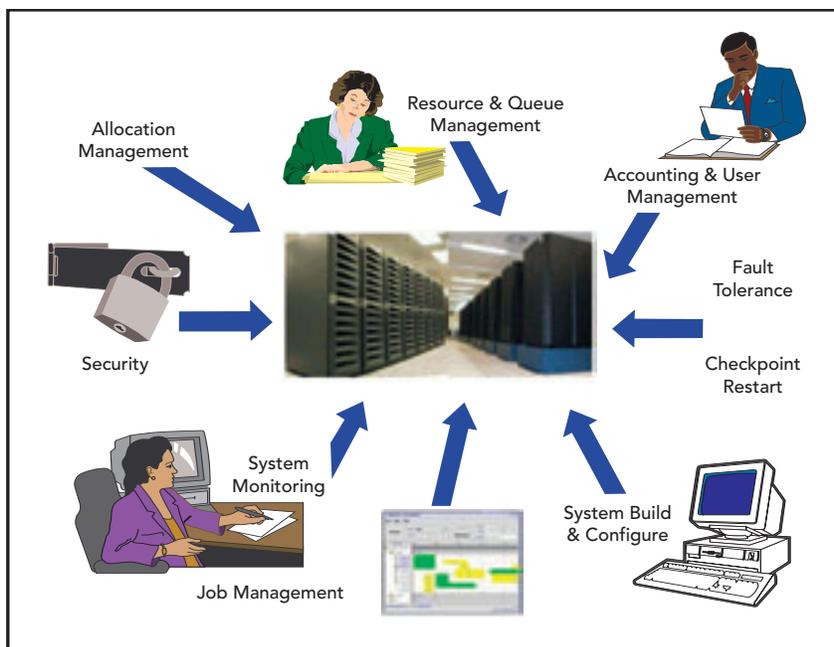


FIGURE 5. The mission of the Scalable Systems Software center is the development of an integrated suite of systems software and tools for the effective management and utilization of terascale computational resources, particularly those at the DOE facilities.

Here are some of the highlights of the project:

Designed modular architecture:

A critical component of the project's software design is its modularity. The ability to plug and play components is important because of the diversity in both high-end systems and the individual site management policies.

Defined XML-based interfaces that are independent of language and wire protocol: The interfaces between all the components have been fully documented and made publicly available, thereby allowing others to write replacement components (or wrap their existing components) as needed. The reference implementation has a mixture of components written in a variety of common programming languages, which allows a great deal of flexibility to the component author and allows the same interface to work on a wide range of hardware architectures and facility policies.

Reference implementation released: Version 1.0 of the fully integrated systems software suite was

released at the SC2005 conference – the leading international meeting of the supercomputing community – and will be followed by quarterly updates for the following year.

Production users: Full system software suites have been put into production on clusters at Ames Laboratory and Argonne National Laboratory. Pacific Northwest National Laboratory and the National Center for Supercomputer Applications have adopted one or more components from the suite, and the project team is in discussion with Department of Defense sites about use of some of the system software components.

Adoption of application programming interface (API): The suite's scheduler component is the widely used Maui Scheduler. The public Maui release (as well as the commercial Moab scheduler) has been updated to use the public XML interfaces and has added new capabilities for fairness, higher system utilization, and improved response time. All new Maui and Moab installations worldwide (more

than 3,000/month) now use the system software interfaces developed in this ISIC. Prominent users include 75 of the world's top 100 supercomputers and commercial industries such as Amazon.com and Ford Motor Co.

Principal Investigator: Al Geist, Oak Ridge National Laboratory

The Scientific Data Management Center

Scientific exploration and discovery typically takes place in two phases: generating or collecting data and then analyzing the data. In the data collection/generation phase, large volumes of data are generated by simulation programs running on supercomputers or collected from experiments. As experimental facilities and computers have become larger and more powerful, the amount of resulting data threatens to overwhelm scientists. Without effective tools for data collection and analysis, scientists can find themselves spending more time on data management than on scientific discovery.

The **Scientific Data Management Center (SDM)** was established to focus on the application of known and emerging data management technologies to scientific applications. The center's goals are to integrate and deploy software-based solutions to the efficient and effective management of large volumes of data generated by scientific applications. The purpose is not only to achieve efficient storage and access to the data using specialized indexing, compression, and parallel storage and access technology, but also to enhance the effective use of the scientists' time by eliminating unproductive simulations, by providing specialized data-mining techniques, by streamlining time-consuming tasks, and by automating the scientists' workflows.

The center's approach is to provide an integrated scientific data

management framework where components can be chosen by scientists and applied to their specific domains. By overcoming the data management bottlenecks and unnecessary information-technology overhead through the use of this integrated framework, scientists are freed to concentrate on their science and achieve new scientific insights.

Today's computer simulations and large-scale experiments can generate data at the terabyte level, equivalent to about 5 percent of all the books in the Library of Congress. Keeping up with such a torrent of data equires efficient parallel data systems. In order to make use of such amounts of data, it is necessary to have efficient indexes and effective analysis tools to find and focus on the information that can be extracted from the data, and the knowledge

learned from that information. Because of the large volume of data, it is also useful to perform analysis as the data are generated. For example, a scientist running a thousand-time-step, 3D simulation can benefit from analyzing the data generated by the individual steps in order to steer the simulation, saving unnecessary computation, and accelerating the discovery process. However, this requires sophisticated workflow tools, as well as efficient dataflow capabilities to move large volumes of data between the analysis components. To enable this, the center uses an integrated framework that provides a scientific workflow capability, supports data mining and analysis tools, and accelerates storage access and data searching. This framework facilitates hiding the details of the underlying parallel and indexing technology, and streamlining the

assembly of modules using process automation technologies.

Since it was established, SDM has adopted, improved and applied various data management technologies to several scientific application areas. By working with scientists, the SDM team not only learned the important aspects of the data management problems from the scientists' point of view, but also provided solutions that led to actual results. The successful results achieved so far include:

- More than a tenfold speedup in writing and reading NetCDF files was achieved by developing Parallel NetCDF software on top of the MPI-IO.
- An improved version of PVFS is now offered by cluster vendors, including Dell, HP, Atipa, and Cray.
- A method for the correct classification of orbits in puncture plots from the National Compact Stellarator eXperiment (NCSX) at PPPL was developed by converting the data into polar coordinates and fitting second-order polynomials to the data points.
- A new specialized method for indexing using bitmaps was used to provide flexible efficient search over billions of collisions (events) in high energy physics applications. A paper on the system received best paper award at the 2004 International Supercomputing Conference.
- The development of a parallel version of the popular statistical package R is being applied to fusion, geographic information systems and proteomics (biochemistry) applications.
- A scientific workflow system was developed and applied to the analysis of microarray data, as well as astrophysics applications. The system greatly increased the efficiency of running repetitive processes.

Principal Investigator: Arie Shoshani,
Lawrence Berkeley National Laboratory

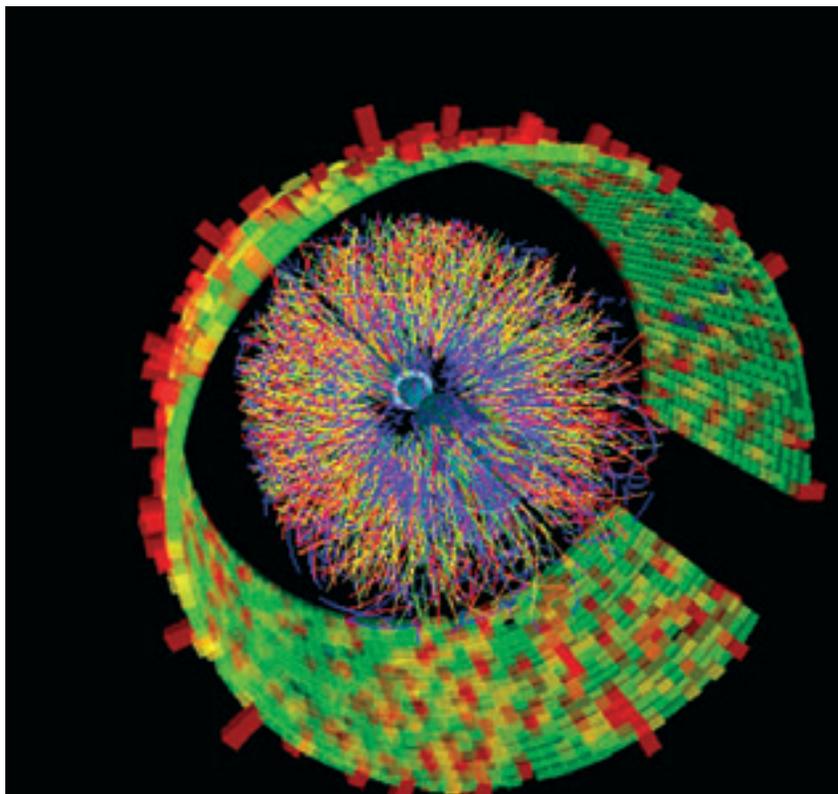


FIGURE 6. The STAR Experiment at Brookhaven National Laboratory generates petabytes of data resulting from the collisions of millions of particles. The Scientific Data Management ISIC developed software to make it easier for scientists to transfer, search and analyze the data from such large-scale experiments.

Creating an Advanced Infrastructure for Increased Collaboration

A hallmark of many DOE research efforts is collaboration among scientists from different disciplines. The multidisciplinary approach is the driving force behind the “big science” breakthroughs achieved at DOE national laboratories and other research institutions. By creating teams comprising physicists, chemists, engineers, applied mathematicians, biologists, earth scientists, computer scientists and others, even the most challenging scientific problems can be defined, analyzed and addressed more effectively than if tackled by one or two scientists working on their own.

With the advent of high-performance computers and networks, such teams can now be assembled virtually, with a multitude of research talents brought to bear on scientific problems of national and global importance. DOE scientists and engineers have been developing technologies in support of these “collaboratories.” The goal is to make collaboration among scientists at geographically dispersed sites as effective as if they were working at the same location. A key aspect of collaboratories is providing seamless and secure access to experimental facilities and computing resources at sites around the country.

As these collaborative experiments, whether a particle accelerator in New York or a massive telescope atop a volcano in Hawaii, come on line, scientists will be inundated with massive amounts of data which must be shared, analyzed, archived and retrieved by multiple users from multiple sites. Large research pro-

grams are producing data libraries at the petabyte level, the equivalent of 500 billion standard-size document pages, or 100 times the data volume of the U.S. Library of Congress. Tools which enable the easy transfer of huge datasets are essential to the success of these research efforts.

Under SciDAC, a number of collaboratory infrastructure projects were established to advance collaborative research. These projects incorporated technologies from distributed systems often referred to as “Grids.” Built upon networks like the Internet, a computing Grid is a service for sharing computer power and data storage capacity over the Internet, resulting in large, distributed computational resources. DOE is expanding the concept of Grids beyond computers to integrate access to experimental facilities and data archives around the world. While the idea is straightforward, developing a unified, secure infra-

structure linking various systems with unique security and access procedures, sometimes across international boundaries, is a daunting task.

Grid technology is evolving to provide the services and infrastructure needed for building “virtual” systems and organizations. A Grid-based infrastructure provides a way to use and manage widely distributed computing and data resources to support science. The result is a standard, large-scale computing, data, instrument, and collaboration environment for science that spans many different projects, institutions, and countries.

National Collaboratory Projects

DOE’s investment in National Collaboratories includes four SciDAC projects focused on the goal of creating collaboratory software environments to enable geographically separated scientists to effectively work together as a team and to facilitate remote access to both facilities and data.

The DOE Science Grid

A significant portion of DOE science is already, or is rapidly becoming, a distributed endeavor involving many collaborators who are frequently working at institutions across the country or around the world. Such collaborations typically have rapidly increasing needs for high-speed data transfer and high-performance computing. More and more, these requirements must be addressed with computing, data,



and instrument resources that are often more widely distributed than the collaborators. Therefore, developing an infrastructure that supports widely distributed computing and data resources is critical to DOE's leading-edge science.

Such infrastructures have become known as Grids, which provide secure and reliable access to these critical resources. While the individual institutions have their own policies and procedures for providing access to users, successful deployment of a Grid requires that a common approach be implemented so that when an authorized user logs into a Grid, all appropriate resources can be easily accessed.

Under SciDAC, the DOE Science Grid was developed and deployed across the DOE research complex to provide persistent Grid services to advanced scientific applications and problem-solving frameworks. By reducing barriers to the use of remote resources, it has made significant contributions to SciDAC and the infrastructure required for the next generation of science. The DOE Science Grid testbed was initially deployed among five project sites: the National Energy Research Scientific Computing Center and Argonne, Lawrence Berkeley, Oak Ridge and Pacific Northwest national laboratories.

In addition to the construction of a Grid across five major DOE facilities with an initial complement of computing and data resources, other accomplishments include:

- Integration into the Grid of the large-scale computing and data storage systems at NERSC, DOE's Office of Science flagship supercomputing center.
- Design and deployment of a Grid security infrastructure that is facilitating collaboration between U.S. and European high energy physics projects, and within the U.S. magnetic fusion community. This infrastructure provides a global, policy-

based method of identifying and authenticating users, which leads to a "single sign-on" so that any system on the Grid can accept a uniform user identity for authorization. This work is currently used by DOE's Particle Physics Data Grid, Earth Systems Grid, and Fusion Grid projects.

- A resource monitoring and debugging infrastructure that facilitates managing this widely distributed system and the building of high performance distributed science applications.
- Development and deployment partnerships established with several key vendors.
- Use of the Grid infrastructure by applications from several disciplines – computational chemistry, groundwater transport, climate modeling, bioinformatics, etc.
- In collaboration with DOE's Energy Sciences Network (ESnet), the design and deployment of the DOE Grids Certification Authority (Grid authentication infrastructure) that provides Grid identity certificates to users, systems, and services, based on a common, science collaboration oriented policy that is defined and administered by the subscribing virtual organizations.

Principal Investigator: William Johnston,
Lawrence Berkeley National Laboratory

The National Fusion Collaboratory



Developing a reliable energy system that is economically and environmentally sustainable is the long-term goal of DOE's Fusion Energy Science (FES) research program. In the U.S., FES experimental research is centered at three large facilities with a replacement value of over \$1 billion. As these experiments have increased in size and complexity, there has been a concurrent growth in the number and importance of collaborations among large groups at the experimental sites and smaller groups located nationwide.

The next generation fusion experimental device is the ITER reactor, an international collaboration with researchers from China, Europe, India, Japan, Korea, Russia and the U.S. ITER will be a burning fusion plasma magnetic confinement experiment located in France. With ITER located outside the U.S., the ability to maximize its value to the U.S. program is closely linked to the development and deployment of collaborative technology. As a result of the highly collaborative present and future nature of FES research, the community is facing new and unique challenges.

The **National Fusion Collaboratory** established under SciDAC unites fusion and computer science researchers to directly address these challenges by creating and deploying collaborative software tools. In particular, the project has developed and deployed a national FES Grid (FusionGrid) that is a system for secure sharing of computation, visualization and data resources over the Internet. This represents a fundamental paradigm shift for the fusion community where data, analysis and simulation codes, and



FIGURE 1. The control rooms of NSTX (a), DIII-D (b), and C-Mod (c) with shared display walls being used to enhance collocated collaboration. On the DIII-D shared display is also video from remote collaborators in Europe who were participating in that day's experiment.

visualization tools are now thought of as network services. In this new paradigm, access to resources (data, codes, visualization tools) is separated from their implementation, freeing the researcher from needing to know about software implementation details and allowing a sharper focus on the physics. The long-term goal of FusionGrid is to allow scientists at remote sites to participate as fully in experiments and computational activities as if they were working on site, thereby creating a unified virtual organization of the geographically dispersed U.S. magnetic fusion community – more than 1,000 researchers from over 40 institutions.

Fusion scientists use the FusionGrid to provide wide access to the complex physics codes and associated data used for simulating and analyzing magnetic fusion experiments. The first FusionGrid service came online in fall 2002. Since then, the service – a transport code known as TRANSP – has been used by scientists to produce over 6,000 code

runs and has supported analysis of fusion devices on three continents. TRANSP was also used to demonstrate the feasibility of between-shot analysis on FusionGrid.

Fusion scientists also use the FusionGrid as an advanced collaborative environment service that provides computer-mediated communications techniques to enhance work environments and to enable increased productivity for collaborative work. In January 2004 a fusion scientist in San Diego remotely led an experiment on the largest European magnetic fusion experiment, located in the United Kingdom, using FusionGrid's remote collaboration services. These services included multiple video images and a unified audio stream along with real-time secure access to data. With the introduction of FusionGrid's remote collaboration services, remote participation in fusion experiments is becoming more commonplace.

The National Fusion Collaboratory project is also working to



FIGURE 2. FusionGrid services are being used to collaborate in real time by fusion scientists in the U.S. and abroad. Pictured in the foreground is Dr. deGrassie (San Diego) leading the JET experiment (England), and on the screen is the JET control room and various JET data traces being discussed in real time.

enhance collaboration within the control rooms of fusion experiments. Shared display walls have been installed in the control rooms of the three large U.S. fusion experiments for enhanced large-group collaboration. The project has created and deployed unique software that presents the scientists with a multi-user environment allowing them to simultaneously share data to the large display and simultaneously interact with the display to edit, arrange, and highlight information. For collocated scientists, the ability to publish and share a scientific graph on the display wall results in easier and broader discussion compared to the old model of a few individuals gathering around a small computer monitor. For remote scientists, the ability of the shared display to accept video combined with audio has allowed them to interact and even lead an experiment.

Taken as a whole, the technologies being deployed by the National Fusion Collaboratory project are further facilitating data sharing and decision-making by both collocated and remote scientists. With the eyes of the worldwide fusion community shifting towards France for the next generation device, ITER, such capability becomes increasingly important for scientific success.

Principal Investigator: David Schissel,
General Atomics

The Particle Physics Data Grid

The Department of Energy's national laboratories are home to some of the world's leading facilities for particle physics experiments. Such experiments typically generate millions of particles which must be extensively analyzed and filtered to find and measure the exotic particles that herald new discoveries.

The **Particle Physics Data Grid** (PPDG) collaboration was launched as a collaboration between physicists and computer scientists to adapt and extend Grid technologies and the physicists' applications to build distributed systems to handle the storage, retrieval and analysis of the particle physics data at DOE's most critical research facilities.

PPDG has been instrumental in the significant progress made over the past five years in deploying Grid-enabled, end-to-end capabilities into the production data processing infrastructure of the participating experiments, and in making extended and robust Grid technologies generally available. Also, PPDG has successfully demonstrated the effectiveness of close collaboration between end users and the developers of new distributed technologies.

The experiments involved span those already in steady data-taking mode – where existing production systems were carefully extended step by step to take advantage of Grid technologies – to those developing new global systems for detector commissioning and the acquisition of data.

As a result of PPDG and other Grid projects in the U.S. and Europe, particle physics collaborations now include Grid-distributed computing as an integral part of their data processing models. By collaboratively deploying their developed technologies into significant production use, the computer science groups have extended their technologies and made them generally available to benefit other research communities.

In addition to the experiment-

specific end-to-end systems, through partnering with the National Science Foundation's iVDGL and GriPhyN projects and working with European counterparts such as the European Data Grid and Enabling Grids for ESciencE, PPDG has helped pave the way for a worldwide multi-science production Grid infrastructure. In the U.S., PPDG has played a leadership role in creating first the Grid3 and now the Open Science Grid infrastructure, which provides a common software installation, operational procedures and sharing of services for 50 laboratory and university sites.

The PPDG-supported improvements in data handling include a tenfold increase in the amount of sustained data distribution, a 50 percent reduction in the effort required to distribute data and analysis among five institutions, and up to a 50 percent increase in available resources due to Grid-enabled sharing. And even though funding for the PPDG has concluded, the collaborations established under the program continue to boost scientific productivity.

Significant scientific results and benefits that the technology developments and collaborations in PPDG have enabled include:

- The shortened turnaround time for analysis of the STAR nuclear physics experiment data produced early results in the first direct measurement of particles known as “open charm” production at the Relativistic Heavy Ion Collider. This was enabled through the experiment's collaboration with the Storage Resource Management group, which for the past few years has provided sustained transfer of more than 5 terabytes a week between Brookhaven National Laboratory in New York and Lawrence Berkeley National Laboratory in California.
- All results from the Fermilab Tevatron experiments, CDF and D0, now rely on distributed

computing using Grid technologies. Up to 20 sites around the world now receive and analyze data from the D0 experiment, which has led to new discoveries, such as results showing the mixing of B_s mesons, mesons that contain a “beauty quark” and a “strange quark.” The PPDG has contributed to such successes by improving a number of the software applications used to transfer and analyze data, more than doubling the efficiency of these operations. CDF seamlessly supports simulation and analysis jobs at up to 10 sites in Europe and the U.S.

- The BaBar experiment at the Stanford Linear Accelerator Center routinely moves all the data collected from the accelerator to computing centers in France, Italy and England. The Storage Resource Broker was extended to provide the distributed catalogs describing and managing the data. These catalogs are necessary to manage the location and selection information for the multipetabyte distributed datasets.
- The Large Hadron Collider (LHC) experiments are participating in a worldwide Grid project to provide global distribution and analysis of data from CERN in Switzerland. The U.S. ATLAS and U.S. CMS collaborations are developing and testing their physics algorithms and distributed infrastructure in preparation for the onslaught of tens of petabytes a year of data to be analyzed by 2008. At the end of 2005, ATLAS produced more than one million fully simulated events on the Open Science Grid with 20 different physics samples to test the data distribution and analysis capabilities. CMS is running “Data Challenges” to distribute data from the CERN accelerator location to the global sites. Data is transferred from

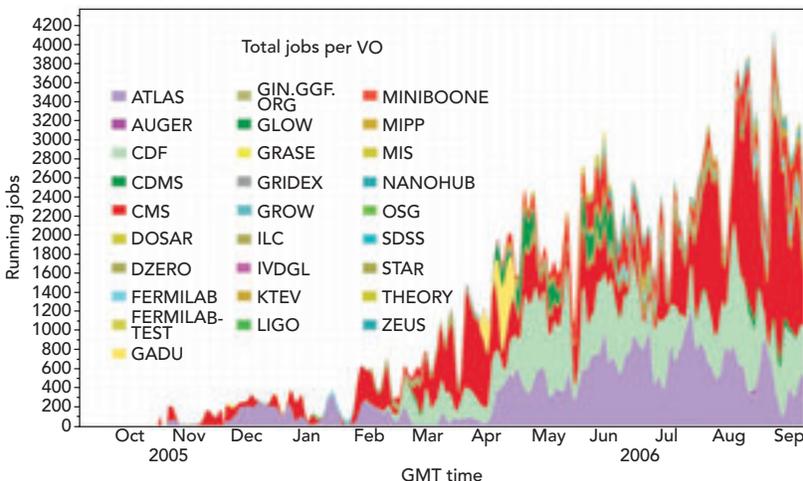


FIGURE 3: Monitored Jobs on Open Science Grid

the U.S. Tier-1 Center to sites around the world as part of the data distribution and analysis system. Peak transfer rates of ~5 Gbps are reached.

- The Open Science Grid Consortium provides a common infrastructure, validated software stack and ongoing operations and collaboration which all the groups in PPDG use and to which PPDG has made significant contributions. The use of Open Science Grid has increased steadily since its launch in July 2005, and it is hoped to continue with this work to benefit all the PPDG collaborators over the next few years. Ideally, the significant progress made by PPDG towards a usable, global computing infrastructure supporting large institutions can be extended to benefit smaller research programs and institutions as well.

Principal Investigators: Richard Mount, Stanford Linear Accelerator Center; Harvey Newman, California Institute of Technology; Miron Livny, University of Wisconsin Madison; Ruth Pordes, Fermi National Accelerator Laboratory (added 2003).

The Earth System Grid II: Turning Climate Model Datasets into Community Resources

One of the most important – and challenging – scientific problems which requires extensive computational simulations is global climate change. The United States’ climate research community has created the Community Climate System Model (CCSM), one of the world’s leading climate modeling codes. By changing various conditions, such as the emission levels of greenhouse gases, scientists can model different climate change scenarios. Because running such simulations is computationally demanding, the code is only being run at a few DOE computing facilities and the National Center for Atmospheric Research. The resulting data archive, distributed over several sites, currently contains upwards of 100 terabytes of simulation data and continues to grow to petabytes by the end of the decade. Making the data available to a broad range of researchers for analysis is key to a comprehensive study of climate change.

SciDAC’s **Earth System Grid (ESG)** is a collaborative interdisciplinary project aimed at addressing

the challenge of enabling management, discovery, access and analysis of these enormous and extremely important data assets. For the modeling teams, the daily management and tracking of their data is already proving to be a significant problem. Then there are climate researchers working at institutions and universities across the U.S. whose ability to discover and use the data is extremely limited. That’s today, and the problem is rapidly escalating.

As computers become more powerful and can run the models faster and with more complex details, such as finer geographic resolution, the amount of data will increase significantly. Much of the current modeling activity is focused upon simulations aimed at the upcoming Intergovernmental Panel on Climate Change (IPCC) assessment, and these simulations have twice the horizontal resolution of the models that have been run for the past several years, which increases the resulting data volume fourfold. Figure 4 depicts new work representing yet another doubling of the resolution.

All of this adds up to an enormous increase in the volume and complexity of data that will be produced. Moving that data will become increasingly costly, which could discourage the sharing of data among the community of climate researchers.

The heart of ESG is a simple, elegant, and powerful Web portal that allows scientists to register, search, browse, and acquire the data they need. This portal was demonstrated at the Supercomputing 2003 conference on high performance computing. The portal incorporates new capabilities for cataloging datasets, registering new users and interoperability between different mass storage architectures.

One of ESG’s strategies is to dramatically reduce the amount of data that needs to be moved over the network, and the project team

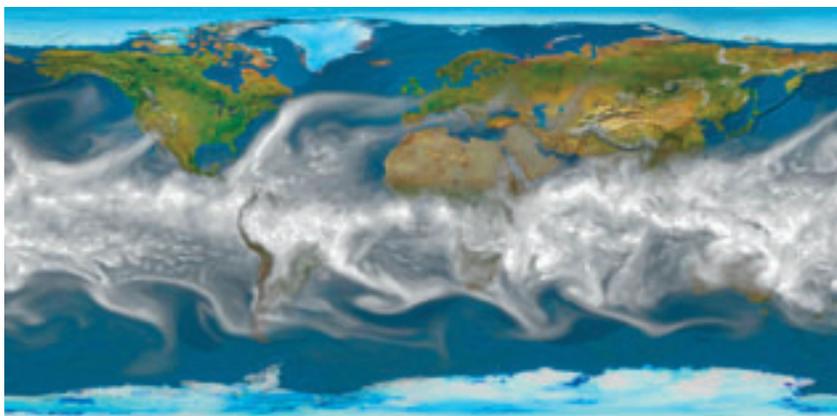


FIGURE 4: From the atmospheric component of the CCSM, this visualization shows precipitable water from a high-resolution experimental simulation (T170 resolution, about 70 km).

has done groundbreaking work in developing generalized remote data access capabilities. This involves heavy collaboration with the SciDAC DataGrid Toolkit project, as well as joint work with the community OpenDAP project.

In summer 2004, ESG began serving IPCC and other model data to a global community in close partnership with the IPCC effort and the World Meteorological Organization.

Principal Investigator: Ian Foster, Argonne National Laboratory

Security and Policy for Group Collaboration

Today, scientific advances are rarely the result of an individual toiling in isolation, but rather the result of collaboration. Such collaborations are using their combined talents to address scientific problems central to DOE's mission in areas such as particle physics experiments, global climate change and fusion science. While considerable work has been done on tools to help perform the work of a collaboration, little attention has been paid to creating mechanisms for establishing and maintaining the structure of such collaborative projects. This structure includes methods for identifying who are members of the

collaboration, what roles they play, what types of activities they are entitled to perform, what community resources are available to members of the collaboration and what are the policies set down by the owners of those resources.

The **Security and Policy for Group Collaboration** project was created to develop scalable, secure and usable methods and tools for defining and maintaining membership, rights, and roles in group collaborations. Such collaborations share common characteristics:

- The participants, as well as the resources, are distributed both geographically and organizationally.
- Collaborations can scale in size from a few individuals to thousands of participants, and membership may very dynamic.
- Collaborations may span areas of expertise, with members filling different roles within the collaboration.
- The work of the team is enabled by providing team members with access to a variety of community resources, including computers, storage systems, datasets, applications, and tools.

Central to this problem of structure is determining the identity of both participants and resources and, based on this identity, determining

the access rights of the participant and the policy of the resource. While mechanisms for authentication and authorization have been defined, the issues of distribution, dynamics and scale complicate their application to collaborative environments. Additionally, sites providing the resources for a collaboration often have overruling security mechanisms and policies in place which must be taken into account, rather than replaced by the collaboration.

The project team has instantiated its research results into the Globus Toolkit's[®] widely used Grid Security Infrastructure. Since the Globus Toolkit is already adopted by many science projects – the Particle Physics Data Grid, Earth Systems Grid, DOE Science Grid and other non-DOE Grid activities like NSF TeraGrid, NASA IPG and the European Data Grid – the project's mechanisms for security and authentication can be easily used by these scientists.

Principal Investigator: Stephen Tuecke, Argonne National Laboratory

Middleware Projects

Two projects are conducting research and development that will address individual technology elements to enable universal, ubiquitous, easy access to remote resources and to contribute to the ease with which distributed teams work together. Enabling high performance for scientific applications is especially important.

Middleware Technology to Support Science Portals: Gateways to the Grid

While extensive networks such as the Internet have been in existence for decades, it took the organization of the World Wide Web and the creation of easy-to-use Web browsers to make the Internet the dynamic information resource that

it is today. In much the same way, networked Grids linking computing, data and experimental resources for scientific research have the potential to greatly accelerate collaborative scientific discovery.

Under SciDAC, the **Middleware Technology to Support Science Portals** project was created to develop a “Science Portal” environment which would allow scientists to program, access and execute distributed Grid applications from a conventional Web browser and other tools on their desktop computers. Portals allow each user to configure their own problem-solving environment for managing their work. The goal is to allow the scientist to focus completely on the science by making the Grid a transparent extension of the user’s desktop computing environment.

The first step in this project was to create a software platform for building the portal and which could be used as a testbed. Then the project team partnered with applications researchers to get feedback on how well the applications worked when accessed by portal users. For applications, the project partnered with several National Science Foundation (NSF) projects. The team also worked with the Global Grid Forum, with NSF funding, to refine the research portal technology for production use.

Principal Investigator: Dennis Gannon, Indiana University

DataGrid Middleware: Enabling Big Science on Big Data

Many of the most challenging scientific problems being addressed by DOE researchers require access to tremendous quantities of data as well as computational resources. Physicists around the world cooperate in the analysis of petabytes of accelerator data. Climate modelers compare massive climate simulation outputs. Output from multi-million dollar online instruments, such as

the Advanced Photon Source or earthquake engineering shake tables, must be visualized in real time so that a scientist can adjust the experiment while it is running.

In order for these applications to be feasible, infrastructure must be in place to support efficient, high performance, reliable, secure, and policy aware management of large-scale data movement. The SciDAC **DataGrid Middleware** project has provided tools in three primary areas in support of this goal. GridFTP, developed primarily at Argonne National Laboratory, provides a secure, robust, high performance transport mechanism that is recognized as the de facto standard for transport on the Grid. The Globus replica tools, developed primarily at the University of Southern California’s Information Sciences Institute, provide tracking of replicated data sets and provide efficiency in the selection of data sets to access. The Condor team at the University of Wisconsin is providing storage resource management, particularly space reservation tools. Together, these three components provide the basis for DataGrid applications.

These DataGrid tools are being used by other SciDAC projects. The Earth System Grid and the Particle Physics Data Grid use GridFTP servers to stage input data and move results to mass storage systems. They also employ the project’s first generation replica catalog to determine the best location from which to store and/or retrieve data. The Laser Interferometer Gravitational Wave Observatory project has moved over 50 terabytes of data and has a replica location service with over 3 million logical files and over 30 million physical filenames. The Grid3 project, part of the Grid Physics Network, moves over 4 terabytes of data a day. GridFTP is now a Global Grid Forum standard and there is a replication working group working on standards as well. The tools developed as part of this proj-

ect have become the de facto standard for data management and, in fact, are opening up new realms of scientific discovery.

For example, participants in the Large Hadron Collider experiment, which will come on line in 2007 in Switzerland, are conducting “data challenges” to ensure that the data coming out of the experiment can be accommodated. In a 2005 challenge, a continuous average data flow of 600 megabytes per second was sustained over 10 days, for a total transfer of 500 terabytes of data – all enabled by GridFTP. Moving the same amount of data over a typical household network connection (at 512 kilobits per second) would take about 250 years.

Principal Investigators: Ian Foster, Argonne National Laboratory; Carl Kesselman, University of Southern California Information Sciences Institute; Miron Livny, University of Wisconsin-Madison

Networking Projects

As high-speed computer networks play an increasingly important role in large-scale scientific research, scientists are becoming more dependent on reliable and robust network connections, whether for transferring experimental data, running applications on distributed supercomputers or collaborating across the country or around the globe. But as anyone who has logged onto the Internet can attest, network connections can unexpectedly break or bog down at any time. To help keep such system problems from disrupting scientific research, SciDAC funded three research projects to develop better methods for analyzing and addressing network performance.

INCITE: Monitoring Performance from the Outside Looking In

The **INCITE (InterNet Control and Inference Tools at the Edge)** project was aimed at developing tools and services to improve the performance of applications running on distributed computing systems linked by computational grids. Although such applications are complex and difficult to analyze, knowledge of the internal network traffic conditions and services could optimize the overall performance. Without special-purpose network support (at every router), the only alternative is to indirectly infer dynamic network characteristics from edge-based network measurements. Therefore, this project researched and developed network tools and services for analyzing, modeling, and characterizing high-performance network end-to-end performance based solely on edge-based measurements made at the hosts and routers at the edges of the network, rather than midpoints on the network. Specifically, the goal was to develop multi-scalable on-line tools to characterize and map network performance as a function of space, time, applications, protocols, and services for end-to-end measurement, prediction, and network diagnosis.

Two of the tools developed by the project have been incorporated into a toolkit used to identify performance problems at a number of major networks and research sites. INCITE users include: Globus, SciDAC's Supernova Science Center and Particle Physics Data Grid Pilot, Scientific Workspaces of the Future, TeraGrid, Transpac at Indiana University, San Diego Supercomputing Center, Telecordia, CAIDA, Autopilot, TAU and the European GridLab project.

Principal Investigator: Richard Baraniuk, Rice University

Logistical Networking: A New Approach for Data Delivery

The project for **Optimizing Performance and Enhancing Functionality of Distributed Applications using Logistical Networking** explored the use of a new technique to provide fast, efficient and reliable data delivery to support the high-performance applications used by SciDAC collaborators. Logistical networking (LN) is a

Mbps between Oak Ridge National Laboratory (ORNL) in Tennessee and North Carolina State University (NCSU). A new, private LN infrastructure is in place, designed specifically for TSI and other SciDAC projects. Depots at ORNL, San Diego Supercomputing Center, SUNY-Stony Brook and NCSU provide 8 terabytes of storage and form the network's backbone.

The second area of focus is support for SciDAC's fusion energy

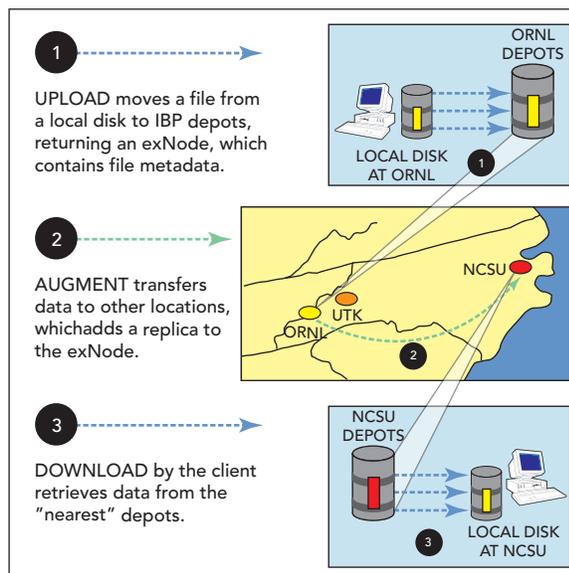


FIGURE 5: The interworking components of the Logistical Runtime System Tools (LoRS Tools) in the high performance distribution of data between IBP depots at Terascale Supernova Initiative sites at ORNL and NCSU.

new way of synthesizing networking and storage to create a communication infrastructure that provides superior control of data movement and management for distributed applications based on shared network storage. LN software tools allow users to create local storage "depots" or utilize shared storage depots deployed worldwide to easily accomplish long-haul data transfers, temporary storage of large datasets (on the order of terabytes), pre-positioning of data for fast on-demand delivery, and high performance content distribution such as streaming video.

A primary research drive of the project was to work with SciDAC's Terascale Supernova Initiative (TSI) group, which uses LN to share massive datasets at speeds up to 220

researchers. Typical fusion plasma experiments require real-time feedback for rapid tuning of experimental parameters, meaning data must be analyzed during the 15-minute intervals between plasma-generating pulses in experimental reactors. Such rapid assimilation of data is achieved by a geographically dispersed research team, a challenge for which LN technologies are well suited.

Principal Investigator: Micah Beck, University of Tennessee

Estimating Bandwidth to Find the Fastest, Most Reliable Path

The **Bandwidth Estimation: Measurement Methodologies and Applications** project focused on the

research, development and deployment of scalable and accurate bandwidth estimation tools for high-capacity network links. Such tools would allow applications to adapt to changing network conditions, finding the fastest, most reliable network path. This adaptive ability would benefit a large class of data-intensive and distributed scientific applications. However, previous tools and methodologies for measuring network capacity, available bandwidth and throughput have been largely ineffective across real Internet infrastructures.

The goal was to develop meas-

urement methodologies and tools for various bandwidth-related metrics, such as per-hop capacity, end-to-end capacity, and end-to-end available bandwidth. Capacity is the maximum throughput that the path can provide to an application when there is no competing traffic. Available bandwidth, on the other hand, is the maximum throughput that the path can provide to an application given the path's current traffic load. Measuring capacity is crucial for debugging, calibrating and managing a path, while measuring available bandwidth is important for predicting end-to-end per-

formance of applications, for dynamic path selection and traffic engineering.

As part of the project, Georgia Tech researchers developed two bandwidth estimation tools, Pathrate (which measures capacity) and Pathload (which measures availability), which were released in January 2004. Pathrate and Pathload have been downloaded by more than 2,000 users around the world.

Principal Investigators: K. C. Klaffy, San Diego Supercomputer Center; Constantinos Davrolis, Georgia Institute of Technology

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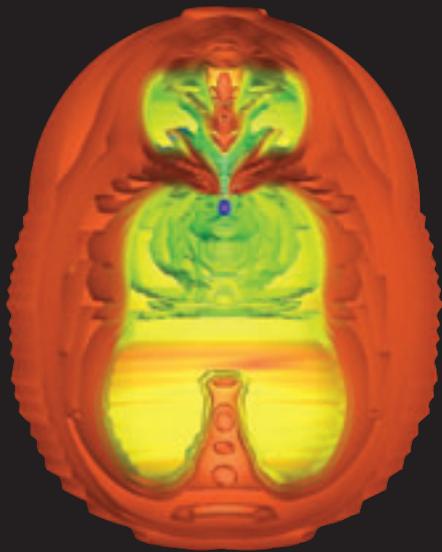
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This report was researched, written and edited by Jon Bashor and John Hules of the LBNL Computing Sciences Communications Group, and designed by the Creative Services Office at LBNL. JO11877
LBNL-61963

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